1	Ice, Cloud, and Land Elevation Satellite-2 (ICESat-2)
2	Almosition Theory (incl. Books, B. company (ATBB)
3 4	Algorithm Theoretical Basis Document (ATBD)
5	for
6	
7	Land - Vegetation Along-Track Products (ATL08)
8	
9	
10	
11	Contributions by Land/Vegetation SDT Team Members
12	and ICESat-2 Project Science Office
13	(Amy Neuenschwander, Katherine Pitts, Benjamin Jelley, John Robbins
14	Jonathan Markel, Sorin Popescu, Ross Nelson, David Harding, Dylan
15	Pederson, Brad Klotz, and Ryan Sheridan)
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17	
18	ATBD prepared by
19	Amy Neuenschwander
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21	
22	15 May 2021
23	(This ATBD Version corresponds to release 005 of the ICESat-2 ATL08
24	data)
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27	Content reviewed: technical approach, assumptions, scientific soundness,
28	maturity, scientific utility of the data product
29	

ATBD Version	Change
2016 Nov	Product segment size changed from 250 signal photons to
2010 NOV	100 m using five 20m segments from ATL03 (Sec 2)
2016 Nov	Filtered signal classification flag removed from
2010 NOV	
2016 Nov	classed_pc_flag (Sec 2.3.2)
2016 Nov	DRAGANN signal flag added (Sec 2.3.5)
2016 Nov	Do not report segment statistics if too few ground photons
2016 N-	within segment (Sec 4.14 (3))
2016 Nov	Product parameters added: h_canopy_uncertainty,
	landsat_flag, d_flag, delta_time_beg, delta_time_end,
2017 Mass	night_flag, msw_flag (Sec 2)
2017 May	Revised region boundaries to be separated by continent (Sec
2017 M	2)
2017 May	Alternative DRAGANN parameter calculation added (Sec
2017 M -	4.3.1)
2017 May	Set canopy flag = 0 when $L$ - $km$ segment is over Antarctica or
2017 Mass	Greenland regions (Sec 4.4 (1))
2017 May	Change initial canopy filter search radius from 3 m to 15 m
2017 M -	(Sec 4.8 (6))
2017 May	Product parameters removed: h_rel_ph, terrain_thresh
2017 May	Product parameters added: segment_id, segment_id_beg,
2017 Index	segment_id_end, dem_flag, surf_type (Sec 2)
2017 July	Urban flag added (Sec 2.4.20)
2017 July	Dynamic point spread function added (Sec 4.10 (6))
2017 July	Methodology for processing <i>L-km</i> segments with buffer
2017 Index	added (Sec 4.1 (2), Sec 4.16)
2017 July	Revised alternative DRAGANN methodology (see bolded text in Sec 4.3.1)
2017 July	,
2017 July 2017 July	Added post-DRAGANN filtering methodology (Sec 4.6) Updated SNR to be estimated from superset of ATL03 and
2017 July	
	DRAGANN found signal used for processing ATL08 (Sec 2.5.18)
2017 September	More details added to DRAGANN description (Sec 4.3), and
2017 September	corrections to DRAGANN implementation (Sec 3.1.1, Sec 4.3
	<u> </u>
2017 September	(9)) Added Appendix A – very detailed DRAGANN description
2017 September	Revised alternative DRAGANN methodology (see bolded text
2017 September	in Sec 4.3.1)
2017 September	Clarified SNR calculation (Sec 2.5.18, Sec 4.3 (18))
2017 September	Added cloud flag filtering option (Sec 1)
2017 September	Added top of canopy median surface filter (Sec 3.5 (a), Sec
2017 deptember	4.9 (3), Sec 4.11 (1-3))
	117 (0), 000 1111 (1 0))

2017 September	Modified 500 canopy photon segment filter (Sec 3.5 (c), Sec 4.11 (6))
2017 November	Added solar_azimuth, solar_elevation, and n_seg_ph to Reference Data group; parameters were already in product (Sec 2.4)
2017 November	Specified number of ground photons threshold for relative canopy product calculations (Sec 4.15 (2)); no number of ground photons threshold for absolute canopy heights (Sec 4.15.1 (1))
2017 November	Changed the ATL03 signal used in superset from all ATL03 signal (signal_conf_ph flags 1-4) to the medium-high confidence flags (signal_conf_ph flags 3-4) (Sec 3.1, Sec 4.3 (17))
2017 November	Removed Date parameter from Table 2.4 since UTC date is in file metadata
2018 March	Clarified that cloud flag filtering option should be turned off by default (Sec 1)
2018 March	Changed h_diff_ref QA threshold from 10 m to 25 m (Table 5.2)
2018 March	Added absolute canopy height quartiles, canopy_h_quartile_abs ( <i>Later removed</i> )
2018 March	Removed psf_flag from main product; psf_flag will only be a QAQC alert (Sec 5.2)
2018 March	Added an Asmooth filter based on the reference DEM value (Sec 4.5 (4-5))
2018 March	Changed relief calculation to 95 <sup>th</sup> – 5 <sup>th</sup> signal photon heights. (Sec 4.5 (6))
2018 March	Adjusted the Asmooth smoothing methodology (Sec 4.5 (8))
2018 March	Recalculate the Asmooth surface after filtering outlying noise from signal, then detrend signal height data (Sec 4.6 (3-4))
2018 March	Added option to run alternative DRAGANN process again in high noise cases (Sec 4.3.3)
2018 March	Changed global land cover reference to MODIS Global Mosaics product (Sec 2.4.16)
2018 March	Adjusted the top of canopy median filter thresholds based on SNR (Sec 4.11 (1-2))
2018 March	Added a final photon classification QA check (Sec 4.13, Table 5.2)
2018 March	Added slope adjusted terrain parameters (Later removed)
2018 June	Replaced slope adjusted terrain parameters with terrain best fit parameter (Sec 2.1.14, 4.14 (2.e))
2018 June	Clarified source for water mask (Sec 2.4.18)
2018 June	Clarified source for urban mask (Sec 2.4.20)
2018 June	Added expansion to the terrain_slope calculation (Sec 4.14)
2018 June	Removed canopy_d_quartile

2018 June	Removed canopy_quartile_heights and			
	canopy_quartile_heights_abs, replaced with canopy_h_metrics (Secs 2.2.3, 4.15 (6), 4.15.1 (5))			
2018 *** draft 1	Delta_time specified as mid-segment time, rather than mean			
2010 draft 1	segment time (Sec 2.4.7)			
2018 *** draft 1	QA/QC products to be reported on a per orbit basis, rather			
	than per region (Sec 5.2)			
2018 *** draft 1	Added more detail to landsat_flag description (Sec. 2.2.23)			
2018 *** draft 1	Added psf_flag back into ATL08 product, as it is also needed for the QA product (Sec 2.5.12)			
2018 *** draft 1	Specified that the sigma_h value reported here is the mean of			
	the ATL03 reported sigma_h values (Sec 2.5.7)			
2018 *** draft 1	Removed n_photons from all subgroups			
2018 *** draft 1	Better defined the interpolation and smoothing methods			
	used throughout:			
	• 1 (3): Interpolation – nearest			
	• 4.5 (5): Interpolation – PCHIP			
	• 4.5 (8): Smoothing – moving average			
	• 4.6 (3): Interpolation – PCHIP			
	• 4.6 (3): Smoothing – moving average			
	• 4.7 (10): Smoothing – moving average			
	• 4.7 (11): Interpolation – linear			
	• 4.7 (12): Smoothing – moving average			
	• 4.7 (13): Interpolation – linear			
	• 4.7 (14): Smoothing – moving average			
	• 4.7 (15): Smoothing – Savitzky-Golay			
	• 4.7 (16): Interpolation – linear			
	• 4.7 (21): Interpolation – PCHIP			
	• 4.9 (10): Interpolation – linear			
	• 4.10 (all): Smoothing – moving average			
	• 4.10 (6.b): Interpolation – linear			
	• 4.11 (1.a): Interpolation – linear			
	• 4.11 (1.c): Smoothing – lowess			
	• 4.11 (4): Interpolation – PCHIP			
	• 4.11 (7): Interpolation – PCHIP			
	• 4.11 (9): Smoothing – moving average			
2018 *** draft 1	4.14 (2.e.i.1): Interpolation – linear  Added ref clay and ref agimuth back in (it was mistakenly)			
2018 · · · drait I	Added ref_elev and ref_azimuth back in (it was mistakenly removed in a previous version; Secs 2.5.3, 2.5.4)			
2018 *** draft 1	Clarified wording of h_canopy_quad definition (Sec 2.2.18)			
2018 *** draft 1	Updated segment_snowcover description to match the			
2010 urait 1	ATL09 snow_ice parameter it references (Sec 2.4.19) and			
	added product reference to Table 4.2			
	duded product reference to rubic 1.2			

2018 *** draft 1	Added ph_ndx_beg (Sec 2.5.22); parameter was already on product
2018 *** draft 1	Added dem_removal_flag for QA purposes (Sec 2.4.13; Table 5.2)
2018 *** draft 2	Reformatted QA/QC trending and trigger alert list into a table for better clarification (Table 5.3)
2018 *** draft 2	Replaced n_photons in Table 5.2 with n_te_photons, n_ca_photons, and n_toc_photons
2018 *** draft 2	Removed beam_number from Table 2.5. Beam number and weak/strong designation within gtx group attributes.
2018 *** draft 2	Clarified calculation of h_te_best_fit (Sec 4.14 (2.e))
2018 *** draft 2	Changed h_canopy and h_canopy_abs to be 98 <sup>th</sup> percentile height (Table 2.2, Sec 2.2.5, Sec 2.2.6, Sec 4.15 (4), Sec 4.15.1 (3))
2018 *** draft 2	Separated h_canopy_metrics_abs from h_canopy_metrics (Table 2.2, Sec 2.2.3, Sec 4.15.1 (5))
2018 October	Removed 99 <sup>th</sup> percentile from h_canopy_metrics and h_canopy_metrics_abs (Table 2.2, Sec 2.2.3, Sec 2.2.4, Sec 4.15 (4), Sec 4.15.1 (5))
2018 December	Renamed and reworded Section 4.3.1 to better indicate that the DRAGANN preprocessing step is not optional
2018 December	Specified that DRAGANN should use along-track time, and added time rescaling step (Sec 4.3 (1 - 4))
2018 December	Added DRAGANN changes made to better capture sparse canopy in cases of low noise rates (Sec 4.3, Appendix A)
2018 December	Made corrections to DRAGANN description regarding the determination of the noise Gaussian (Sec 3.1.1, Sec 4.3)
2018 December	Removed h_median_canopy and h_median_canopy_abs, as they are equivalent to canopy_h_metrics(50) and canopy_h_metrics_abs(50) (Table 2.2, Sec 4.15 (5), Sec 4.15.1 (4))
2018 December	Removed the requirement that > 5% ground photons required to calculate relative canopy height parameters (Table 2.2, Sec 4.15 (2))
2018 December	Added canopy relative height confidence flag (canopy_rh_conf) based on the percentage of ground and canopy photons in a segment (Table 2.2, Sec 4.15 (2))
2018 December	Added ATL09 layer_flag to ATL08 output (Table 2.5, Table 4.2)
2019 February	Adjusted cloud filtering to be based on ATL09 backscatter analysis rather than cloud flags (Sec 4.1)
2019 March 5	Updated ATL09-based product descriptions reported on ATL08 product (Secs 2.5.13, 2.5.14, 2.5.15, 2.5.16)
2019 March 5	Updated cloud-based low signal filter methodology, and moved to first step of ATL08 processing (Sec 4.1)

(Table 2.2, Sec 2.2.24)  2019 March 13  Change ATL08 product output regions to match ATL03 regions (Sec 2), but keep ATL08 regions internally and report in new parameter atl08_regions (Table 2.4, Sec 2.4.22)  2019 March 13  Add methodology for handling short ATL08 processing segments at the end of an ATL03 granule (Sec 4.2), and output distance the processing segment length is extended into new parameter last_seg_extend (Table 2.4, Sec 2.4.23)  2019 March 13  Add preprocessing step for removing atmospheric and ocean tide corrections from ATL03 heights (Later removed)  Remove preprocessing step for removing atmospheric and ocean tide corrections from ATL03 heights, since those values are now removed from the ATL03 photon heights.  2019 March 27  Replaced ATL03 region figure with corrected version (Figure 2.2)  2019 March 27  Specified that at least 50 classed photons are required to create the 100 m land and canopy products (Secs 2, 4.14(1), 4.15(1))  2019 March 27  Clarified that any non-extended segments would report a land_seg_extend value of 0 (Sec 4.2, Sec 2.4.23)  2019 April 30  Fixed the error in Eqn 1.4 for the sigma topo value  2019 May 13  Specified for cloud flag carry-over from ATL09 that ATL08 will report the highest cloud flag if an 08 segment straddles two 09 segments. (Section 2.5)  2019 May 13  Changed parameter cloud_flag_asr to cloud_flag_atm since the cloud_flag_asr is likely not to work over land due to varying surface reflectance (Sec, 2.5)  2019 May 13  Clarification on the calculation of gradient for slope that feeds into the calculation of the point spread function (Sec 4.11)  Changed Landsat canopy cover percentage to 3 % (from original value of 5%) (Section 4.4)  Added a QA method for DRAGANN flags to help remove false positives (now Section 4.3.1)		
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positives (now Section 4.3.1)	2019 July 8	
2019 July 8 Set the window size to 9 rather than Smooth Size for the final	2019 July 8	
ground finding step. (Section 4.11 and 4.12)	2019 July 8	Set the window size to 9 rather than SmoothSize for the final ground finding step. (Section 4.11 and 4.12)
2019 July 8 Added a brightness flag to land segments. (Section 2.4.21)	2019 July 8	Added a brightness flag to land segments. (Section 2.4.21)
2019 November Added subset_te_flag to (Section 2.1) which indicate 100 m	2019 November	Added subset_te_flag to (Section 2.1) which indicate 100 m
segments that are populated by less than 100 m worth of data	12	segments that are populated by less than 100 m worth of

2019 November 12	Added subset_can_flag (section 2.2) which indicate 100 m segments that are populated by less than 100 m worth of data
2020 January 5	Clarified the interpolation of values (latitude, longitude, delta time) when the 100 m segments are populated by less than 100 m worth of data. (Section 2.4.3 and 2.4.4)
2020 January 13	Fine-tuned the methodology to improve ground finding by first histogramming the photons to improve detecting the ground in cases of dense canopy. (Section 4.8)
2020 January 13 2020 February 14	Updated ATL08 HDF5 file organization figure in Section 2.1 Added sentence to avoid ATL03 data having a degraded PPD flag to beginning of Section 4
2020 February 14	Added documentation for removing signal photons due to cloud contamination by checking the reference DEM to beginning of Section 4
2020 February 14	Added full saturation flag and near saturation flag from ATL03 to ATL08 data product to Section 2.
2020 February 14	Added statement to clarify handling of remaining geosegments that do not fit within a 100 m window at the end of a 10-km processing window in Section 4.2
2020 April 15	Added ph_h parameter to photon group on data structure. ph_h is the photon height above the interpolated ground surface.
2020 May 15	Added sat_flag which is derived from the ATL03 product. The saturation flag indicates that the ATL08 segment experienced some saturation which is often an indicator for water
2020 May 15	Canopy height metrics (relative and absolute heights) were expanded to every 5% ranging from 5 – 95%.
2020 May 15	The Landsat canopy cover check to determine whether the algorithm should search for both ground and canopy or just ground has been disabled. Now the ATL08 algorithm will search for both ground and canopy points everywhere.
2020 June 15	Corrected the calculation of the absolute canopy heights
2020 June 15	Changed the search radius for initial top of canopy determination (Section 4.9)
2020 September 1	Incorporate the quality_ph flag from ATL03 into the ATL08 workflow (beginning of Section 4)
2020 September 1	Added the calculation of Terrain photon rate (photon_rate_te) for each ATL08 segment to the land product (Section 2.1.16)
2020 September 1	Added the calculation of canopy photon rate (photon_rate_can) for each ATL08 segment to the land product (Section 2.2.26)

2020 September 1	Changed the k-d tree search radius for the top of canopy from 15 m to 100 m. Section 4.9.6
2020 September 15	Added new parameter for terrain heights (h_te_rh25) which represents the height of the 25% of ground cumulative distribution.
2021 March 15	Added terrain_best_fit_geosegment (h_te_best_fit_20m) parameter to the data product. 20 m estimate of best fit terrain height
2021 March 15	Added canopy_height_geosegment (h_canopy_20m) to the data product. 20 m estimate of relative canopy height
2021 March 15	Added latitude_20m to the data product.
2021 March 15	Added longitude_20m to the data product
2021 March 15	Updated the urban_flag parameter. Inclusion of the DLR Global Urban Footprint (GUF) as a potential indicator of man-made/built structures. Section 2.4.20
2021 March 15	Updated the Segment_landcover with Copernicus. Replace the MODIS landcover value with the landcover classification from the 100 m Copernicus landcover. Section 2.4.16
2021 March 15	Added the Segment_Woody_Vegetation_Fractional_cover. Inclusion of a woody vegetation fraction cover derived from the 2019 Copernicus fractional cover data products. Section 2.4.17
2021 March 15	Removed Landsat_perc (Landsat Percentage Calculation), Landsat_flag, and Canopy_flag from the ATL08 data product and from the algorithm. Removed all reference to Landsat from the ATBD.

# **Contents**

36	List of Table	es	16
37	List of Figur	res	17
38	1 INTRO	DUCTION	19
39	1.1. Ba	ckground	20
40	1.2 Ph	oton Counting Lidar	22
41	1.3 Th	e ICESat-2 concept	23
42	1.4 He	ight Retrieval from ATLAS	26
43	1.5 Acc	curacy Expected from ATLAS	28
44	1.6 Ad	ditional Potential Height Errors from ATLAS	30
45	1.7 De	nse Canopy Cases	30
46	1.8 Spa	arse Canopy Cases	31
47	2. ATL08	: DATA PRODUCT	32
48	2.1 Sul	bgroup: Land Parameters	35
49	2.1.1	Georeferenced_segment_number_beg	36
50	2.1.2	Georeferenced_segment_number_end	37
51	2.1.3	Segment_terrain_height_mean	37
52	2.1.4	Segment_terrain_height_med	37
53	2.1.5	Segment_terrain_height_min	38
54	2.1.6	Segment_terrain_height_max	38
55	2.1.7	Segment_terrain_height_mode	38
56	2.1.8	Segment_terrain_height_skew	38
57	2.1.9	Segment_number_terrain_photons	39
58	2.1.10	Segment height_interp	39
59	2.1.11	Segment h_te_std	39
60	2.1.12	Segment_terrain_height_uncertainty	39
61	2.1.13	Segment_terrain_slope	39

62	2.1.14	Segment_terrain_height_best_fit	40
63	2.1.15	Segment_terrain_height_25	40
64	2.1.16	Subset_te_flag {1:5}	40
65	2.1.17	Segment Terrain Photon Rate	41
66	2.1.18	Terrain Best Fit GeoSegment {1:5}	41
67	2.2 Sub	ogroup: Vegetation Parameters	41
68	2.2.1	Georeferenced_segment_number_beg	44
69	2.2.2	Georeferenced_segment_number_end	44
70	2.2.3	Canopy_height_metrics_abs	44
71	2.2.4	Canopy_height_metrics	45
72	2.2.5	Absolute_segment_canopy_height	45
73	2.2.6	Segment_canopy_height	46
74	2.2.7	canopy_height GeoSegment {1:5}	46
75	2.2.8	Absolute_segment_mean_canopy	46
76	2.2.9	Segment_mean_canopy	47
77	2.2.10	Segment_dif_canopy	47
78	2.2.11	Absolute_segment_min_canopy	47
79	2.2.12	Segment_min_canopy	47
80	2.2.13	Absolute_segment_max_canopy	47
81	2.2.14	Segment_max_canopy	48
82	2.2.15	Segment_canopy_height_uncertainty	48
83	2.2.16	Segment_canopy_openness	49
84	2.2.17	Segment_top_of_canopy_roughness	49
85	2.2.18	Segment_canopy_quadratic_height	49
86	2.2.19	Segment_number_canopy_photons	49
87	2.2.20	Segment_number_top_canopy_photons	50
88	2 2 21	Centroid height	50

89	2.2.22	Segment_rel_canopy_conf	50
90	2.2.23	Subset_can_flag {1:5}	50
91	2.2.24	Segment Canopy Photon Rate	51
92	2.3 Sub	ogroup: Photons	51
93	2.3.1	Indices_of_classed_photons	52
94	2.3.2	Photon_class	52
95	2.3.3	Georeferenced_segment_number	52
96	2.3.4	Photon Height	53
97	2.3.5	DRAGANN_flag	53
98	2.4 Sub	ogroup: Reference data	53
99	2.4.1	Georeferenced_segment_number_beg	55
100	2.4.2	Georeferenced_segment_number_end	55
101	2.4.3	Segment_latitude	55
102	2.4.4	Geosegment_latitude{1:5}	56
103	2.4.5	Segment_longitude	56
104	2.4.6	Geosegment_longitude{1:5}	57
105	2.4.7	Delta_time	57
106	2.4.8	Delta_time_beg	57
107	2.4.9	Delta_time_end	57
108	2.4.10	Night_Flag	57
109	2.4.11	Segment_reference_DTM	57
110	2.4.12	Segment_reference_DEM_source	58
111	2.4.13	Segment_reference_DEM_removal_flag	58
112	2.4.14	Segment_terrain_difference	58
113	2.4.15	Segment_terrain flag	58
114	2.4.16	Segment_landcover	58
115	2.4.17	Segment Woody Vegetation Fractional Cover	60

116	2.4.18	Segment_watermask	60
117	2.4.19	Segment_snowcover	60
118	2.4.20	Urban_flag	60
119		Surface_Type	60
120	2.4.21	ATL08_region	61
121	2.4.22	Last_segment_extend	61
122	2.4.23	Brightness_flag	62
123	2.5 Sub	ogroup: Beam data	62
124	2.5.1	Georeferenced_segment_number_beg	65
125	2.5.2	Georeferenced_segment_number_end	65
126	2.5.3	Beam_coelevation	65
127	2.5.4	Beam_azimuth	65
128	2.5.5	ATLAS_Pointing_Angle	66
129	2.5.6	Reference_ground_track	66
130	2.5.7	Sigma_h	66
131	2.5.8	Sigma_along	66
132	2.5.9	Sigma_across	66
133	2.5.10	Sigma_topo	67
134	2.5.11	Sigma_ATLAS_LAND	67
135	2.5.12	PSF_flag	67
136	2.5.13	Layer_flag	67
137	2.5.14	Cloud_flag_atm	68
138	2.5.15	MSW	68
139	2.5.16	Cloud Fold Flag	68
140	2.5.17	Computed_Apparent_Surface_Reflectance	68
141	2.5.18	Signal_to_Noise_Ratio	69
142	2.5.19	Solar Azimuth	69

143	2.5.20	Solar_Elevation	69
144	2.5.21	Number_of_segment_photons	69
145	2.5.22	Photon_Index_Begin	69
146	2.5.23	Saturation Flag	69
147	3 ALGOF	RITHM METHODOLOGY	71
148	3.1 No	ise Filtering	71
149	3.1.1	DRAGANN	72
150	3.2 Su	rface Finding	76
151	3.2.1	De-trending the Signal Photons	78
152	3.2.2	Canopy Determination	78
153	3.2.3	Variable Window Determination	79
154	3.2.4	Compute descriptive statistics	80
155	3.2.5	Ground Finding Filter (Iterative median filtering)	82
156	3.3 To	p of Canopy Finding Filter	83
157	3.4 Cla	ssifying the Photons	84
158	3.5 Re	fining the Photon Labels	84
159	3.6 Ca	nopy Height Determination	89
160	3.7 Lir	nk Scale for Data products	89
161	4. ALGOF	RITHM IMPLEMENTATION	90
162	4.1 Clo	oud based filtering	93
163	4.2 Pro	eparing ATL03 data for input to ATL08 algorithm	95
164	4.3 No	ise filtering via DRAGANN	97
165	4.3.1	DRAGANN Quality Assurance	99
166	4.3.2	Preprocessing to dynamically determine a DRAGANN parameter	101
167	4.3.3	Iterative DRAGANN processing	104
168	4.4 Co	mpute Filtering Window	104
169	4.5 De	-trend Data	105

170	4.6	Filter outlier noise from signal	106
171	4.7	Finding the initial ground estimate	106
172	4.8	Find the top of the canopy (if canopy_flag = 1)	109
173	4.9	Compute statistics on de-trended (Asmooth) data	110
174	4.10	Refine Ground Estimates	111
175	4.11	Canopy Photon Filtering	113
176	4.12	Compute individual Canopy Heights	115
177	4.13	Final photon classification QA check	116
178	4.14	Compute segment parameters for the Land Products	116
179	4.15	Compute segment parameters for the Canopy Products	119
180	4.1	15.1 Canopy Products calculated with absolute heights	120
181	4.16	Record final product without buffer	121
182	5 DA	ATA PRODUCT VALIDATION STRATEGY	122
183	5.1	Validation Data	122
184	5.2	Internal QC Monitoring	125
185	6 RE	EFERENCES	131
186			

187	List of Tables	
188	Table 2.1. Summary table of land parameters on ATL08	35
189	Table 2.2. Summary table of canopy parameters on ATL084	12
190	Table 2.3. Summary table for photon parameters for the ATL08 product	51
191	Table 2.4. Summary table for reference parameters for the ATL08 product5	53
192	Table 2.5. Summary table for beam parameters for the ATL08 product	52
193 194	Table 3.1. Standard deviation ranges utilized to qualify the spread of photons within moving window	
195	Table 4.1. Input parameters to ATL08 classification algorithm	<del>)</del> 1
196	Table 4.2. Additional external parameters referenced in ATL08 product	<del>)</del> 2
197 198	Table 5.1. Airborne lidar data vertical height (Z accuracy) requirements for validation data	22
199	Table 5.2. ATL08 parameter monitoring.	25
200	Table 5.3. QA/QC trending and triggers12	29
201		

202	List of Figures
203	Figure 1.1. Various modalities of lidar detection. Adapted from Harding, 200923
204 205 206	Figure 1.2. Schematic of 6-beam configuration for ICESat-2 mission. The laser energy will be split into 3 laser beam pairs – each pair having a weak spot (1X) and a strong spot (4X)25
207 208 209	Figure 1.3. Illustration of off-nadir pointing scenarios. Over land (green regions) in the mid-latitudes, ICESat-2 will be pointed away from the repeat ground tracks to increase the density of measurements over terrestrial surfaces26
210 211	Figure 1.4. Illustration of the point spread function, also referred to as Znoise, for a series of photons about a surface
212	Figure 2.1. HDF5 data structure for ATL08 products33
213	Figure 2.2. ATL03 granule regions; graphic from ATL03 ATBD (Neumann et al.)34
214	Figure 2.3. ATL08 product regions35
215 216 217	Figure 2.4. Illustration of canopy photons (red dots) interaction in a vegetated area. Relative canopy heights, H <sub>i</sub> , are computed by differencing the canopy photon height from an interpolated terrain surface
218 219	Figure 3.1. Combination of noise filtering algorithms to create a superset of input data for surface finding algorithms
220 221	Figure 3.2. Histogram of the number of photons within a search radius. This histogram is used to determine the threshold for the DRAGANN approach74
222	Figure 3.3. Output from DRAGANN filtering. Signal photons are shown as blue76
223	Figure 3.4. Flowchart of overall surface finding method
224 225	Figure 3.5. Plot of Signal Photons (black) from 2014 MABEL flight over Alaska and de-trended photons (red)78
226	Figure 3.6. Shape Parameter for variable window size80
227 228 229	Figure 3.7. Illustration of the standard deviations calculated for each moving window to identify the amount of spread of signal photons within a given window.
230 231	Figure 3.8. Three iterations of the ground finding concept for <i>L-km</i> segments with canopy83

<ul><li>232</li><li>233</li></ul>	Figure 3.9. Example of the intermediate ground and top of canopy surfaces calculated from MABEL flight data over Alaska during July 201486
234 235 236 237	Figure 3.10. Example of classified photons from MABEL data collected in Alaska 2014. Red photons are photons classified as terrain. Green photons are classified as top of canopy. Canopy photons (shown as blue) are considered as photons lying between the terrain surface and top of canopy
238 239 240 241	Figure 3.11. Example of classified photons from MABEL data collected in Alaska 2014. Red photons are photons classified as terrain. Green photons are classified as top of canopy. Canopy photons (shown as blue) are considered as photons lying between the terrain surface and top of canopy
242 243 244 245	Figure 3.12. Example of classified photons from MABEL data collected in Alaska 2014. Red photons are photons classified as terrain. Green photons are classified as top of canopy. Canopy photons (shown as blue) are considered as photons lying between the terrain surface and top of canopy
246 247	Figure 5.1. Example of <i>L-km</i> segment classifications and interpolated ground surface
248	

#### 1 INTRODUCTION

This document describes the theoretical basis and implementation of the processing algorithms and data parameters for Level 3 land and vegetation heights for the non-polar regions of the Earth. The ATL08 product contains heights for both terrain and canopy in the along-track direction as well as other descriptive parameters derived from the measurements. At the most basic level, a derived surface height from the ATLAS instrument at a given time is provided relative to the WGS-84 ellipsoid. Height estimates from ATL08 can be compared with other geodetic data and used as input to higher-level ICESat-2 products, namely ATL13 and ATL18. ATL13 will provide estimates of inland water-related heights and associated descriptive parameters. ATL18 will consist of gridded maps for terrain and canopy features.

The ATL08 product will provide estimates of terrain heights, canopy heights, and canopy cover at fine spatial scales in the along-track direction. Along-track is defined as the direction of travel of the ICESat-2 satellite in the velocity vector. Parameters for the terrain and canopy will be provided at a fixed step-size of 100 m along the ground track referred to as a segment. A fixed segment size of 100 m was chosen to provide continuity of data parameters on the ATL08 data product. From an analysis perspective, it is difficult and cumbersome to attempt to relate canopy cover over variable lengths. Furthermore, a segment size of 100 m will facilitate a simpler combination of along-track data to create the gridded products.

We anticipate that the signal returned from the weak beam will be sufficiently weak and may prohibit the determination of both a terrain and canopy segment height, particularly over areas of dense vegetation. However, in more arid regions we anticipate producing a terrain height for both the weak and strong beams.

In this document, section 1 provides a background of lidar in the ecosystem community as well as describing photon counting systems and how they differ from discrete return lidar systems. Section 2 provides an overview of the Land and Vegetation parameters and how they are defined on the data product. Section 3 describes the basic methodology that will be used to derive the parameters for ATL08.

Section 4 describes the processing steps, input data, and procedure to derive the data parameters. Section 5 will describe the test data and specific tests that NASA's implementation of the algorithm should pass in order to determine a successful implementation of the algorithm.

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#### 1.1. Background

The Earth's land surface is a complex mosaic of geomorphic units and land cover types resulting in large variations in terrain height, slope, roughness, vegetation height and reflectance, often with the variations occurring over very small spatial scales. Documentation of these landscape properties is a first step in understanding the interplay between the formative processes and response to changing conditions. Characterization of the landscape is also necessary to establish boundary conditions for models which are sensitive to these properties, such as predictive models of atmospheric change that depend on land-atmosphere interactions. Topography, or land surface height, is an important component for many height applications, both to the scientific and commercial sectors. The most accurate global terrain product was produced by the Shuttle Radar Topography Mission (SRTM) launched in 2000; however, elevation data are limited to non-polar regions. The accuracy of SRTM derived elevations range from 5 – 10 m, depending upon the amount of topography and vegetation cover over a particular area. ICESat-2 will provide a global distribution of geodetic measurements (of both the terrain surface and relative canopy heights) which will provide a significant benefit to society through a variety of applications including sea level change monitoring, forest structural mapping and biomass estimation, and improved global digital terrain models.

In addition to producing a global terrain product, monitoring the amount and distribution of above ground vegetation and carbon pools enables improved characterization of the global carbon budget. Forests play a significant role in the terrestrial carbon cycle as carbon pools. Events, such as management activities (Krankina et al. 2012) and disturbances can release carbon stored in forest above

ground biomass (AGB) into the atmosphere as carbon dioxide, a greenhouse gas that contributes to climate change (Ahmed et al. 2013). While carbon stocks in nations with continuous national forest inventories (NFIs) are known, complications with NFI carbon stock estimates exist, including: (1) ground-based inventory measurements are time consuming, expensive, and difficult to collect at large-scales (Houghton 2005; Ahmed et al. 2013); (2) asynchronously collected data; (3) extended time between repeat measurements (Houghton 2005); and (4) the lack of information on the spatial distribution of forest AGB, required for monitoring sources and sinks of carbon (Houghton 2005). Airborne lidar has been used for small studies to capture canopy height and in those studies canopy height variation for multiple forest types is measured to approximately 7 m standard deviation (Hall et al., 2011).

Although the spatial extent and changes to forests can be mapped with existing satellite remote sensing data, the lack of information on forest vertical structure and biomass limits the knowledge of biomass/biomass change within the global carbon budget. Based on the global carbon budget for 2015 (Quere et al., 2015), the largest remaining uncertainties about the Earth's carbon budget are in its terrestrial components, the global residual terrestrial carbon sink, estimated at  $3.0\pm0.8$  GtC/year for the last decade (2005-2014). Similarly, carbon emissions from land-use changes, including deforestation, afforestation, logging, forest degradation and shifting cultivation are estimated at  $0.9\pm0.5$  GtC /year. By providing information on vegetation canopy height globally with a higher spatial resolution than previously afforded by other spaceborne sensors, the ICESat-2 mission can contribute significantly to reducing uncertainties associated with forest vegetation carbon.

Although ICESat-2 is not positioned to provide global biomass estimates due to its profiling configuration and somewhat limited detection capabilities, it is anticipated that the data products for vegetation will be complementary to ongoing biomass and vegetation mapping efforts. Synergistic use of ICESat-2 data with other space-based mapping systems is one solution for extended use of ICESat-2 data. Possibilities include NASA's Global Ecosystems Dynamics Investigation (GEDI) lidar

planned to fly onboard the International Space Station (ISS) or imaging sensors, such as Landsat 8, or NASA/ISRO –NISAR radar mission.

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### 1.2 Photon Counting Lidar

Rather than using an analog, full waveform system similar to what was utilized on the ICESat/GLAS mission, ICESat-2 will employ a photon counting lidar. Photon counting lidar has been used successfully for ranging for several decades in both the science and defense communities. Photon counting lidar systems operate on the concept that a low power laser pulse is transmitted and the detectors used are sensitive at the single photon level. Due to this type of detector, any returned photon whether from the reflected signal or solar background can trigger an event within the detector. A discussion regarding discriminating between signal and background noise photons is discussed later in this document. A question of interest to the ecosystem community is to understand where within the canopy is the photon likely to be reflected. Figure 1.1 is an example of three different laser detector modalities: full waveform, discrete return, and photon counting. Full waveform sensors record the entire temporal profile of the reflected laser energy through the canopy. In contrast, discrete return systems have timing hardware that record the time when the amplitude of the reflected signal energy exceeds a certain threshold amount. A photon counting system, however, will record the arrival time associated with a single photon detection that can occur anywhere within the vertical distribution of the reflected signal. If a photon counting lidar system were to dwell over a surface for a significant number of shots (i.e. hundreds or more), the vertical distribution of the reflected photons will resemble a full waveform. Thus, while an individual photon could be reflected from anywhere within the vertical canopy, the probability distribution function (PDF) of that reflected photon would be the full waveform. Furthermore, the probability of detecting the top of the tree is not as great as detecting reflective surfaces positioned deeper into the canopy where the bulk of leaves and branches are located. As one might imagine, the PDF will differ according to canopy structure and vegetation physiology. For example, the PDF of a conifer tree will look different than broadleaf trees.

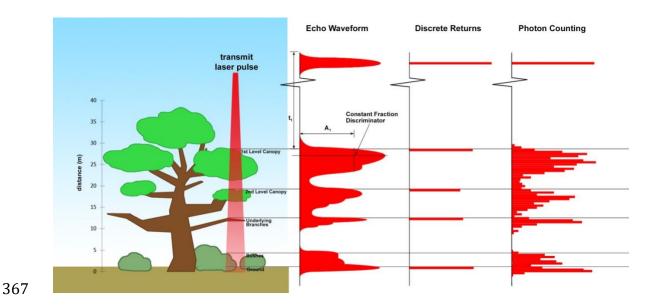


Figure 1.1. Various modalities of lidar detection. Adapted from Harding, 2009.

A cautionary note, the photon counting PDF that is illustrated in Figure 1.1 is merely an illustration if enough photons (i.e. hundreds of photons or more) were to be reflected from a target. In reality, due to the spacecraft speed, ATLAS will record 0 – 4 photons per transmit laser pulse over vegetation.

#### 1.3 The ICESat-2 concept

The Advanced Topographic Laser Altimeter System (ATLAS) instrument designed for ICESat-2 will utilize a different technology than the GLAS instrument used for ICESat. Instead of using a high-energy, single-beam laser and digitizing the entire temporal profile of returned laser energy, ATLAS will use a multi-beam, micropulse laser (sometimes referred to as photon-counting). The travel time of each detected photon is used to determine a range to the surface which, when combined with satellite attitude and pointing information, can be geolocated into a unique XYZ location on or near the Earth's surface. For more information on how the photons from ICESat-2 are geolocated, refer to ATLO3 ATBD. The XYZ positions from ATLAS

are subsequently used to derive surface and vegetation properties. The ATLAS instrument will operate at 532 nm in the green range of the electromagnetic (EM) spectrum and will have a laser repetition rate of 10 kHz. The combination of the laser repetition rate and satellite velocity will result in one outgoing laser pulse approximately every 70 cm on the Earth's surface and each spot on the surface is  $\sim$ 13 m in diameter. Each transmitted laser pulse is split by a diffractive optical element in ATLAS to generate six individual beams, arranged in three pairs (Figure 1.2). The beams within each pair have different transmit energies ('weak' and 'strong', with an energy ratio of approximately 1:4) to compensate for varying surface reflectance. The beam pairs are separated by  $\sim$ 3.3 km in the across-track direction and the strong and weak beams are separated by  $\sim$ 2.5 km in the along-track direction. As ICESat-2 moves along its orbit, the ATLAS beams describe six tracks on the Earth's surface; the array is rotated slightly with respect to the satellite's flight direction so that tracks for the fore and aft beams in each column produce pairs of tracks – each separated by approximately 90 m.

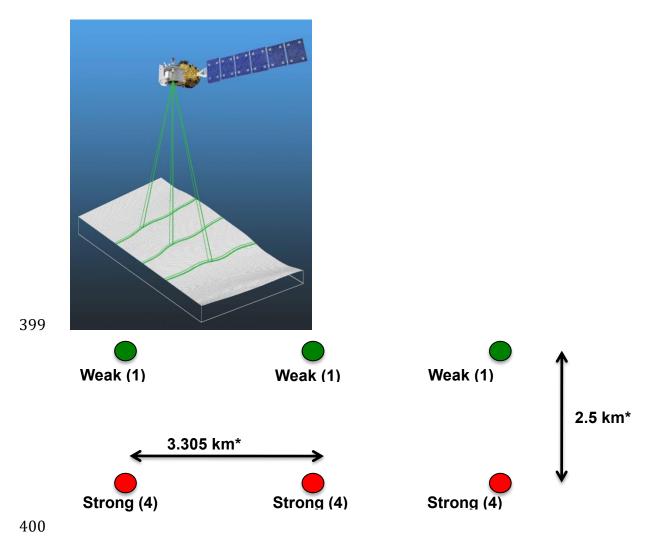


Figure 1.2. Schematic of 6-beam configuration for ICESat-2 mission. The laser energy will be split into 3 laser beam pairs – each pair having a weak spot (1X) and a strong spot (4X).

The motivation behind this multi-beam design is its capability to compute cross-track slopes on a per-orbit basis, which contributes to an improved understanding of ice dynamics. Previously, slope measurements of the terrain were determined via repeat-track and crossover analysis. The laser beam configuration as proposed for ICESat-2 is also beneficial for terrestrial ecosystems compared to GLAS as it enables a denser spatial sampling in the non-polar regions. To achieve a spatial sampling goal of no more than 2 km between equatorial ground tracks, ICESat-2 will be off-nadir pointed a maximum of 1.8 degrees from the reference ground track during the entire mission.

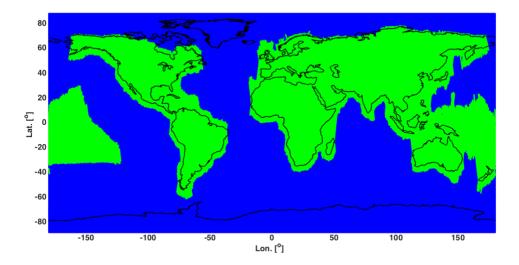


Figure 1.3. Illustration of off-nadir pointing scenarios. Over land (green regions) in the mid-latitudes, ICESat-2 will be pointed away from the repeat ground tracks to increase the density of measurements over terrestrial surfaces.

ICESat-2 is designed to densely sample the Earth's surface, permitting scientists to measure and quantitatively characterize vegetation across vast expanses, e.g., nations, continents, globally. ICESat-2 will acquire synoptic measurements of vegetation canopy height, density, the vertical distribution of photosynthetically active material, leading to improved estimates of forest biomass, carbon, and volume. In addition, the orbital density, i.e., the number of orbits per unit area, at the end of the three year mission will facilitate the production of gridded global products. ICESat-2 will provide the means by which an accurate "snapshot" of global biomass and carbon may be constructed for the mission period.

#### 1.4 Height Retrieval from ATLAS

Light from the ATLAS lasers reaches the earth's surface as flat disks of down-traveling photons approximately 50 cm in vertical extent and spread over approximately 14 m horizontally. Upon hitting the earth's surface, the photons are reflected and scattered in every direction and a handful of photons return to the

ATLAS telescope's focal plane. The number of photon events per laser pulse is a function of outgoing laser energy, surface reflectance, solar conditions, and scattering and attenuation in the atmosphere. For highly reflective surfaces (such as land ice) and clear skies, approximately 10 signal photons from a single strong beam are expected to be recorded by the ATLAS instrument for a given transmit laser pulse. Over vegetated land where the surface reflectance is considerably less than snow or ice surfaces, we expect to see fewer returned photons from the surface. Whereas snow and ice surfaces have high reflectance at 532 nm (typical Lambertian reflectance between 0.8 and 0.98 (Martino, GSFC internal report, 2010)), canopy and terrain surfaces have much lower reflectance (typically around 0.3 for soil and 0.1 for vegetation) at 532 nm. As a consequence we expect to see 1/3 to 1/9 as many photons returned from terrestrial surfaces as from ice and snow surfaces. For vegetated surfaces, the number of reflected signal photon events per transmitted laser pulse is estimated to range between 0 to 4 photons.

The time measured from the detected photon events are used to compute a range, or distance, from the satellite. Combined with the precise pointing and attitude information about the satellite, the range can be geolocated into a XYZ point (known as a geolocated photon) above the WGS-84 reference ellipsoid. In addition to recording photons from the reflected signal, the ATLAS instrument will detect background photons from sunlight which are continually entering the telescope. A primary objective of the ICESat-2 data processing software is to correctly discriminate between signal photons and background photons. Some of this processing occurs at the ATL03 level and some of it also occurs within the software for ATL08. At ATL03, this discrimination is done through a series of three steps of progressively finer resolution with some processing occurring onboard the satellite prior to downlink of the raw data. The ATL03 data product produces a classification between signal and background (i.e. noise) photons, and further discussion on that classification process can be read in the ATLO3 ATBD. In addition, not all geophysical corrections (e.g. ocean tide) are applied to the position of the individual geolocated photons at the ATL03 level, but they are provided on the ATL03 data product if there exists a need to apply them. Thus, in general, all of the heights processed in the ATL08 algorithm consists of the ATL03 heights with respect to the WGS-84 ellipsoid, with geophysical corrections applied, as specified in Chapter 6 of the ATL03 ATBD.

#### 1.5 Accuracy Expected from ATLAS

There are a variety of elements that contribute to the elevation accuracy that are expected from ATLAS and the derived data products. Elevation accuracy is a composite of ranging precision of the instrument, radial orbital uncertainty, geolocation knowledge, forward scattering in the atmosphere, and tropospheric path delay uncertainty. The ranging precision seen by ATLAS will be a function of the laser pulse width, the surface area potentially illuminated by the laser, and uncertainty in the timing electronics. The requirement on radial orbital uncertainty is specified to be less than 4 cm and tropospheric path delay uncertainty is estimated to be 3 cm. In the case of ATLAS, the ranging precision for flat surfaces, is expected to have a standard deviation of approximately 25 cm. The composite of each of the errors can also be thought of as the spread of photons about a surface (see Figure 1.4) and is referred to as the point spread function or Znoise.

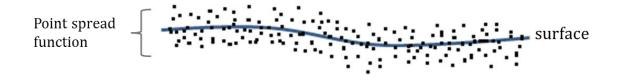


Figure 1.4. Illustration of the point spread function, also referred to as Znoise, for a series of photons about a surface.

The estimates of  $\sigma_{Orbit}$ ,  $\sigma_{troposphere}$ ,  $\sigma_{forwardscattering}$ ,  $\sigma_{pointing}$ , and  $\sigma_{timing}$  for a photon will be represented on the ATL03 data product as the final geolocated accuracy in the X, Y, and Z (or height) direction. In reality, these parameters have different temporal and spatial scales, however until ICESat-2 is on orbit, it is uncertain how these parameters will vary over time. As such, Equation 1.1 may change once the

temporal aspects of these parameters are better understood. For a preliminary quantification of the uncertainties, Equation 1.1 is valid to incorporate the instrument related factors.

489 
$$\sigma_Z = \sqrt{\sigma_{Orbit}^2 + \sigma_{trop}^2 + \sigma_{forwardscattering}^2 + \sigma_{pointing}^2 + \sigma_{timing}^2}$$
 Eqn. 1.1

Although  $\sigma_Z$  on the ATL03 product represents the best understanding of the uncertainty for each geolocated photon, it does not incorporate the uncertainty associated with local slope of the topography. The slope component to the geolocation uncertainty is a function of both the geolocation knowledge of the pointing (which is required to be less than 6.5 m) multiplied by the tangent of the surface slope. In a case of flat topography (<=1 degree slope),  $\sigma_Z$ <= 25 cm, whereas in the case of a 10 degree surface slope,  $\sigma_Z$  =119 cm. The uncertainty associated with the local slope will be combined with  $\sigma_Z$  to produce the term  $\sigma_{Atlas_{Land}}$ .

499 
$$\sigma_{Atlas_{Land}} = \sqrt{\sigma_Z^2 + \sigma_{topo}^2}$$
 Eqn. 1.2

500 
$$\sigma_{topo} = \sigma_{topo} = \sqrt{\left(6.5tan(\theta_{surfaceslope})\right)^2}$$
 Eqn. 1.3

Ultimately, the uncertainty that will be reported on the data product ATL08 will include the  $\sigma_{Atlas_{Land}}$  term and the local rms values of heights computed within each data parameter segment. For example, calculations of terrain height will be made on photons classified as terrain photons (this process is described in the following sections). The uncertainty of the terrain height for a segment is described in Equation 1.4, where the root mean square term of  $\sigma_{Atlas_{Land}}$  and rms of terrain heights are normalized by the number of terrain photons for that given segment.

508 
$$\sigma_{ATL08_{segment}} = \sqrt{\sigma_{Atlas_{Land}}^2 + \sigma_{Zrms_{segment\_class}}^2}$$
 Eqn. 1.4

#### 1.6 Additional Potential Height Errors from ATLAS

Some additional potential height errors in the ATL08 terrain and vegetation product can come from a variety of sources including:

- a. Vertical sampling error. ATLAS height estimates are based on a random sampling of the surface height distribution. Photons may be reflected from anywhere within the PDF of the reflecting surface; more specifically, anywhere from within the canopy. A detailed look at the potential effect of vertical sampling error is provided in Neuenschwander and Magruder (2016).
- b. Background noise. Random noise photons are mixed with the signal photons so classified photons will include random outliers.
- c. Complex topography. The along-track product may not always represent complex surfaces, particularly if the density of ground photons does not support an accurate representation.
- d. Vegetation. Dense vegetation may preclude reflected photon events from reaching the underlying ground surface. An incorrect estimation of the underlying ground surface will subsequently lead to an incorrect canopy height determination.
- e. Misidentified photons. The product from ATL03 combined with additional noise filtering may not identify the correct photons as signal photons.

#### 1.7 Dense Canopy Cases

Although the height accuracy produced from ICESat-2 is anticipated to be superior to other global height products (e.g. SRTM), for certain biomes photon counting lidar data as it will be collected by the ATLAS instrument present a challenge for extracting both the terrain and canopy heights, particularly for areas of dense

vegetation. Due to the relatively low laser power, we anticipate that the along-track signal from ATLAS may lose ground signal under dense forest (e.g. >96% canopy closure) and in situations where cloud cover obscures the terrestrial signal. In areas having dense vegetation, it is likely that only a handful of photons will be returned from the ground surface with the majority of reflections occurring from the canopy. A possible source of error can occur with both the canopy height estimates and the terrain heights if the vegetation is particularly dense and the ground photons were not correctly identified.

## 1.8 Sparse Canopy Cases

Conversely, sparse canopy cases also pose a challenge to vegetation height retrievals. In these cases, expected reflected photon events from sparse trees or shrubs may be difficult to discriminate between solar background noise photons. The algorithms being developed for ATL08 operate under the assumption that signal photons are close together and noise photons will be more isolated in nature. Thus, signal (in this case canopy) photons may be incorrectly identified as solar background noise on the data product. Due to the nature of the photon counting processing, canopy photons identified in areas that have extremely low canopy cover <15% will be filtered out and reassigned as noise photons.

#### **2. ATL08: DATA PRODUCT**

The ATL08 product will provide estimates of terrain height, canopy height, and canopy cover at fine spatial scales in the along-track direction. In accordance with the HDF-driven structure of the ICESat-2 products, the ATL08 product will characterize each of the six Ground Tracks (GT) associated with each Reference Ground Track (RGT) for each cycle and orbit number. Each ground track group has a distinct beam number, distance from the reference track, and transmit energy strength, and all beams will be processed independently using the same sequence of steps described within ATL08. Each ground track group (GT) on the ATL08 product contains subgroups for land and canopy heights segments as well as beam and reference parameters useful in the ATL08 processing. In addition, the labeled photons that are used to determine the data parameters will be indexed back to the ATL03 products such that they are available for further, independent analysis. A layout of the ATL08 HDF product is shown in Figure 2.1. The six GTs are numbered from left to right, regardless of satellite orientation.

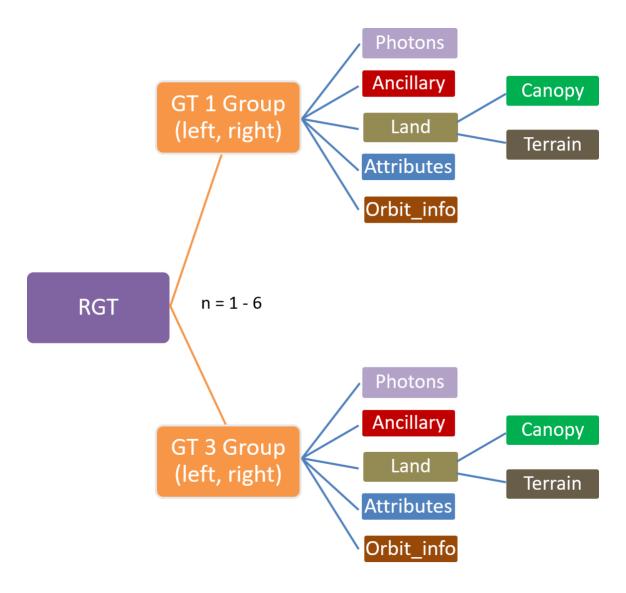


Figure 2.1. HDF5 data structure for ATL08 products

For each data parameter, terrain surface elevation and canopy heights will be provided at a fixed segment size of 100 meters along the ground track. Based on the satellite velocity and the expected number of reflected photons for land surfaces, each segment should have more than 100 signal photons, but in some instances there may be less than 100 signal photons per segment. If a segment has less than 50 classed (i.e., labeled by ATL08 as ground, canopy, or top of canopy) photons we feel this would not accurately represent the surface. Thus, an invalid value will be reported in

all height fields. In the event that there are more than 50 classed photons, but a terrain height cannot be determined due to an insufficient number of ground photons, (e.g. lack of photons penetrating through dense canopy), the only reported terrain height will be the interpolated surface height.

The ATL08 product will be produced per granule based on the ATL03 defined regions (see Figure 2.2). Thus, the ATL08 file/name convention scheme will match the file/naming convention for ATL03 –in attempt for reducing complexity to allow users to examine both data products.

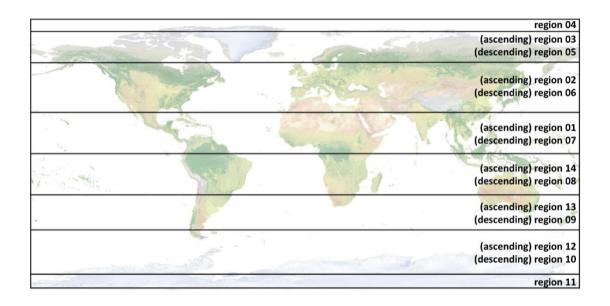


Figure 2.2. ATL03 granule regions; graphic from ATL03 ATBD (Neumann et al.).

The ATL08 product additionally has its own internal regions, which are roughly assigned by continent, as shown by Figure 2.3. For the regions covering Antarctica (regions 7, 8, 9, 10) and Greenland (region 11), the ATL08 algorithm will assume that no canopy is present. These internal ATL08 regions will be noted in the ATL08 product (see parameter atl08\_region in Section 2.4.22). Note that the regions for each ICESat-2 product are not the same.

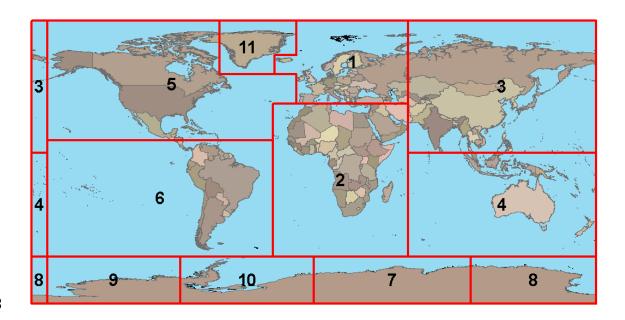


Figure 2.3. ATL08 product regions.

# 2.1 Subgroup: Land Parameters

ATL08 terrain height parameters are defined in terms of the absolute height above the reference ellipsoid.

Table 2.1. Summary table of land parameters on ATL08.

Group	Data type	Description	Source
segment_id_beg	Integer	First along-track segment_id number in 100-m segment	ATL03
segment_id_end	Integer	Last along-track segment_id number in 100-m segment	ATL03
h_te_mean	Float	Mean terrain height for segment	computed
h_te_median	Float	Median terrain height for segment	computed
h_te_min	Float	Minimum terrain height for segment	computed
h_te_max	Float	Maximum terrain height for segment	computed
h_te_mode	Float	Mode of terrain height for segment	computed
h_te_skew	Float	Skew of terrain height for segment	computed

n_te_photons	Intogor	Number of ground photons in	computed
n_te_photons	Integer	segment	computed
h_te_interp	Float	Interpolated terrain surface height at mid-point of segment	computed
h_te_std	Float	Standard deviation of ground heights about the interpolated ground surface	computed
h_te_uncertainty	Float	Uncertainty of ground height estimates. Includes all known uncertainties such as geolocation, pointing angle, timing, radial orbit errors, etc.	computed from Equation 1.4
terrain_slope	Float	Slope of terrain within segment	computed
h_te_best_fit	Float	Best fit terrain elevation at the 100 m segment mid-point location	computed
h_te_best_fit_20m	Float	Best fit terrain elevation at the 20 m geosegment mid-point location	computed
h_te_rh25	float	The relative height from classified canopy photons are sorted into a cumulative distribution, and the height associated with the 98% height above the h_te_bestfit for that segment is reported.	computed
subset_te_flag	Integer	Quality flag indicating the terrain photons populating the 100 m segment statistics are derived from less than 100 m worth of photons	computed
photon_rate_te	Float	Calculated photon rate for ground photons within each segment	computed

## **2.1.1** Georeferenced\_segment\_number\_beg

(parameter = segment\_id\_beg). The first along-track segment\_id in each 100-m segment. Each 100-m segment consists of five sequential 20-m segments provided from the ATL03 product, which are labeled as segment\_id. The segment\_id is a seven digit number that uniquely identifies each along track segment, and is written at the along-track geolocation segment rate (i.e. ~20m along track). The four digit RGT

number can be combined with the seven digit segment\_id number to uniquely define any along-track segment number. Values are sequential, with 0000001 referring to the first segment after the equatorial crossing of the ascending node.

### **2.1.2** Georeferenced segment number end

(parameter = segment\_id\_end). The last along-track segment\_id in each 100-m segment. Each 100-m segment consists of five sequential 20-m segments provided from the ATL03 product, which are labeled as segment\_id. The segment\_id is a seven digit number that uniquely identifies each along track segment, and is written at the along-track geolocation segment rate (i.e. ~20m along track). The four digit RGT number can be combined with the seven digit segment\_id number to uniquely define any along-track segment number. Values are sequential, with 0000001 referring to the first segment after the equatorial crossing of the ascending node.

### **2.1.3** Segment\_terrain\_height\_mean

(parameter = h\_te\_mean). Estimated mean of the terrain height above the reference ellipsoid derived from classified ground photons within the 100 m segment. If a terrain height cannot be directly determined within the segment (i.e. there are not a sufficient number of ground photons), only the interpolated terrain height will be reported. Required input data is classified point cloud (i.e. photons labeled as either canopy or ground in the ATL08 processing). This parameter will be derived from only classified ground photons.

#### **2.1.4** Segment\_terrain\_height\_med

(parameter = h\_te\_median). Median terrain height above the reference ellipsoid derived from the classified ground photons within the 100 m segment. If there are not a sufficient number of ground photons, an invalid value will be reported –no interpolation will be done. Required input data is classified point cloud (i.e. photons labeled as either canopy or ground in the ATL08 processing). This parameter will be derived from only classified ground photons.

### **2.1.5** Segment\_terrain\_height\_min

(parameter = h\_te\_min). Minimum terrain height above the reference ellipsoid derived from the classified ground photons within the 100 m segment. If there are not a sufficient number of ground photons, an invalid value will be reported –no interpolation will be done. Required input data is classified point cloud (i.e. photons labeled as either canopy or ground in the ATL08 processing). This parameter will be derived from only classified ground photons.

#### **2.1.6** Segment\_terrain\_height\_max

(parameter = h\_te\_max). Maximum terrain height above the reference ellipsoid derived from the classified ground photons within the 100 m segment. If there are not a sufficient number of ground photons, an invalid value will be reported –no interpolation will be done. Required input data is classified point cloud (i.e. photons labeled as either canopy or ground in the ATL08 processing). This parameter will be derived from only classified ground photons.

### **2.1.7** Segment\_terrain\_height\_mode

(parameter = h\_te\_mode). Mode of the classified ground photon heights above the reference ellipsoid within the 100 m segment. If there are not a sufficient number of ground photons, an invalid value will be reported –no interpolation will be done. Required input data is classified point cloud (i.e. photons labeled as either canopy or ground in the ATL08 processing). This parameter will be derived from only classified ground photons.

### **2.1.8** Segment\_terrain\_height\_skew

(parameter = h\_te\_skew). The skew of the classified ground photons within the 100 m segment. If there are not a sufficient number of ground photons, an invalid value will be reported –no interpolation will be done. Required input data is classified point cloud (i.e. photons labeled as either canopy or ground in the ATL08 processing). This parameter will be derived from only classified ground photons.

### **2.1.9** Segment\_number\_terrain\_photons

(parameter = n\_te\_photons). Number of terrain photons identified in segment.

### **2.1.10** Segment height\_interp

(parameter = h\_te\_interp). Interpolated terrain surface height above the reference ellipsoid from ATL08 processing at the mid-point of each segment. This interpolated surface is the FINALGROUND estimate (described in section 4.9).

#### **2.1.11** Segment h te std

(parameter = h\_te\_std). Standard deviations of terrain points about the interpolated ground surface within the segment. Provides an indication of surface roughness.

#### **2.1.12** Segment\_terrain\_height\_uncertainty

(parameter = h\_te\_uncertainty). Uncertainty of the mean terrain height for the segment. This uncertainty incorporates all systematic uncertainties (e.g. timing, orbits, geolocation, etc.) as well as uncertainty from errors of identified photons. This parameter is described in Section 1, Equation 1.4. If there are not a sufficient number of ground photons, an invalid value will be reported –no interpolation will be done. Required input data is classified point cloud (i.e. photons labeled as either canopy or ground in the ATL08 processing). This parameter will be derived from only classified ground photons. The  $\sigma_{segmentclass}$  term in Equation 1.4 represents the standard deviation of the terrain height residuals about the FINALGROUND estimate.

#### **2.1.13** Segment terrain slope

(parameter = terrain\_slope). Slope of terrain within each segment. Slope is computed from a linear fit of the terrain photons. It estimates the rise [m] in relief over each segment [100 m]; e.g., if the slope value is 0.04, there is a 4 m rise over the 100 m segment. Required input data are the classified terrain photons.

### **2.1.14** Segment\_terrain\_height\_best\_fit

(parameter = h\_te\_best\_fit). The best fit terrain elevation at the mid-point location of each 100 m segment. The mid-segment terrain elevation is determined by selecting the best of three fits – linear,  $3^{\text{rd}}$  order and  $4^{\text{th}}$  order polynomials – to the terrain photons and interpolating the elevation at the mid-point location of the 100 m segment. For the linear fit, a slope correction and weighting is applied to each ground photon based on the distance to the slope height at the center of the segment.

# **2.1.15** Segment\_terrain\_height\_25

(parameter = h\_te\_rh25). The terrain elevation from the 25% height. The classified ground photons are sorted into a cumulative distribution and the height associated with the 25% height for that segment is reported.

# **2.1.16** Subset\_te\_flag {1:5}

(parameter = subset\_te\_flag). This flag indicates the quality distribution of identified terrain photons within each 100 m on a geosegment basis. The purpose of this flag is to provide the user with an indication whether the photons contributing to the terrain estimate are evenly distributed or only partially distributed (i.e. due to cloud cover or signal attenuation). A 100 m ATL08 segment is comprised of 5 geosegments and we are populating a flag for each geosegment. subset\_te\_flags:

-1: no data within geosegment available for analysis

0: indicates no ground photons within geosegment

1: indicates ground photons within geosegment

For example, an 100 m ATL08 segment might have the following subset\_te\_flags: {-1 -1 0 1 1} which would translate that no signal photons (canopy or ground) were available for processing in the first two geosegments. Geosegment 3 was found to have photons, but none were labeled as ground photons. Geosegment 4 and 5 had valid labeled ground photons. Again, the motivation behind this flag is to

inform the user that, in this example, the 100 m estimate are being derived from only 40 m worth of data.

# 2.1.17 Segment Terrain Photon Rate

(parameter = photon\_rate\_te). This value indicates the terrain photon rate within each ATL08 segment. This value is calculated as the total number of terrain photons divided by the total number of laser shots within each ATL08 segment. The number of laser shots is defined as the number of unique Delta\_Time values within each segment.

### **2.1.18** Terrain Best Fit GeoSegment {1:5}

(parameter = h\_te\_best\_fit\_20m). The best fit terrain elevation at the mid-point location of each 20 m geosegment. The mid-segment terrain elevation is determined by selecting the best of three fits – linear,  $3^{rd}$  order and  $4^{th}$  order polynomials – to the terrain photons and interpolating the elevation at each 20 m along a 100 m segment. For the linear fit, a slope correction and weighting is applied to each ground photon based on the distance to the slope height at the center of the segment. For segments that do not have a sufficient number of photons, an invalid (or fill) value will be reported. Each 20 m geo-segment shall have 10 signal photons as a minimum number to be used for calculations and a minimum of 3 terrain photons are required to estimate a height.

#### 2.2 Subgroup: Vegetation Parameters

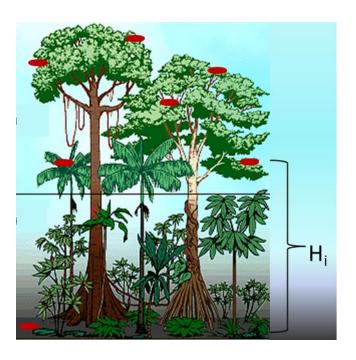
Canopy parameters will be reported on the ATL08 data product in terms of both the absolute height above the reference ellipsoid as well as the relative height above an estimated ground. The relative canopy height,  $H_i$ , is computed as the height from an identified canopy photon minus the interpolated ground surface for the same horizontal geolocation (see Figure 2.3). Thus, each identified signal photon above an interpolated surface (including a buffer distance based on the instrument point spread function) is by default considered a canopy photon. Canopy parameters will

only be computed for segments where more than 5% of the classed photons are classified as canopy photons.



745

746



748

750

751

Figure 2.4. Illustration of canopy photons (red dots) interaction in a vegetated area.

Relative canopy heights, H<sub>i</sub>, are computed by differencing the canopy photon height from an interpolated terrain surface.

Table 2.2. Summary table of canopy parameters on ATL08.

Group	Data type	Description	Source
segment_id_beg	Integer	First along-track segment_id number in 100-m segment	ATL03
segment_id_end	Integer	Last along-track segment_id number in 100-m segment	ATL03
canopy_h_metrics_abs	Float	Absolute (H##) canopy height metrics calculated at the following percentiles: 5, 10, 15, 20, 25, 30, 35, 40, 45, 50, 55, 60, 65, 70, 75, 80, 85, 90, 95.	computed
canopy_h_metrics	Float	Relative (RH##) canopy height metrics calculated at the following percentiles: 5, 10, 15, 20, 25, 30, 35, 40, 45, 50, 55, 60, 65, 70, 75, 80, 85, 90, 95.	computed

h_canopy_abs	Float	98% height of all the individual	computed
_ 10_		absolute canopy heights (height above	
		WGS84 ellipsoid) for segment.	
h_canopy	Float	98% height of all the individual relative	computed
		canopy heights (height above terrain) for segment.	
h_canopy_20m	Float	98% height of all the individual relative	
		canopy heights (height above terrain)	
		for 20m geosegment.	
h_mean_canopy_abs	Float	Mean of individual absolute canopy	computed
h_mean_canopy	Float	heights within segment Mean of individual relative canopy	computed
n_mean_eanopy	11000	heights within segment	computed
h_dif_canopy	Float	Difference between h_canopy and	computed
		canopy_h_metrics(50)	
h_min_canopy_abs	Float	Minimum of individual absolute canopy	computed
h_min_canopy	Float	heights within segment Minimum of individual relative canopy	computed
n_mm_canopy	11000	heights within segment	compated
h_max_canopy_abs	Float	Maximum of individual absolute	computed
		canopy heights within segment. Should	
h may canany	Float	be equivalent to H100	computed
h_max_canopy	rioat	Maximum of individual relative canopy heights within segment. Should be	computed
		equivalent to RH100	
h_canopy_uncertainty	Float	Uncertainty of the relative canopy	computed
	nı .	height (h_canopy)	. 1
canopy_openness	Float	STD of relative heights for all photons classified as canopy photons within the	computed
		segment to provide inference of canopy	
		openness	
toc_roughness	Float	STD of relative heights of all photons	computed
		classified as top of canopy within the	
h_canopy_quad	Float	segment Quadratic mean canopy height	computed
n_ca_photons	Integer4	Number of canopy photons within 100	computed
<u>-</u> -		m segment	<b>F</b>
n_toc_photons	Integer4	Number of top of canopy photons	computed
	nı .	within 100 m segment	. 1
centroid_height	Float	Absolute height above reference ellipsoid associated with the centroid of	computed
		all signal photons	
canopy_rh_conf	Integer	Canopy relative height confidence flag	computed
		based on percentage of ground and	
		canopy photons within a segment: 0	
		(<5% canopy), 1 (>5% canopy, <5% ground), 2 (>5% canopy, >5% ground)	
		grounds, 2 (>3% canopy, >3% ground)	

subset_can_flag	Integer	Quality flag indicating the canopy photons populating the 100 m segment statistics are derived from less than 100 m worth of photons	computed
photon_rate_can	Float	Photon rate of canopy photons within each 100 m segment	computed

#### **2.2.1** Georeferenced segment number beg

(parameter = segment\_id\_beg). The first along-track segment\_id in each 100-m segment. Each 100-m segment consists of five sequential 20-m segments provided from the ATL03 product, which are labeled as segment\_id. The segment\_id is a seven digit number that uniquely identifies each along track segment, and is written at the along-track geolocation segment rate (i.e. ~20m along track). The four digit RGT number can be combined with the seven digit segment\_id number to uniquely define any along-track segment number. Values are sequential, with 0000001 referring to the first segment after the equatorial crossing of the ascending node.

### **2.2.2** Georeferenced\_segment\_number\_end

(parameter = segment\_id\_end). The last along-track segment\_id in each 100-m segment. Each 100-m segment consists of five sequential 20-m segments provided from the ATL03 product, which are labeled as segment\_id. The segment\_id is a seven digit number that uniquely identifies each along track segment, and is written at the along-track geolocation segment rate (i.e. ~20m along track). The four digit RGT number can be combined with the seven digit segment\_id number to uniquely define any along-track segment number. Values are sequential, with 0000001 referring to the first segment after the equatorial crossing of the ascending node.

# **2.2.3** Canopy\_height\_metrics\_abs

(parameter = canopy\_h\_metrics\_abs). The absolute height metrics (H##) of classified canopy photons (labels 2 and 3) above the ellipsoid. The height metrics are sorted based on a cumulative distribution and calculated at the following percentiles: 10, 15, 20, 25, 30, 35, 40, 45, 50, 55, 60, 65, 70, 75, 80, 85, 90, 95. These height metrics

are often used in the literature to characterize vertical structure of vegetation. One important distinction of these canopy height metrics compared to those derived from other lidar systems (e.g., LVIS or GEDI) is that the ICESat-2 canopy height metrics are heights above the ground surface. These metrics do not include the ground photons. Required input data are the relative canopy heights of all canopy photons above the estimated terrain surface and the mid-segment elevation. The absolute canopy heights metrics are determined by adding the relative canopy height metric to the best-fit terrain (h\_te\_bestfit). For cases where the h\_te\_bestfit is invalid, the cumulative distribution will be calculated for the absolute canopy heights (not the relative canopy heights) and those cumulative heights will be reported.

### **2.2.4** Canopy\_height\_metrics

(parameter = canopy\_h\_metrics). Relative height metrics above the estimated terrain surface (RH##) of classified canopy photons (labels 2 and 3). The height metrics are sorted based on a cumulative distribution and calculated at the following percentiles: 10, 15, 20, 25, 30, 35, 40, 45, 50, 55, 60, 65, 70, 75, 80, 85, 90, 95. These height metrics are often used in the literature to characterize vertical structure of vegetation. One important distinction of these canopy height metrics compared to those derived from other lidar systems (e.g., LVIS or GEDI) is that the ICESat-2 canopy height metrics are heights above the ground surface. These metrics do not include the ground photons. Required input data are relative canopy heights above the estimated terrain surface for all canopy photons.

### **2.2.5** Absolute\_segment\_canopy\_height

(parameter = h\_canopy\_abs). The absolute 98% height of classified canopy photon heights (labels 2 and 3) above the ellipsoid. The relative height from classified canopy photons are sorted into a cumulative distribution, and the height associated with the 98% height above the h\_te\_bestfit for that segment is reported. For cases

where the h\_te\_bestfit is invalid, the cumulative distribution will be calculated for the absolute canopy heights and the 98% absolute height will be reported.

### **2.2.6** Segment\_canopy\_height

(parameter = h\_canopy). The relative 98% height of classified canopy photon heights (labels 2 and 3) above the estimated terrain surface. Relative canopy heights have been computed by differencing the canopy photon height from the estimated terrain surface in the ATL08 processing. The relative canopy heights are sorted into a cumulative distribution, and the height associated with the 98% height is reported.

### **2.2.7** canopy\_height GeoSegment {1:5}

(parameter = h\_canopy\_20m). The relative 98% height of classified canopy photon heights (labels 2 and 3) above the estimated terrain surface in each 20 m geosegment. Relative canopy heights have been computed by differencing the canopy photon height from the estimated terrain surface in the ATL08 processing. The relative canopy heights are sorted into a cumulative distribution, and the height associated with the 98% height is reported. For segments that do not have a sufficient number of photons, an invalid (or fill) value will be reported. Each 20 m geo-segment shall have 10 signal photons as a minimum number to be used for calculations and a minimum of 3 canopy photons are required to estimate a height.

#### **2.2.8** Absolute segment mean canopy

(parameter = h\_mean\_canopy\_abs). The absolute mean canopy height for the segment. relative canopy heights are the photons heights for canopy photons (labels 2 and 3) above the estimated terrain surface. These relative heights are averaged and then added to h\_te\_bestfit.

### **2.2.9** Segment\_mean\_canopy

(parameter = h\_mean\_canopy). The mean canopy height for the segment. Relative canopy heights have been computed by differencing the canopy photon height (labels 2 and 3) from the estimated terrain surface in the ATL08 processing. These heights are averaged.

#### **2.2.10** Segment\_dif\_canopy

(parameter = h\_dif\_canopy). Difference between h\_canopy and canopy\_h\_metrics(50). This parameter is one metric used to describe the vertical distribution of the canopy within the segment.

### **2.2.11** Absolute\_segment\_min\_canopy

(parameter = h\_min\_canopy\_abs). The minimum absolute canopy height for the segment. Relative canopy heights are the photons heights for canopy photons (labels 2 and 3) above the estimated terrain surface. Required input data is classified point cloud (i.e. photons labeled as either canopy or ground in the ATL08 processing). The minimum relative canopy height for each segment is added to h\_te\_bestfit and reported as the absolute minimum canopy height.

#### **2.2.12** Segment min canopy

(parameter = h\_min\_canopy). The minimum relative canopy height for the segment. Canopy heights are the photons heights for canopy photons (labels 2 and 3) differenced from the estimated terrain surface. Required input data is classified point cloud (i.e. photons labeled as either canopy or ground in the ATL08 processing).

#### **2.2.13** Absolute\_segment\_max\_canopy

(parameter = h\_max\_canopy\_abs). The maximum absolute canopy height for the segment. This parameter is equivalent to H100 metric reported in the literature. This parameter, however, has the potential for error as random solar background noise may not have been fully rejected. It is recommended that h\_canopy or h\_canopy\_abs (i.e., the 98% canopy height) be considered as the top of canopy

measurement. Required input data is classified point cloud (i.e. photons labeled as either canopy or ground in the ATL08 processing). The absolute max canopy height is the maximum relative canopy height added to h\_te\_bestfit.

### **2.2.14** Segment\_max\_canopy

(parameter = h\_max\_canopy). The maximum relative canopy height for the segment. Canopy heights are the photons heights for canopy photons (labels 2 and 3) differenced from the estimated terrain surface. This product is equivalent to RH100 metric reported in the literature. This parameter, however, has the potential for error as random solar background noise may not have been fully rejected. It is recommended that h\_canopy or h\_canopy\_abs (i.e., the 98% canopy height) be considered as the top of canopy measurement. Required input data is classified point cloud (i.e. photons labeled as either canopy or ground in the ATL08 processing).

### **2.2.15** Segment\_canopy\_height\_uncertainty

(parameter = h\_canopy\_uncertainty). Uncertainty of the relative canopy height for the segment. This uncertainty incorporates all systematic uncertainties (e.g. timing, orbits, geolocation, etc.) as well as uncertainty from errors of identified photons. This parameter is described in Section 1, Equation 1.4. If there are not a sufficient number of ground photons, an invalid value will be reported –no interpolation will be done. In the case for canopy height uncertainty, the parameter  $\sigma_{segmentclass}$  is comprised of both the terrain uncertainty within the segment but also the top of canopy residuals. Required input data is classified point cloud (i.e. photons labeled as either top of canopy or ground in the ATL08 processing). This parameter will be derived from only classified top of canopy photons, label = 3. The canopy height uncertainty is derived from Equation 1.4, shown below as Equation 1.5, represents the standard deviation of the terrain points and the standard deviation of the top of canopy height photons.

$$\sigma_{ATL08_{segment\_ch}} = \frac{\sqrt{\sigma_{Atlas_{Land}}^2 + \sigma_{Zrms_{segment\_terrain}}^2 + \sigma_{Zrms_{segment\_toc}}^2}}{n_{photons_{segment_{terrain}}} + n_{photons_{segment_{toc}}}}$$
Eqn 1.5

#### **2.2.16** Segment\_canopy\_openness

(parameter = canopy\_openness). Standard deviation of relative canopy heights within each segment. This parameter will potentially provide an indicator of canopy openness (label = 2 and 3) as a greater standard deviation of heights indicates greater penetration of the laser energy into the canopy. Required input data is classified point cloud (i.e. photons labeled as either canopy or ground in the ATL08 processing).

### **2.2.17** Segment\_top\_of\_canopy\_roughness

(parameter = toc\_roughness). Standard deviation of relative top of canopy heights (label = 3) within each segment. This parameter will potentially provide an indicator of canopy variability. Required input data is classified point cloud (i.e. photons labeled as the top of the canopy in the ATL08 processing).

# **2.2.18** Segment\_canopy\_quadratic\_height

(parameter = h\_canopy\_quad). The quadratic mean relative height of relative canopy heights. The quadratic mean height is computed as:

$$qmh = \sqrt{\sum_{i=1}^{n\_ca\_photons} \frac{h_i^2}{n\_ca\_photons}}$$

### **2.2.19** Segment\_number\_canopy\_photons

(parameter =  $n_{ca}$ \_photons). Number of canopy photons (label 2 and 3) within each segment. Required input data is classified point cloud (i.e. photons labeled as either canopy or ground in the ATL08 processing).

# 904 **2.2.20** Segment\_number\_top\_canopy\_photons 905 (parameter = n\_toc\_photons). Number of top of canopy photons (label = 3) 906 within each segment. Required input data is classified point cloud (i.e. photons 907 labeled as top of canopy in the ATL08 processing). 908 2.2.21 Centroid height 909 (parameter = centroid\_height). Optical centroid of all photons classified as 910 either canopy or ground points (label = 1 2 or 3) within a segment. The heights used 911 in this calculation are absolute heights above the reference ellipsoid. This parameter 912 is equivalent to the centroid height produced on ICESat GLA14. 913 **2.2.22** Segment rel canopy conf 914 (parameter = canopy rh conf). Canopy relative height confidence flag based 915 on percentage of ground photons and percentage of canopy photons (label 2 and 3), 916 relative to the total classified (ground and canopy, label = 1 2 and 3) photons within 917 a segment: 0 (<5% canopy), 1 (>5% canopy and <5% ground), 2 (>5% canopy and 918 >5% ground). This is a measure based on the quantity, not the quality, of the 919 classified photons in each segment. 920 **2.2.23** Subset can flag {1:5} 921 (parameter = subset can flag). This flag indicates the distribution of identified 922 canopy photons (label 2 and 3) within each 100 m. The purpose of this flag is to 923 provide the user with an indication whether the photons contributing to the canopy 924 height estimates are evenly distributed or only partially distributed (i.e. due to cloud 925 cover or signal attenuation). A 100 m ATL08 segment is comprised of 5 geo-segments. 926 subset\_can\_flags: 927 -1: no data within geosegment available for analysis 928 0: indicates no canopy photons within geosegment

1: indicates canopy photons within geosegment

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For example, a 100 m ATL08 segment might have the following subset\_can\_flags: {-1 -1 -1 1 1} which would translate that no photons (canopy or ground) were available for processing in the first three geosegments. Geosegment 4 and 5 had valid labeled canopy photons. Again, the motivation behind this flag is to inform the user that, in this example, the 100 m estimate are being derived from only 40 m worth of data.

### 2.2.24 Segment Canopy Photon Rate

(parameter = photon\_rate\_can). This value indicates the canopy photon rate within each ATL08 segment. This value is calculated as the total number of canopy photons (label =2 and 3) divided by the total number of unique laser shots within each ATL08 segment. The number of laser shots is defined as the number of unique Delta Time values within each segment.

#### 2.3 Subgroup: Photons

The subgroup for photons contains the classified photons that were used to generate the parameters within the land or canopy subgroups. Each photon that is identified as being likely signal will be classified as: 0 = noise, 1 = ground, 2 = canopy, or 3 = top of canopy. The index values for each classified photon will be provided such that they can be extracted from the ATL03 data product for independent evaluation.

Table 2.3. Summary table for photon parameters for the ATL08 product.

Group	Data Type	Description	Source
classed_PC_indx	Float	Indices of photons tracking back to ATL03 that surface finding software identified and used within the creation of the data products.	ATL03
classed_PC_flag	Integer	Classification flag for each photon as either noise,	computed

ph_segment_id	Integer	ground, canopy, or top of canopy. Georeferenced bin number (20-m) associated with each	ATL03
ph_h	Float	photon Height of photon above interpolated ground surface	computed
d_flag	Integer	Flag indicating whether DRAGANN labeled the photon as noise or signal	computed

#### **2.3.1** Indices of classed photons

(parameter = classed\_PC\_indx). Indices of photons tracking back to ATL03 that surface finding software identified and used within the creation of the data products for a given segment.

#### **2.3.2** Photon\_class

(parameter = classed\_PC\_flag). Classification flags for a given segment. 0 = noise, 1 = ground, 2 = canopy, 3 = top of canopy. The final ground and canopy classification are flags 1-3. The full canopy is the combination of flags 2 and 3.

#### **2.3.3** Georeferenced\_segment\_number

(parameter = ph\_segment\_id). The segment\_id associated with every photon in each 100-m segment. Each 100-m segment consists of five sequential 20-m segments provided from the ATL03 product, which are labeled as segment\_id. The segment\_id is a seven digit number that uniquely identifies each along track segment, and is written at the along-track geolocation segment rate (i.e. ~20m along track). The four digit RGT number can be combined with the seven digit segment\_id number to uniquely define any along-track segment number. Values are sequential, with 0000001 referring to the first segment after the equatorial crossing of the ascending node.

# **2.3.4** Photon Height

(parameter = ph\_h). Height of the photon above the interpolated ground surface at the location of the photon.

### **2.3.5** DRAGANN\_flag

(parameter =  $d_flag$ ). Flag indicating the labeling of DRAGANN noise filtering for a given photon. 0 = noise, 1 = signal.

## 2.4 Subgroup: Reference data

The reference data subgroup contains parameters and information that are useful for determining the terrain and canopy heights that are reported on the product. In addition to position and timing information, these parameters include the reference DEM height, reference landcover type, and flags indicating water or snow.

Table 2.4. Summary table for reference parameters for the ATL08 product.

Group	Data Type	Description	Source
segment_id_beg	Integer	First along-track segment_id number in 100-m segment	ATL03
segment_id_end	Integer	Last along-track segment_id number in 100-m segment	ATL03
latitude	Float	Center latitude of signal photons within each segment	ATL03
longitude	Float	Center longitude of signal photons within each segment	ATL03
delta_time	Float	Mid-segment GPS time in seconds past an epoch. The epoch is provided in the metadata at the file level	ATL03
delta_time_beg	Float	Delta time of the first photon in the segment	ATL03
delta_time_end	Float	Delta time of the last photon in the segment	ATL03
night_flag	Integer	Flag indicating whether the measurements were	computed

		acquired during night time conditions	
dem_h	Float4	Reference DEM elevation	external
dem_flag		Source of reference DEM	external
dem_removal_flag	Integer	Quality check flag to indicate	computed
	C	> 20% photons removed due	•
		to large distance from dem_h	
h_dif_ref	Float4	Difference between	computed
		h_te_median and dem_h	•
terrain_flg	Integer	Terrain flag quality check to	computed
_ 0	C	indicate a deviation from the	•
		reference DTM	
segment_landcover	Integer4	Reference landcover for	external
_		segment derived from best	
		global landcover product	
		available	
segment_watermask	Integer4	Water mask indicating inland	external
		water produced from best	
		sources available	
segment_snowcover	Integer4	Daily snow cover mask	external
		derived from best sources	
urban_flag	Integer	Flag indicating segment is	external
		located in an urban area	
surf_type	Integer1	Flags describing surface	ATL03
		types: 0=not type, 1=is type.	
		Order of array is land, ocean,	
		sea ice, land ice, inland	
		water.	
atl08_region	Integer	ATL08 region(s)	computed
		encompassed by ATL03	
		granule being processed	
last_seg_extend	Float	The distance (km) that the	computed
		last ATL08 processing	
		segment in a file is either	
		extended or overlapped with	
		the previous ATL08	
	<b>.</b>	processing segment	. 1
brightness_flag	Integer	Flag indicating that the	computed
		ground surface is bright (e.g.	
		snow-covered or other	
		bright surfaces)	

### **2.4.1** Georeferenced\_segment\_number\_beg

(parameter = segment\_id\_beg). The first along-track segment\_id in each 100-m segment. Each 100-m segment consists of five sequential 20-m segments provided from the ATL03 product, which are labeled as segment\_id. The segment\_id is a seven digit number that uniquely identifies each along track segment, and is written at the along-track geolocation segment rate (i.e. ~20m along track). The four digit RGT number can be combined with the seven digit segment\_id number to uniquely define any along-track segment number. Values are sequential, with 0000001 referring to the first segment after the equatorial crossing of the ascending node.

### **2.4.2** Georeferenced\_segment\_number\_end

(parameter = segment\_id\_end). The last along-track segment\_id in each 100-m segment. Each 100-m segment consists of five sequential 20-m segments provided from the ATL03 product, which are labeled as segment\_id. The segment\_id is a seven digit number that uniquely identifies each along track segment, and is written at the along-track geolocation segment rate (i.e. ~20m along track). The four digit RGT number can be combined with the seven digit segment\_id number to uniquely define any along-track segment number. Values are sequential, with 0000001 referring to the first segment after the equatorial crossing of the ascending node.

#### **2.4.3** Segment latitude

(parameter = latitude). Center latitude of signal photons within each segment. Each 100 m segment consists of 5 20m ATL03 geosegments. In most cases, there will be signal photons in each of the 5 geosegments necessary for calculating a latitude value. For instances where the 100 m ATL08 is not fully populated with photons (e.g. photons drop out due to clouds or signal attenuation), the latitude will be interpolated to the mid-point of the 100 m segment. To implement this interpolation, we confirm that each 100 m segment is comprised of at least 3 unique ATL03 geosegments IDs, indicating that data is available near the mid-point of the land segment. If less than 3 ATL03 segments are available, the coordinate is interpolated based on the ratio of delta time at the centermost ATL03 segment and that of the centermost photon, thus

applying the centermost photon's coordinates to represent the land segment with a slight adjustment. In some instances, the latitude and longitude will require extrapolation to estimate a mid-100 m segment location. It is possible that in these extremely rare cases, the latitude and longitude could not represent the true center of the 100 m segment. We encourage the user to investigate the parameters segment\_te\_flag and segment\_can\_flag which provide information as to the number and distribution of signal photons within each 100 m segment.

### **2.4.4** Geosegment\_latitude{1:5}

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1021 (parameter = latitude\_20m). Interpolated center latitude of each 20 m 1022 geosegment.

#### **2.4.5** Segment longitude

(parameter = longitude). Center longitude of signal photons within each segment. Each 100 m segment consists of 5 20m geosegments. In most cases, there will be signal photons in each of the 5 geosegments necessary for calculating a longitude value. For instances where the 100 m ATL08 is not fully populated with photons (e.g. photons drop out due to clouds or signal attenuation), the latitude will be interpolated to the mid-point of the 100 m segment. To implement this interpolation, we confirm that each 100 m segment is comprised of at least 3 unique ATL03 geosegments IDs, indicating that data is available near the mid-point of the land segment. If less than 3 ATL03 segments are available, the coordinate is interpolated based on the ratio of delta time at the centermost ATL03 segment and that of the centermost photon, thus applying the centermost photon's coordinates to represent the land segment with a slight adjustment. In some instances, the latitude and longitude will require extrapolation to estimate a mid-100 m segment location. It is possible that in these extremely rare cases, the latitude and longitude could not represent the true center of the 100 m segment. We encourage the user to investigate the paramters segment\_te\_flag and segment\_can\_flag which provide information as to the number and distribution of signal photons within each 100 m segment.

1041	<b>2.4.6</b> Geosegment_longitude{1:5}
1042	(parameter = longitude_20m). Interpolated center longitude of each 20 m
1043	geosegment.
1044	
1045	2.4.7 Delta_time
1046	(parameter = delta_time). Mid-segment GPS time for the segment in seconds
1047	past an epoch. The epoch is listed in the metadata at the file level.
1048	2.4.8 Delta_time_beg
1049	(parameter = delta_time_beg). Delta time for the first photon in the segment
1050	in seconds past an epoch. The epoch is listed in the metadata at the file level.
1051	2.4.9 Delta_time_end
1052	(parameter = delta_time_end). Delta time for the last photon in the segment
1053	in seconds past an epoch. The epoch is listed in the metadata at the file level.
1054	<b>2.4.10</b> Night_Flag
1055	(parameter = night_flag). Flag indicating the data were acquired in night
1056	conditions: $0 = \text{day}$ , $1 = \text{night}$ . Night flag is set when solar elevation is below 0.0
1057	degrees.
1058	2.4.11 Segment_reference_DTM
1059	(parameter = dem_h). Reference terrain height value for segment determined
1060	by the "best" DEM available based on data location. All heights in ICESat-2 are
1061	referenced to the WGS 84 ellipsoid unless clearly noted otherwise. DEM is taken from
1062	a variety of ancillary data sources: MERIT, GIMP, GMTED, MSS. The DEM source flag
1063	indicates which source was used.

### **2.4.12** Segment\_reference\_DEM\_source

1065 (parameter = dem\_flag). Indicates source of the reference DEM height. Values: 1066 0=None, 1=GIMP, 2=GMTED, 3=MSS, 4=MERIT.

### **2.4.13** Segment\_reference\_DEM\_removal\_flag

1068 (parameter = dem\_removal\_flag). Quality check flag to indicate > 20% classified photons removed from land segment due to large distance from dem\_h.

### **2.4.14** Segment\_terrain\_difference

(parameter = h\_dif\_ref). Difference between h\_te\_median and dem\_h. Since the mean terrain height is more sensitive to outliers, the median terrain height will be evaluated against the reference DEM. This parameter will be used as an internal data quality check with the notion being that if the difference exceeds a threshold (TBD) a terrain quality flag (terrain\_flg) will be triggered.

# **2.4.15** Segment\_terrain flag

(parameter = terrain\_flg). Terrain flag to indicate confidence in the derived terrain height estimate. If  $h_{dif_{ref}}$  exceeds a threshold (TBD) the terrain\_flg parameter will be set to 1. Otherwise, it is 0.

### **2.4.16** Segment\_landcover

(parameter = segment\_landcover). Updating the segment landcover with the 2019 Copernicus Landcover 100 m discrete landcover product which incorporates 23 discrete landcover classes which follow the UN-FAO's Land Cover Classification System. The ATL08 landcover segment will be the Copernicus Landcover value at the segment latitude/longitude. <a href="https://land.copernicus.eu/global/products/lc">https://land.copernicus.eu/global/products/lc</a> (<a href="https://land.copernicus.eu/global/products/lc">https://land.copernicus.eu/global/products/lc</a> (<a href="https://doi.org/10.5281/zenodo.3939050">https://doi.org/10.5281/zenodo.3939050</a>).

Map Code	Landcover Class	Definition according to UN LCCS
0	No data	

111	Closed forest, evergreen needle leaf	Tree canopy >70%, almost all needle leaf trees remain green all year. Canopy is never without green foliage		
113	Closed forest, deciduous needle leaf	Tree canopy >70%, consists of seasonal needle leaf communities with an annual cycle of leaf-on and leaf-off periods.		
112	Closed forest, evergreen broad leaf	Tree canopy >70%, almost all broadleaf trees remain green year round. Canopy is never without green foliage		
114	Closed forest, deciduous broad leaf	Tree canopy >70%, consists of seasonal broad leaf communities with an annual cycle of leaf-on and leaf-off periods.		
115	Closed forest, mixed	Closed forest, mix of types		
116	Closed forest, unknown	Closed forest, not matching any of the other definitions		
121	Open forest, evergreen needle leaf	Top layer- trees 15-70% and second layer mixed of shrubs and grassland, almost all needle leaf trees remain green all year. Canopy is never without green foliage		
123	Open forest, deciduous needle leaf	Top layer- trees 15-70% and second layer mixed of shrubs and grassland, consists of seasonal needle leaf tree communities with an annual cycle of leaf-on and leaf-off		
122	Open forest, evergreen broad leaf	Top layer- trees 15-70% and second layer mixed of shrubs and grassland, almost all broad leaf trees remain green all year. Canopy is never without green foliage		
124	Open forest, deciduous broad leaf	Top layer- trees 15-70% and second layer mixed of shrubs and grassland, consists of seasonal broad leaf tree communities with an annual cycle of leaf-on and leaf-off		
125	Open forest, mixed	Open forest, mix of types		
126	Open forest, unknown	Open forest, not matching any of the other definitions		
20	Shrubs	Woody perennial plants with persistent and woody stems and without a main stem being less than 5m. The shrub foliage can be either evergreen or deciduous.		
30	Herbaceous	Plants without persistent stems or shoots above ground and lacking firm structure. Tree and shrub cover is less than 10%		
90	Herbaceous Wetland	Lands with a permanent mixture of water and herbaceous or woody vegetation. The vegetation can be present in salt, brackish, or fresh water.		
100	Moss and lichen	Moss and lichen		
60	Bare/sparse vegetation	Lands with exposed soil, sand, or rocks and never has more than 10% vegetation cover during any time of the year		
40	Cultivated and managed vegetation/agriculture	Lands covered with temporary crops followed by harvest and a bare soil period.		
50	Urban/built up	Land covered by buildings or other man-made structures		
70	Snow and ice	Land under snow or ice throughout the year		
80	Permanent water bodies	Lakes, reservoirs, and rivers. Can be either fresh or salt-water bodies		

200	Open sea	Oceans, seas. Can be either fresh of salt-water bodies.
	L	<u>I</u>
2.4	.17 Segment_Woody Vegeta	ntion Fractional Cover
(1	parameter = segment_cove	r). Woody vegetation fractional cover derived
from the	2019 Copernicus 100 m shr	rub and forest fractional cover data products. The
woody c	over fractional cover is the s	simple addition of the forest fractional cover with
the shru	b fractional cover. The ATL	08 woody vegetation fractional cover value shall
be the p	ixel value at the segment la	titude/longitude. The Copernicus data products
can be fo	ound at https://lcviewer.vite	o.be/download
2.4	.18 Segment_watermask	
<b>(</b> 1	parameter = segment_water	rmask). Water mask (i.e., flag) indicating inland
water as	referenced from the Globa	l Raster Water Mask at 250 m spatial resolution
(Carroll	et al, 2009; available online	e at http://glcf.umd.edu/data/watermask/). 0 =
no water	; 1 = water.	
2.4	.19 Segment_snowcover	
1)	oarameter = segment_sno	owcover). Daily snowcover mask (i.e., flag)
indicatin	g a likely presence of snow	or ice within each segment produced from best
available	e source used for reference.	The snow mask will be the same snow mask as
used for	ATL09 Atmospheric Proc	ducts: NOAA snow-ice flag. 0=ice free water;
1=snow	free land; 2=snow; 3=ice.	
2.4	.20 Urban_flag	
2.4	.21 (parameter = urban_fl	ag). Segment estimated urban cover flag as
	derived from the Glob	al Urban Footprint (GUF) data product. GUF
	is a global mapping of	urban areas derived from the TerraSAR-X

and TanDEM-X satellites. The GUF maps at a resolution of ~12 m (0.4 arcseconds). Due to differences in resolution, the ATL08 GUF value is set based upon a 4x4 block of pixels about the 100 m segment latitude/longitude. If ANY of the pixels the GUF pixels are labeled as urban, the ATL08 GUF value is set to urban. The GUF urban flag is set as -1 = undetermined, 0 = not urban, 1 = urban. The GUF data are available from DLR https://www.dlr.de/eoc/en/desktopdefault.aspx/tabid-**9628/16557 read-40454/**Surface Type

(parameter = surf\_type). The surface type for a given segment is determined at the major frame rate (every 200 shots, or  $\sim$ 140 meters along-track) and is a two-dimensional array surf\_type(n, nsurf), where n is the major frame number, and nsurf is the number of possible surface types such that surf\_type(n,isurf) is set to 0 or 1 indicating if surface type isurf is present (1) or not (0), where isurf = 1 to 5 (land, ocean, sea ice, land ice, and inland water) respectively.

### **2.4.22** ATL08\_region

(parameter = atl08\_region). The ATL08 regions that encompass the ATL03 granule being processed through the ATL08 algorithm. The ATL08 regions are shown by Figure 2.3.

#### 2.4.23 Last segment extend

(parameter = last\_seg\_extend). The distance (km) that the last ATL08 10 km processing segment is either extended beyond 10 km or uses data from the previous 10 km processing segment to allow for enough data for processing the ATL03 photons through the ATL08 algorithm. If the last portion of an ATL03 granule being processed would result in a segment with less than 3.4 km (170 geosegments) worth of data, that last portion is added to the previous 10 km processing window to be processed together as one extended ATL08 processing segment. The resulting last\_seg\_extend value would be a positive value of distance beyond 10 km that the ATL08 processing segment was extended by. If the last ATL08 processing segment would be less than

10 km but greater than 3.4 km, a portion extending from the start of current ATL08 processing segment backwards into the previous ATL08 processing segment would be added to the current ATL08 processing segment to make it 10 km in length. The distance of this backward data gathering would be reported in last\_seg\_extend as a negative distance value. Only new 100 m ATL08 segment products generated from this backward extension would be reported. All other segments that are not extended will report a last\_seg\_extend value of 0.

### **2.4.24** Brightness\_flag

(parameter = brightness\_flag). Based upon the classification of the photons within each 100 m, this parameter flags ATL08 segments where the mean number of ground photons per shot exceed a value of 3. This calculation can be made as the total number of ground photons divided by the number of ATLAS shots within the 100 m segment. A value of 0 = indicates non-bright surface, value of 1 indicates bright surface, and a value of 2 indicates "undetermined" due to clouds or other factors. The brightness is computed initially on the 10 km processing segment. If the ground surface is determined to be bright for the entire 10 km segment, the brightness is then calculated at the 100 m segment size.

### 2.5 Subgroup: Beam data

The subgroup for beam data contains basic information on the geometry and pointing accuracy for each beam.

Table 2.5. Summary table for beam parameters for the ATL08 product.

Group	Data Type	Units	Description	Source
segment_id_beg	Integer		First along-track segment_id number in 100-m segment	ATL03
segment_id_end	Integer		Last along-track segment_id number in 100-m segment	ATL03

ref_elev	Float	Elevation of the unit	ATL03
		pointing vector for the	
		reference photon in the	
		local ENU frame in	
		radians. The angle is	
		measured from East-	
		North plane and positive	
		towards up	
ref_azimith	Float	Azimuth of the unit	ATL03
		pointing vector for the	
		reference photon in the	
		ENU frame in radians.	
		The angle is measured	
		from North and positive	
		toward East.	
atlas_pa	Float	Off nadir pointing angle	ATL03
		of the spacecraft	
rgt	Integer	The reference ground	ATL03
S	O	track (RGT) is the track	
		on the earth at which	
		the vector bisecting	
		laser beams 3 and 4 is	
		pointed during repeat	
		operations	
sigma_h	Float	Total vertical	ATL03
51 <b>gu</b> 1	11000	uncertainty due to PPD	111200
		and POD	
sigma_along	Float	Total along-track	ATL03
Sigma_urong	11000	uncertainty due to PPD	1111100
		and POD knowledge	
sigma_across	Float	Total cross-track	ATL03
signia_act 033	Tioat	uncertainty due to PPD	MILOS
		and POD knowledge	
sigma_topo	Float	Uncertainty of the	computed
Signia_topo	rivat	geolocation knowledge	computeu
		due to local topography	
		(Equation 1.3)	
sigma_atlas_land	Float	Total uncertainty that	computed
sigilia_atlas_iallu	rivat	includes sigma_h plus	computeu
		• ·	
		the geolocation	
		uncertainty due to local	
	in to an	slope Equation 1.2	
psf_flag	integer	Flag indicating	computed
		sigma_atlas_land (aka	
		PSF) as computed in	
		Equation 1.2 exceeds a	
		value of 1m.	

layer_flag	Integer	Cloud flag indicating presence of clouds or blowing snow	ATL09
cloud_flag_atm	Integer	Cloud confidence flag from ATL09 indicating clear skies	ATL09
msw_flag	Integer	Multiple scattering warning product produced on ATL09	ATL09
cloud_fold_flag	integer	Cloud flag to indicate potential of high clouds that have "folded" into the lower range bins	ATL09
asr	Float	Apparent surface reflectance	ATL09
snr	Float	Background signal to noise level	Computed
solar_azimuth  solar_elevation	Float	The azimuth (in degrees) of the sun position vector from the reference photon bounce point position in the local ENU frame. The angle is measured from North and is positive towards East.  The elevation of the sun position vector from the	ATL03g
		reference photon bounce point position in the local ENU frame. The angle is measured from the East-North plane and is positive Up.	
n_seg_ph	Integer	Number of photons within each land segment	computed
ph_ndx_beg	Integer	Photon index begin	computed
sat_flag	Integer	Flag derived from full_sat_fract and near_sat_fract on the ATL03 data product	computed

### **2.5.1** Georeferenced\_segment\_number\_beg

(parameter = segment\_id\_beg). The first along-track segment\_id in each 100-m segment. Each 100-m segment consists of five sequential 20-m segments provided from the ATL03 product, which are labeled as segment\_id. The segment\_id is a seven digit number that uniquely identifies each along track segment, and is written at the along-track geolocation segment rate (i.e. ~20m along track). The four digit RGT number can be combined with the seven digit segment\_id number to uniquely define any along-track segment number. Values are sequential, with 0000001 referring to the first segment after the equatorial crossing of the ascending node.

#### **2.5.2** Georeferenced\_segment\_number\_end

(parameter = segment\_id\_end). The last along-track segment\_id in each 100-m segment. Each 100-m segment consists of five sequential 20-m segments provided from the ATL03 product, which are labeled as segment\_id. The segment\_id is a seven digit number that uniquely identifies each along track segment, and is written at the along-track geolocation segment rate (i.e. ~20m along track). The four digit RGT number can be combined with the seven digit segment\_id number to uniquely define any along-track segment number. Values are sequential, with 0000001 referring to the first segment after the equatorial crossing of the ascending node.

#### **2.5.3** Beam coelevation

(parameter = ref\_elev). Elevation of the unit pointing vector for the reference photon in the local ENU frame in radians. The angle is measured from East-North plane and positive towards up.

#### **2.5.4** Beam\_azimuth

(parameter = ref\_azimuth). Azimuth of the unit pointing vector for the reference photon in the ENU frame in radians. The angle is measured from North and positive toward East.

# **2.5.5** ATLAS\_Pointing\_Angle

(parameter = atlas\_pa). Off nadir pointing angle (in radians) of the satellite to increase spatial sampling in the non-polar regions.

#### **2.5.6** Reference ground track

(parameter = rgt). The reference ground track (RGT) is the track on the earth at which the vector bisecting laser beams 3 and 4 (or GT2L and GT2R) is pointed during repeat operations. Each RGT spans the part of an orbit between two ascending equator crossings and are numbered sequentially. The ICESat-2 mission has 1387 RGTs, numbered from 0001xx to 1387xx. The last two digits refer to the cycle number.

### **2.5.7** Sigma\_h

(parameter = sigma\_h). Total vertical uncertainty due to PPD (Precise Pointing Determination), POD (Precise Orbit Determination), and geolocation errors. Specifically, this parameter includes radial orbit error,  $\sigma_{Orbit}$ , tropospheric errors,  $\sigma_{Trop}$ , forward scattering errors,  $\sigma_{forwardscattering}$ , instrument timing errors,  $\sigma_{timing}$ , and off-nadir pointing geolocation errors. The component parameters are pulled from ATL03 and ATL09. Sigma\_h is the root sum of squares of these terms as detailed in Equation 1.1. The sigma\_h reported here is the mean of the sigma\_h values reported within the five ATL03 geosegments that are used to create the 100 m ATL08 segment.

### **2.5.8** Sigma\_along

(parameter = sigma\_along). Total along-track uncertainty due to PPD and PODknowledge. This parameter is pulled from ATL03.

# **2.5.9** Sigma\_across

1212 (parameter = sigma\_across). Total cross-track uncertainty due to PPD and 1213 POD knowledge. This parameter is pulled from ATL03.

### **2.5.10** Sigma\_topo

(parameter = sigma\_topo). Uncertainty in the geolocation due to local surface slope as described in Equation 1.3. The local slope is multiplied by the 6.5 m geolocation uncertainty factor that will be used to determine the geolocation uncertainty. The geolocation error will be computed from a 100 m sample due to the local slope calculation at that scale.

#### **2.5.11** Sigma\_ATLAS\_LAND

(parameter = sigma\_atlas\_land). Total vertical geolocation error due to ranging, and local surface slope. The parameter is computed for ATL08 as described in Equation 1.2. The geolocation error will be computed from a 100 m sample due to the local slope calculation at that scale.

# **2.5.12** PSF\_flag

1226 (parameter = psf\_flag). Flag indicating that the point spread function 1227 (computed as sigma\_atlas\_land) has exceeded 1m.

# **2.5.13** Layer\_flag

(parameter = layer\_flag). Flag is a combination of multiple ATL09 flags and takes daytime/nighttime into consideration. A value of 1 means clouds or blowing snow is likely present. A value of 0 indicates the likely absence of clouds or blowing snow. If no ATL09 product is available for an ATL08 segment, an invalid value will be reported. Since the cloud flags from the ATL09 product are reported at an along-track distance of 250 m, we will report the highest value of the ATL09 flags at the ATL08 resolution (100 m). Thus, if a 100 m ATL08 segment straddles two values from ATL09, the highest cloud flag value will be reported on ATL08. This reporting strategy holds for all the cloud flags reported on ATL08.

1238 2.5.14 Cloud\_flag\_atm (parameter = cloud\_flag\_atm). Cloud confidence flag from ATL09 that indicates 1239 1240 the number of cloud or aerosol layers identified in each 25Hz atmospheric profile. If 1241 the flag is greater than 0, aerosols or clouds could be present. 1242 **2.5.15** MSW 1243 (parameter = msw\_flag). Multiple scattering warning flag with values from -1 to 1244 5 as computed in the ATL09 atmospheric processing and delivered on the ATL09 data 1245 product. If no ATL09 product is available for an ATL08 segment, an invalid value will 1246 be reported. MSW flags: 1247 -1 = signal to noise ratio too low to determine presence of 1248 cloud or blowing snow 1249 0 = no\_scattering 1250 1 = clouds at > 3 km1251 2 =clouds at 1-3km 1252 3 = clouds at < 1 km1253 4 = blowing snow at < 0.5 optical depth 1254 5 = blowing snow at >= 0.5 optical depth1255 2.5.16 Cloud Fold Flag 1256 (parameter = cloud\_fold\_flag). Clouds occurring higher than 14 to 15 km in the 1257 atmosphere will be folded down into the lower portion of the atmospheric profile. 1258 **2.5.17** Computed\_Apparent\_Surface\_Reflectance 1259 (parameter = asr). Apparent surface reflectance computed in the ATL09 1260 atmospheric processing and delivered on the ATL09 data product. If no ATL09 1261 product is available for an ATL08 segment, an invalid value will be reported.

1262 **2.5.18** Signal\_to\_Noise\_Ratio 1263 (parameter = snr). The Signal to Noise Ratio of geolocated photons as 1264 determined by the ratio of the superset of ATL03 signal and DRAGANN found signal 1265 photons used for processing the ATL08 segments to the background photons (i.e., 1266 noise) within the same ATL08 segments. 1267 **2.5.19** Solar\_Azimuth 1268 (parameter = solar\_azimuth). The azimuth (in degrees) of the sun position 1269 vector from the reference photon bounce point position in the local ENU frame. The angle is measured from North and is positive towards East. 1270 1271 **2.5.20** Solar Elevation 1272 (parameter = solar elevation). The elevation of the sun position vector from 1273 the reference photon bounce point position in the local ENU frame. The angle is 1274 measured from the East-North plane and is positive up. 1275 **2.5.21** Number of segment photons 1276 (parameter = n seg ph). Number of photons in each land segment. 1277 2.5.22 Photon\_Index\_Begin 1278 (parameter = ph ndx beg). Index (1-based) within the photon-rate data of 1279 the first photon within this each land segment. 1280 **2.5.23** Saturation Flag 1281 (parameter = sat\_flag) Saturation flag derived from the ATL03 saturation 1282 flags full\_sat\_frac. The saturation flags on the ATL03 data product (full\_sat\_fract) 1283 are the percentage of photons determined to be saturated within each geosement. 1284 For the ATL08 saturation flag, a value of 0 will indicate no saturation. A value of 1 1285 will indicate the average of all 5 geosement full\_sat\_fract values was over 0.2. This 1286 value of 1 is an indication of standing water or saturated soils. If an ATL08 segment 1287 is not fully populated with 5 values for full\_sat\_fract, a value of -1 will be set.

1288	sat_flag:	-1 indicates not enough valid data to make determination
1289		0 indicates no saturation in ATL08 segment
1290		1 indicates saturation in ATL08 segment
1291		
1292		
1293		
1294		

#### 3 ALGORITHM METHODOLOGY

For the ecosystem community, identification of the ground and canopy surface is by far the most critical task, as meeting the science objective of determining global canopy heights hinges upon the ability to detect both the canopy surface and the underlying topography. Since a space-based photon counting laser mapping system is a relatively new instrument technology for mapping the Earth's surface, the software to accurately identify and extract both the canopy surface and ground surface is described here. The methodology adopted for ATL08 establishes a framework to potentially accept multiple approaches for capturing both the upper and lower surface of signal photons. One method used is an iterative filtering of photons in the along-track direction. This method has been found to preserve the topography and capture canopy photons, while rejecting noise photons. An advantage of this methodology is that it is self-parameterizing, robust, and works in all ecosystems if sufficient photons from both the canopy and ground are available. For processing purposes, along-track data signal photons are parsed into *L*-km segment of the orbit which is recommended to be 10 km in length.

#### 3.1 Noise Filtering

Solar background noise is a significant challenge in the analysis of photon counting laser data. Range measurement data created from photon counting lidar detectors typically contain far higher noise levels than the more common photon integrating detectors available commercially in the presence of passive, solar background photons. Given the higher detection sensitivity for photon counting devices, a background photon has a greater probability of triggering a detection event over traditional integral measurements and may sometimes dominate the dataset. Solar background noise is a function of the surface reflectance, topography, solar elevation, and atmospheric conditions. Prior to running the surface finding algorithms used for ATL08 data products, the superset of output from the GSFC medium-high confidence classed photons (ATL03 signal\_conf\_ph: flags 3-4) and the

output from DRAGANN will be considered as the input data set. ATL03 input data requirements include the latitude, longitude, height, segment delta time, segment ID, and a preliminary signal classification for each photon. The motivation behind combining the results from two different noise filtering methods is to ensure that all of the potential signal photons for land surfaces will be provided as input to the surface finding software. The description of the methodology for the ATL03 classification is described separately in the ATL03 ATBD. The methodology behind DRAGANN is described in the following section.

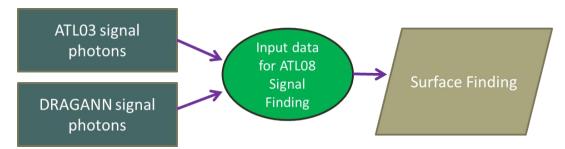


Figure 3.1. Combination of noise filtering algorithms to create a superset of input data for

#### 3.1.1 DRAGANN

surface finding algorithms.

The Differential, Regressive, and Gaussian Adaptive Nearest Neighbor (DRAGANN) filtering technique was developed to identify and remove noise photons from the photon counting data point cloud. DRAGANN utilizes the basic premise that signal photons will be closer in space than random noise photons. The first step of the filtering is to implement an adaptive nearest neighbor search. By using an adaptive method, different thresholds can be applied to account for variable amounts of background noise and changing surface reflectance along the data profile. This search finds an effective radius by computing the probability of finding P number of points within a search area. For MABEL and mATLAS, P=20 points within the search area

was empirically derived but found to be an effective and efficient number of neighbors.

There may be cases, however, where the value of P needs to be changed. For example, during night acquisitions it is anticipated that the background noise rate will be considerably low. Since DRAGANN is searching for two distributions in neighborhood searching space, the software could incorrectly identify signal photons as noise photons. The parameter P, however, can be determined dynamically from estimations of the signal and noise rates from the photon cloud. In cases of low background noise (night), P would likely be changed to a value lower than 20. Similarly, in cases of high amounts of solar background, P may need to be increased to better capture the signal and avoid classifying small, dense clusters of noise as signal. In this case, however, it is likely that noise photons near signal photons will also be misclassified as signal. The method for dynamically determining a P value is explained further in section 4.3.1.

After P is defined, a histogram of the number of neighbors within a search radius for each point is generated. The distribution of neighbor radius occurrences is analyzed to determine the noise threshold.

$$\frac{P}{N_{total}} = \frac{V}{V_{total}}$$
 Eqn. 3.1

where  $N_{total}$  is the total number of photons in the point cloud, V is the volume of the nearest neighborhood search, and  $V_{total}$  is the bounding volume of the enclosed point cloud. For a 2-dimensional data set, V becomes

1370 
$$V = \pi r^2$$
 Eqn. 3.2

where r is the radius. A good practice is to first normalize the data set along each dimension before running the DRAGANN filter. Normalization prevents the algorithm from favoring one dimension over the others in the radius search (e.g., when the latitude and longitude are in degrees and height is in meters).

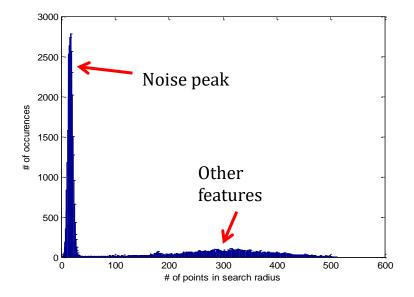


Figure 3.2. Histogram of the number of photons within a search radius. This histogram is used to determine the threshold for the DRAGANN approach.

Once the radius has been computed, DRAGANN counts the number of points within the radius for each point and histograms that set of values. The distribution of the number of points, Figure 3.2, reveals two distinct peaks; a noise peak and a signal peak. The motivation of DRAGANN is to isolate the signal photons by determining a threshold based on the number of photons within the search radius. The noise peak is characterized as having a large number of occurrences of photons with just a few neighboring photons within the search radius. The signal photons comprise the broad second peak. The first step in determining the threshold between the noise and signal is to implement Gaussian fitting to the number of photons distribution (i.e., the distribution shown in Figure 3.2). The Gaussian function has the form

1392 
$$g(x) = ae^{\frac{-(x-b)^2}{2c^2}}$$
 Eqn. 3.3

where a is the amplitude of the peak, b is the center of the peak, and c is the standard deviation of the curve. A first derivative sign crossing method is one option to identify peaks within the distribution.

To determine the noise and signal Gaussians, up to ten Gaussian curves are fit to the histogram using an iterative process of fitting and subtracting the maxamplitude peak component from the histogram until all peaks have been extracted. Then, the potential Gaussians pass through a rejection process to eliminate those with poor statistical fits or other apparent errors (Goshtasby and O'Neill, 1994; Chauve et al. 2008). A Gaussian with an amplitude less than 1/5 of the previous Gaussian and within two standard deviations of the previous Gaussian should be rejected. Once the errant Gaussians are rejected, the final two remaining are assumed to represent the noise and signal. These are separated based on the remaining two Gaussian components within the histogram using the logic that the leftmost Gaussian is noise (low neighbor counts) and the other is signal (high neighbor counts).

The intersection of these two Gaussians (noise and signal) determines a data threshold value. The threshold value is the parameter used to distinguish between noise points and signal points when the point cloud is re-evaluated for surface finding. In the event that only one curve passes the rejection process, the threshold is set at  $1 \sigma$  above the center of the noise peak.

An example of the noise filtered product from DRAGANN is shown in Figure 3.3. The signal photons identified in this process will be combined with the coarse signal finding output available on the ATL03 data product.

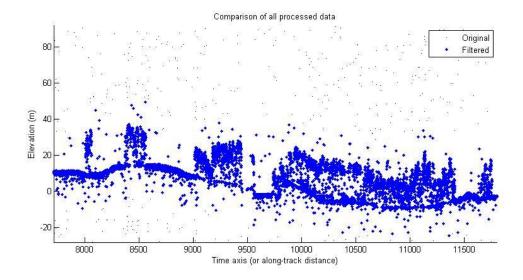


Figure 3.3. Output from DRAGANN filtering. Signal photons are shown as blue.

Figure 3.3 provides an example of along-track (profiling) height data collected in September 2012 from the MABEL (ICESat-2 simulator) over vegetation in North Carolina. The photons have been filtered such that the signal photons returned from vegetation and the ground surface are remaining. Noise photons that are adjacent to the signal photons are also retained in the input dataset; however, these should be classified as noise photons during the surface finding process. It is possible that some additional outlying noise may be retained during the DRAGANN process when noise photons are densely grouped, and these photons should be filtered out before the surface finding process. Estimates of the ground surface and canopy height can then be derived from the signal photons.

#### 3.2 Surface Finding

Once the signal photons have been determined, the objective is to find the ground and canopy photons from within the point cloud. With the expectation that one algorithm may not work everywhere for all biomes, we are employing a framework that will allow us to combine the solutions of multiple algorithms into one final composite solution for the ground surface. The composite ground surface solution will then be utilized to classify the individual photons as ground, canopy, top

of canopy, or noise. Currently, the framework described here utilizes one algorithm for finding the ground surface and canopy surface. Additional methods, however, could be integrated into the framework at a later time. Figure 3.4 below describes the framework.

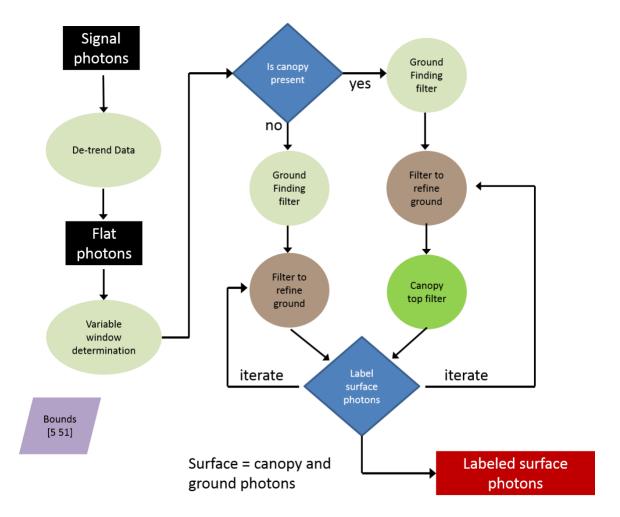


Figure 3.4. Flowchart of overall surface finding method.

#### 3.2.1 De-trending the Signal Photons

An important step in the success of the surface finding algorithm is to remove the effect of topography on the input data, thus improving the performance of the algorithm. This is done by de-trending the input signal photons by subtracting a heavily smoothed "surface" that is derived from the input data. Essentially, this is a low pass filter of the original data and most of the analysis to detect the canopy and ground will subsequently be implemented on the high pass data. The amount of smoothing that is implemented in order to derive this first surface is dependent upon the relief. For segments where the relief is high, the smoothing window size is decreased so topography isn't over-filtered.

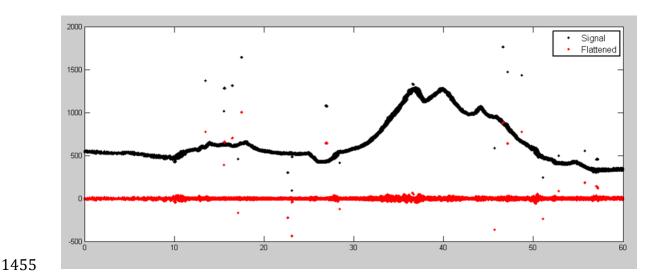


Figure 3.5. Plot of Signal Photons (black) from 2014 MABEL flight over Alaska and detrended photons (red).

### 3.2.2 Canopy Determination

A key factor in the success of the surface finding algorithm is for the software to automatically **account for the presence of canopy** along a given *L*-km segment. Due to the large volume of data, this process has to occur in an automated fashion, allowing the correct methodology for extracting the surface to be applied to the data. In the absence of canopy, the iterative filtering approach to finding ground works

extremely well, but if canopy does exist, we need to accommodate for that fact when we are trying to recover the ground surface.

For ATL08 product regions over Antarctica (regions 7, 8, 9, 10) and Greenland (region 11), the algorithm will assume only ground photons (canopy flag = 0) (see Figure 2.2).

#### 3.2.3 Variable Window Determination

The method for generating a best estimated terrain surface will vary depending upon whether canopy is present. *L-km segments* without canopy are much easier to analyze because the ground photons are usually continuous. *L-km* segments with canopy, however, require more scrutiny as the number of signal photons from ground are fewer due to occlusion by the vegetation.

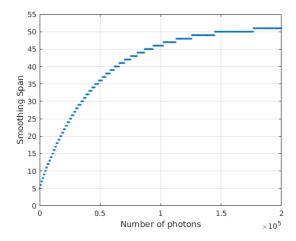
There are some common elements for finding the terrain surface for both cases (canopy/no canopy) and with both methods. In both cases, we will use a variable windowing span to compute statistics as well as filter and smooth the data. For clarification, the window size is variable for each *L-km* segment, but it is constant within the *L-km* segment. For the surface finding algorithm, we will employ a Savitzky-Golay smoothing/median filtering method. Using this filter, we compute a variable smoothing parameter (or window size). It is important to bound the filter appropriately as the output from the median filter can lose fidelity if the scan is over-filtered.

We have developed an empirically-determined shape function, bound between [5 51], that sets the window size (Sspan) based on the number of photons within each *L-km* segment.

1489 
$$Sspan = ceil[5 + 46 * (1 - e^{-a*length})]$$
 Eqn. 3.4

1490 
$$a = \frac{\log(1 - \frac{21}{51 - 5})}{-28114} \approx 21x10^{-6}$$
 Eqn. 3.5

where a is the shape parameter and length is the total number of photons in the L-km segment. The shape parameter, a, was determined using data collected by MABEL and is shown in Figure 3.6. It is possible that the model of the shape function, or the filtering bounds, will need to be adjusted once ICESat-2/ATLAS is on orbit and collecting data.



1497 Figure 3.6. Shape Parameter for variable window size.

#### 3.2.4 Compute descriptive statistics

To help characterize the input data and initialize some of the parameters used in the algorithm, we employ a moving window to compute descriptive statistics on the de-trended data. The moving window's width is the smoothing span function computed in Equation 5 and the window slides ¼ of its size to allow of overlap between windows. By moving the window with a large overlap helps to ensure that the approximate ground location is returned. The statistics computed for each window step include:

- Mean height
  - Min height
  - Max height
- Standard deviation of heights

Dependent upon the amount of vegetation within each window, the estimated ground height is estimated using different statistics. A standard deviation of the photon elevations computed within each moving window are used to classify the vertical spread of photons as belonging to one of four classes with increasing amounts of variation: open, canopy level 1, canopy level 2, canopy level 3. The canopy indices are defined in Table 3.1.

Table 3.1. Standard deviation ranges utilized to qualify the spread of photons within moving window.

Name	Definition	Lower Limit	Upper Limit
Open	Areas with little or no spread in signal photons determined due to low standard deviation	N/A	Photons falling within 1st quartile of Standard deviation
Canopy Level 1	Areas with small spread in signal photons	1 <sup>st</sup> quartile	Median
Canopy Level 2	Areas with a medium amount of spread	Median	3 <sup>rd</sup> quartile
Canopy Level 3	Areas with high amount of spread in signal photons	3 <sup>rd</sup> quartile	N/A

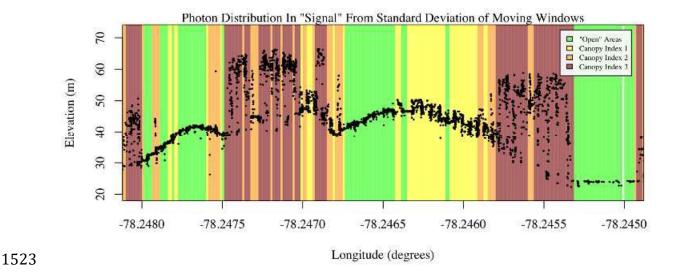
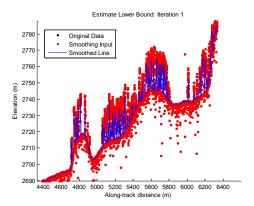
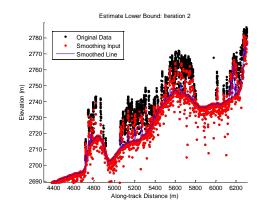


Figure 3.7. Illustration of the standard deviations calculated for each moving window to identify the amount of spread of signal photons within a given window.

#### 3.2.5 Ground Finding Filter (Iterative median filtering)

A combination of an iterative median filtering and smoothing filter approach will be employed to derive the output solution of both the ground and canopy surfaces. The input to this process is the set of de-trended photons. Finding the ground in the presence of canopy often poses a challenge because often there are fewer ground photons underneath the canopy. The algorithm adopted here uses an iterative median filtering approach to retain/eliminate photons for ground finding in the presence of canopy. When canopy exists, a smoothed line will lay somewhere between the canopy top and the ground. This fact is used to iteratively label points above the smoothed line as canopy. The process is repeated five times to eliminate canopy points that fall above the estimated surface as well as noise points that fall below the ground surface. An example of iterative median filtering is shown in Figure 3.8. The final median filtered line is the preliminary surface estimate. A limitation of this approach, however, is in cases of dense vegetation and few photons reaching the ground surface. In these instances, the output of the median filter may lie within the canopy.





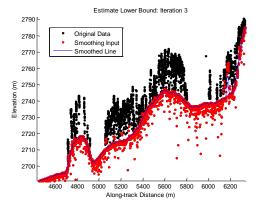


Figure 3.8. Three iterations of the ground finding concept for *L-km* segments with canopy.

# 3.3 Top of Canopy Finding Filter

Finding the top of the canopy surface uses the same methodology as finding the ground surface, except now the de-trended data are "flipped" over. The "flip" occurs by multiplying the photons heights by -1 and adding the mean of all the heights back to the data. The same procedure used to find the ground surface can be used to find the indices of the top of canopy points.

#### 3.4 Classifying the Photons

Once a composite ground surface is determined, photons falling within the point spread function of the surface are labeled as ground photons. Based on the expected performance of ATLAS, the point spread function should be approximately 35 cm rms. Signal photons that are not labeled as ground and are below the ground surface (buffered with the point spread function) are considered noise, but keep the signal label.

The top of canopy photons that are identified can be used to generate an upper canopy surface through a shape-preserving surface fitting method. All signal photons that are not labeled ground and lie above the ground surface (buffered with the point spread function) and below the upper canopy surface are considered to be canopy photons (and thus labeled accordingly). Signal photons that lie above the top of canopy surface are considered noise, but keep the signal label.

1570 FLAGS, 0 = noise 1571 1 = ground 1572 2 = canopy 1573 3 = TOC (top of canopy)

The final ground and canopy classifications are flags 1 - 3. The full canopy is the combination of flags 2 and 3.

#### 3.5 Refining the Photon Labels

During the first iteration of the algorithm, it is possible that some photons are mislabeled; most likely this would be noise photons mislabeled as canopy. To reject these mislabeled photons, we apply three criteria:

- a) If top of canopy photons are 2 standard deviations above a smoothed median top of canopy surface
- b) If there are less than 3 canopy indices within a 15m radius

There are also instances where the ground points will be redefined. This reassigning of ground points is based on how the final ground surface is determined. Following the "iterate" steps in the flowchart shown in Figure 3.4, if there are no canopy indices identified for the *L-km* segment, the final ground surface is interpolated from the identified ground photons and then will undergo a final round of median filtering and smoothing.

If canopy photons are identified, the final ground surface is interpolated based upon the level/amount of canopy at that location along the segment. The final ground surface is a composite of various intermediate ground surfaces, defined thusly:

ASmooth heavily smoothed surface used to de-trend the signal data

Interp\_Aground interpolated ground surface based upon the identified ground photons

AgroundSmooth median filtered and smoothed version of Interp\_Aground

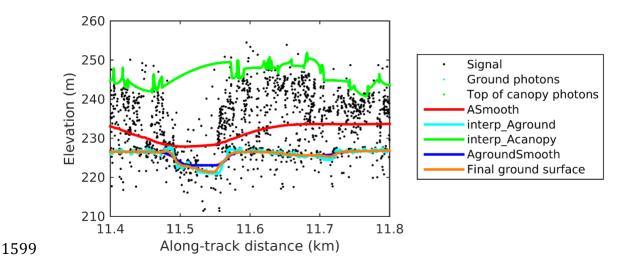


Figure 3.9. Example of the intermediate ground and top of canopy surfaces calculated from MABEL flight data over Alaska during July 2014.

During the first round of ground surface refinement, where there are canopy photons identified in the segment, the ground surface at that location is defined by the smoothed ground surface (AgroundSmooth) value. Else, if there is a location along-track where the standard deviation of the ground-only photons is greater than the 75% quartile for all signal photon standard deviations (i.e., canopy level 3), then the ground surface at that location is a weighted average between the interpolated ground surface (Interp\_Aground\*1/3) and the smoothed interpolated ground surface (AgroundSmooth\*2/3). For all remaining locations long the segment, the ground surface is the average of the interpolated ground surface (Interp\_Aground) and the heavily smoothed surface (Asmooth).

The second round of ground surface refinement is simpler than the first. Where there are canopy photons identified in the segment, the ground surface at that location is defined by the smoothed ground surface (AgroundSmooth) value again. For all other locations, the ground surface is defined by the interpolated ground surface (Interp\_Aground). This composite ground surface is run through the median and smoothing filters again.

The pseudocode for this surface refining process can be found in section 4.10.

Examples of the ground and canopy photons for several MABEL lines are shown in Figures 3.10 - 3.12.

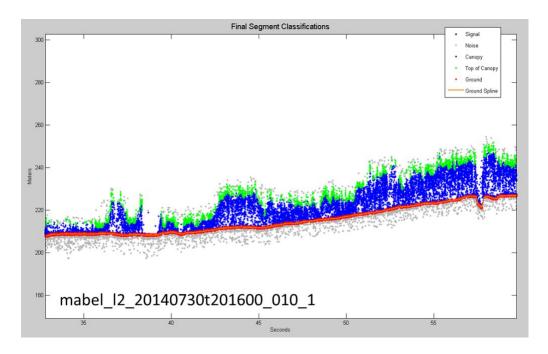


Figure 3.10. Example of classified photons from MABEL data collected in Alaska 2014. Red photons are photons classified as terrain. Green photons are classified as top of canopy. Canopy photons (shown as blue) are considered as photons lying between the terrain surface and top of canopy.

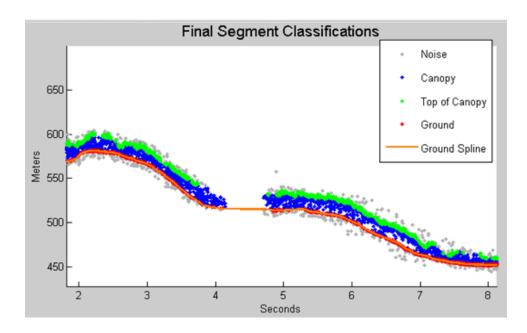


Figure 3.11. Example of classified photons from MABEL data collected in Alaska 2014. Red photons are photons classified as terrain. Green photons are classified as top of canopy. Canopy photons (shown as blue) are considered as photons lying between the terrain surface and top of canopy.

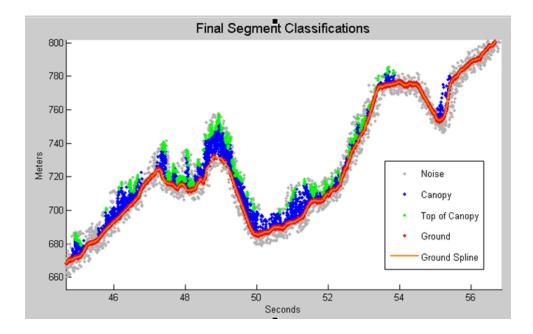


Figure 3.12. Example of classified photons from MABEL data collected in Alaska 2014. Red photons are photons classified as terrain. Green photons are classified as top of canopy.

Canopy photons (shown as blue) are considered as photons lying between the terrain surface and top of canopy.

#### 3.6 Canopy Height Determination

Once a final ground surface is determined, canopy heights for individual photons are computed by removing the ground surface height for that photon's latitude/longitude. These relative canopy height values will be used to compute the canopy statistics on the ATL08 data product.

#### 3.7 Link Scale for Data products

The link scale for each segment within which values for vegetation parameters will be derived will be defined over a fixed distance of 100 m. A fixed segment length ensures that canopy and terrain metrics are consistent between segments, in addition to increased ease of use of the final products. A size of 100 m was selected as it should provide approximately 140 photons (a statistically sufficient number) from which to make the calculations for terrain and canopy height.

#### 4. ALGORITHM IMPLEMENTATION

Prior to running the surface finding algorithms used for ATL08 data products, the superset of output from the GSFC medium-high confidence classed photons (ATL03 signal\_conf\_ph: flags 3-4) and the output from DRAGANN will be considered as the input data set. ATL03 input data requirements include the along-track time, latitude, longitude, height, and classification for each photon. The motivation behind combining the results from two different noise filtering methods is to ensure that all of the potential signal photons for land surfaces will be provided as input to the surface finding software.

Some additional quality checks are also described here prior to implementing the ATL08 software. The first check utilizes the POD\_PPD flag on ATL03. In instances where the satellite is maneuvering or the pointing/ranging solutions are suspect, ATL08 will not use those data. Thus, data will only flow to the ATL08 algorithm when the POD\_PPD flag is set to 0 which indicates 'nominal' conditions.

A second quality check pertains to the flags set on the ATL03 photon quality flag (quality\_ph). Currently, ATL03 quality\_ph flags are described as:

0 = nominal conditions

2 = possible late impulse response effect (this flag identifies additional detector1672 effects 5 - 50 m below the surface).

1673 3 = possible TEP crossing.

For this release of the software, we want to mention that there are cases of after-pulsing that occur 0.5 - 2.3 m below the surface that are considered nominal with the quality\_ph flag. The output from the DRAGANN algorithm (i.e. the DRAGANN flag) will be set to a value of 0 when ATL03 quality\_ph flags are greater than 0 such that they are ignored in the ATL08 algorithm.

A third quality check pertains to the signal photons (DRAGANN + ATL03 signal confidence photons) and whether those heights are near the surface heights. To pass this check, signal photons that lie 120 m above the reference DEM will be disregarded. Signal photons lying below the reference DEM will be allowed to continue for additional ATL08 processing. The motivation for this quality check is to eliminate ICESat-2 photons that are reflecting from clouds rather than the true surface.

Table 4.1. Input parameters to ATL08 classification algorithm.

Name	Data Type	Long Name	Units	Description	Source
delta_time	DOUBLE	GPS elapsed time	seconds	Elapsed GPS seconds since start of the granule for a given photon. Use the metadata attribute granule_start_seconds to compute full gps time.	ATL03
lat_ph	FLOAT	latitude of photon	degrees	Latitude of each received photon. Computed from the ECEF Cartesian coordinates of the bounce point.	ATL03
lon_ph	FLOAT	longitude of photon	degrees	Longitude of each received photon. Computed from the ECEF Cartesian coordinates of the bounce point.	ATL03
h_ph	FLOAT	height of photon	meters	Height of each received photon, relative to the WGS-84 ellipsoid.	ATL03
sigma_h	FLOAT	height uncertainty	m	Estimated height uncertainty (1-sigma) for the reference photon.	ATL03
signal_conf_ph	UINT_1_L E	photon signal confidence	counts	Confidence level associated with each photon event selected as signal (0-noise. 1- added to allow for buffer but algorithm classifies as background, 2-low, 3-med, 4-high).	ATL03
segment_id	UNIT_32	along-track segment ID number	unitless	A seven-digit number uniquely identifying each along-track segment. These are sequential, starting with one for the first segment after an ascending equatorial crossing node.	ATL03

cab_prof	FLOAT	Calibrated Attenuated Backscatte r	unitless	Calibrated Attenuated Backscatter from 20 to -1 km with vertical resolution of 30m	ATL09
dem_h	FLOAT	DEM Height	meters	Best available DEM (in priority of GIMP/ANTARCTIC/GMTED/M SS) value at the geolocation point. Height is in meters above the WGS84 Ellipsoid.	ATL09

# Table 4.2. Additional external parameters referenced in ATL08 product.

Name	Data Type	Long Name	Units	Description	Source
atlas_pa				Off nadir pointing angle of the spacecraft	
ground_track				Ground track, as numbered from left to right: 1 = 1L, 2 = 1R, 3 = 2L, 4 = 2R, 5 = 3L, 6 = 3R	
dem_h				Reference DEM height	ANC06
ref_azimuth	FLOAT	azimuth	radians	Azimuth of the unit pointing vector for the reference photon in the local ENU frame in radians. The angle is measured from north and positive towards east.	ATL03
ref_elev	FLOAT	elevation	radians	Elevation of the unit pointing vector for the reference photon in the local ENU frame in radians. The angle is measured from eastnorth plane and positive towards up.	ATL03
rgt	INTEGER_2	reference ground track	unitless	The reference ground track (RGT) is the track on the Earth at which a specified unit vector within the observatory is pointed. Under nominal operating conditions, there will be no data collected along the RGT, as the RGT is spanned by GT2L and GT2R. During slews or off-pointing, it is possible that ground tracks	ATL03

				may intersect the RGT. The ICESat-2 mission has 1,387 RGTs.	
sigma_along	DOUBLE	along-track geolocation uncertainty	meters	Estimated Cartesian along-track uncertainty (1-sigma) for the reference photon.	ATL03
sigma_across	DOUBLE	across-track geolocation uncertainty	meters	Estimated Cartesian across- track uncertainty (1-sigma) for the reference photon.	ATL03
surf_type	INTEGER_ 1	surface type	unitless	Flags describing which surface types this interval is associated with. 0=not type, 1=is type. Order of array is land, ocean, sea ice, land ice, inland water.	ATL03 , Section 4
layer_flag	Integer	Consolidated cloud flag	unitless	Flag indicating the presence of clouds or blowing snow with good confidence	ATL09
cloud_flag_asr	Integer(3)	Cloud probability from ASR	unitless	Cloud confidence flag, from 0 to 5, indicating low, med, or high confidence of clear or cloudy sky	ATL09
msw_flag	Byte(3)	Multiple scattering warning flag	unitless	Flag with values from 0 to 5 indicating presence of multiple scattering, which may be due to blowing snow or cloud/aerosol layers.	ATL09
asr	Float(3)	Apparent surface reflectance	unitless	Surface reflectance as modified by atmospheric transmission	ATL09
snow_ice	INTEGER_ 1	Snow Ice Flag	unitless	NOAA snow-ice flag. 0=ice free water; 1=snow free land; 2=snow; 3=ice	ATL09

## 4.1 Cloud based filtering

It is possible for the presence of clouds to affect the number of surface photon returns through signal attenuation, or to cause false positive classifications of ground or canopy photons on low cloud returns. Either of these cases would reduce the accuracy of the ATL08 product. To improve the performance of the ATL08 algorithm, ideally all clouds would be identified prior to processing through the ATL08 algorithm. There will be instances, however, where low lying clouds (e.g.

1697	<800 m above the ground surface) may be difficult to identify. Currently, ATL08
1698	provides an ATL09 derived cloud flag (layer_flag) on its 100 m product and
1699	encourages the user to make note of the presence of clouds when using ATL08
1700	output. Unfortunately at present, a review of on-orbit data from ATL03 and ATL09
1701	indicate that the cloud layer flag is not being set correctly in the ATL09 algorithm.
1702	Ultimately, the final cloud based filtering process used in the ATL08 algorithm will
1703	most likely be derived from parameters/flag on the ATL09 data product. Until the
1704	ATL09 cloud flags are proven reliable, however, a preliminary cloud screening
1705	method is presented below. This methodology utilizes the calibrated attenuated
1706	backscatter on the ATL09 data product to identify (and subsequently remove for
1707	processing) clouds or other problematic issues (i.e. incorrectly telemetered
1708	windows). Using this new method, telemetered windows identified as having either
1709	low or no surface signal due to the presence of clouds (likely above the telemetered
1710	band), as well as photon returns suspected to be clouds instead of surface returns,
1711	will be omitted from the ATL08 processing. This process, however, will not identify
1712	the extremely low clouds (i.e. <800 m). The steps are as follows:

- 1713 1. Match up the ATL09 calibrated attenuated backscatter (cab\_prof) columns to the ATL03 granule being processed using segment ID.
- 17152. Flip the matching cab\_prof vertical columns so that the elevation bins go1716from low to high.
- 3. For each of the matching ATL09 cab\_prof vertical columns, perform a cubic Savitsky-Golay smoothing filter with a span size of 15 vertical bins. Call this cab\_smooth.
- 4. Perform the same smoothing filter on each horizontal row of the cab\_smooth output, this time using a span size of 7 horizontal bins. Call this cab\_smoother.
- 5. Create a low\_signal logical array the length of the number of matching ATL09columns and set to false.
- 1725 6. For each column of cab\_smoother:

a. Set any values below 0 to 0.

1727 b. Set a logical array of cab smoother bins that are below 15 km in 1728 elevation to true. Call this cab15. 1729 c. Using the ATL09 dem h value for that column, find the ATL09 1730 cab\_smoother bins that are 240 m above and 240 m below (~8 ATL09 1731 vertical bins each direction) the dem h value. The bins found here that 1732 are also within cab15 are designated as sfc\_bins. 1733 d. Find the maximum peak value of cab smoother within the sfc bins, if 1734 any. This will represent the surface peak. 1735 e. Find the maximum value of cab smoother that is higher in elevation 1736 than the sfc bins and within cab15, if any. This will represent the 1737 cloud peak. 1738 f. If there is no surface peak, set the low signal flag to true. 1739 g. If there are both surface and cloud peak values returned, determine a 1740 surface peak / cloud peak ratio. If that ratio is less than or equal to 0.4, 1741 set low signal flag for that column to true. 1742 7. After each matching ATL09 column of cab\_smoother has been analyzed for 1743 low signal, assign the low signal flag to an ATL03 photon resolution logical 1744 array by matching up the ATL03 photon segment\_id values to the ATL09 1745 range of segment IDs for each ATL09 cab\_prof column. 1746 8. For each ATL09 cab prof column where the low signal flag was not set, check 1747 for any ATL03 photons greater than 800 meters (TBD) in elevation away 1748 (higher or lower) from the ATL09 dem\_h value. Assign an ATL03 photon 1749 resolution too\_far\_signal flag to true when this conditional is met. 1750 9. A logical array mask is created for any ATL03 photons that have either the 1751 low signal flag or the too far signal flag set to true such that those photons 1752 will not be further processed by the ATL08 function. 1753 1754 4.2 Preparing ATL03 data for input to ATL08 algorithm 1755 1. At times, cloud attenuation will lead to a reduced L-km with a length that is 1756 not a multiple of 100 meters. If the last 100m land segment of the L-km

segment contains fewer than 5 ATL03 20m geosegments and the current L-1758 km segment is not the last one of the granule, do not report output for this 1759 last 100m land segment. Retain the starting geosegment of this land segment 1760 and begin the next L-km segment here. 1761 2. Break up data into *L-km* segments. Segments equivalent of 10 km in along-1762 track distance of an orbit would be appropriate. 1763 a. If the last portion of an ATL03 granule being processed would result 1764 in an *L-km* segment with less than 3.4 km (170 geosegments) worth of 1765 data, that last portion is added to the previous *L-km* processing 1766 window to be processed together as one extended *L-km* processing 1767 segment. 1768 i. The resulting **last\_seg\_extend** value would be reported as a 1769 positive value of distance beyond 10 km that the ATL08 1770 processing segment was extended by. 1771 b. If the last *L-km* segment would be less than 10 km but greater than 3.4 1772 km, a portion extending from the start of current *L-km* processing 1773 segment backwards into the previous *L-km* processing segment would 1774 be added to the current ATL08 processing segment to make it 10 km 1775 in length. Only new 100 m ATL08 segment products generated from 1776 this backward extension would be reported. 1777 i. The distance of this backward data gathering would be 1778 reported in **last\_seg\_extend** as a negative distance value. 1779 c. All other segments that are not extended will report a last\_seg\_extend 1780 value of 0. 1781 3. Add a buffer of 200 m (or 10 segment id's) to both ends of each *L-km* 1782 segment. The total processing segment length is (L-km + 2\*buffer), but will 1783 be referred to as *L-km* segments for simplicity. 1784 a. The first *L-km* segment from an ATL03 granule would only have a 1785 buffer at the end, and the last *L-km* segment from an ATL03 granule 1786 would only have a buffer at the beginning. 1787 4. The input data for ATL08 algorithm is X, Y, Z, T (where T is time).

1788	
1789	4.3 Noise filtering via DRAGANN
1790	DRAGANN will use ATL03 photons with all signal classification flags (0-4). These
1791	will include both signal and noise photons. This section give a broad overview of the
1792	DRAGANN function. See Appendix A for more details.
1793	1. Determine the relative along-track time, ATT, of each geolocated photon
1794	from the beginning of each $L$ - $km$ segment.
1795	2. Rescale the ATT with equal-time spacing between each data photon, keeping
1796	the relative beginning and end time values the same.
1797	3. Normalize the height and rescaled ATT data from $0-1$ for each $L$ - $km$
1798	segment based on the min/max of each field. So, normtime = (time -
1799	mintime)/(maxtime - mintime).
1800	4. Build a kd-tree based on normalized Z and normalized and rescaled ATT.
1801	5. Determine the search radius starting with Equation 3.1. P=[determined by
1802	preprocessor; see Sec 4.3.1], and $V_{\text{total}}$ =1. $N_{\text{total}}$ is the number of photons
1803	within the data $L$ - $km$ segment. Solve for $V$ .
1804	6. Now that you know V, determine the radius using Equation 3.2.
1805	7. Compute the number of neighbors for each photon using this search radius.
1806	8. Generate a histogram of the neighbor count distribution. As illustrated in
1807	Figure 3.2, the noise peak is the first peak (usually with the highest
1808	amplitude).
1809	9. Determine the 10 highest peaks of the histogram.
1810	10. Fit Gaussians to the 10 highest peaks. For each peak,
1811	a. Compute the amplitude, a, which is located at peak position b.
1812	b. Determine the width, c, by stepping one bin at a time away from b and
1813	finding the last histogram value that is $> \frac{1}{2}$ the amplitude, a.
1814	c. Use the amplitude and width to fit a Gaussian to the peak of the
1815	histogram, as described in Equation 3.3.
1816	d. Subtract the Gaussian from the histogram, and move on to calculate
1817	the next highest peak's Gaussian.

1818	e. Reject Gaussians that are too near (< 2 standard deviations) and
1819	amplitude too low (<1/5 previous amplitude) from the previous
1820	signal Gaussian.
1821	11. Reject any of the returned Gaussians with imaginary components.
1822	12. Determine if there is a narrow noise Gaussian at the beginning of the
1823	histogram. These typically occur when there is little noise, such as during
1824	nighttime passes.
1825	a. Search for the Gaussian with the highest amplitude, a, in the first $5\%$
1826	of the histogram
1827	b. Check if the highest amplitude is $ >= 1/10 $ of the maximum of all
1828	Gaussian amplitudes
1829	c. Check if the width, c, of the Gaussian with the highest amplitude is <=
1830	4 bins
1831	d. If these three conditions are met, save the [a,b,c] values as [a0,b0,c0].
1832	e. If the three conditions are not met, search again within the first $10\%$ .
1833	Repeat the process, incrementing the percentage of histogram
1834	searched by $5\%$ up to $30\%$ . As soon as the conditions are met, save
1835	the $[a_0,b_0,c_0]$ values and break out of the percentage histogram search
1836	loop.
1837	13. If a narrow noise peak was found, sort the remaining Gaussians from largest
1838	to smallest area, estimated by $a^*c$ , then append $[a_0,b_0,c_0]$ to the beginning of
1839	the sorted [a,b,c] arrays. If a narrow noise peak was not found, sort all
1840	Gaussians by largest to smallest area.
1841	a. If a narrow noise peak was not found, check in sorted order if one of
1842	the Gaussians are in the first $10\%$ of the histogram. If so, it becomes
1843	the first Gaussian.
1844	b. Reject any Gaussians that are fully contained within another.
1845	c. Reject Gaussians whose centers are within 3 standard deviations of
1846	another, unless only two Gaussians remain
1847	14. If there are two or more Gaussians remaining, they are referred to as
1848	Gaussian 1 and Gaussian 2, assumed to be the noise and signal Gaussians.

1849	15. Determine the threshold value that will define the cutoff between noise and
1850	signal.
1851	a. If the absolute difference of the two Gaussians becomes near zero,
1852	defined as < 1e-8, set the first bin index where that occurs, past the
1853	first Gaussian peak location, as the threshold. This would typically be
1854	set if the two Gaussians are far away from each other.
1855	b. Else, the threshold value is the intersection of the two Gaussians,
1856	which can be estimated as the first bin index past the first Gaussian
1857	peak location and before the second Gaussian where there is a
1858	minimum absolute difference between the two Gaussians.
1859	c. If there is only one Gaussian, it is assumed to be the noise Gaussian,
1860	and the threshold is set to b + c.
1861	16. Label all photons having a neighbor count above the threshold as signal.
1862	17. Label all photons having a neighbor count below the threshold as noise.
1863	18. Reject noise photons.
1864	19. Retain signal photons for feeding into next step of processing.
1865	20. Use Logical OR to combine DRAGANN signal photons with ATL03 medium-
1866	high confidence signal photons (flags 3-4) as ATL08 signal photons.
1867	21. Calculate a signal to noise ratio (SNR) for the $L$ - $km$ segment by dividing the
1868	number of ATL08 signal photons by the number of noise (i.e., all - signal)
1869	photons.
1870	4.3.1 DRAGANN Quality Assurance
1871	Based upon on-orbit data, there are instances where only noise photons are selected
1872	as signal photons following running through DRAGANN. These instances usually
1873	occur to telemetered windows with low signal, signal attenuation near the surface
1874	due to fog, haze (or other atmospheric properties). If any d_flag results in the $10\ km$
1875	= 1
1876	1. For each 20 m segment_id that has a d_flag = 1, build a histogram of 5 m
1877	height bins using the height of only the DRAGANN-flagged photons
1878	(d_flag=1)

1879 2. If the number of bins indicates that all d\_flag photons fall within the same 1880 vertical 60 m, do nothing and move to the next geosement. 1881 3. If the d flag photons fall outside of 60 m, calculate the median and 1882 standard deviation of the histogram counts. 1883 4. If the maximum value of the histogram counts is greater than the median 1884 + 3\*standard deviation, a surface peak has been detected based on the 1885 relative photon density within the 5 meter steps. Else, set all d flag = 0 1886 for this geosegment. 1887 5. Set all d\_flag = 0 from 3 height bins below the detected peak to the bottom 1888 of the telemetry window. 1889 6. Starting with the peak count bin (surface), step upwards bin by bin and 1890 check if 12 bin counts (60 meters of height bins) above surface are less 1891 than 0.5 \* histogram median. If so, for all photons above current height in 1892 loop + 60 meters, set all d flag = 0 and exit bin-by-bin loop. 1893 7. Starting with one bin above the peak count bin (surface), again step 1894 upwards bin by bin. For each iteration, calculate the standard deviation of the bin counts including only the current bin to the highest height bin and 1895 1896 call this noise standard deviation. If all remaining vertical height bins 1897 from current bin to highest height bin are less than 2\* histogram 1898 standard deviation, or if the noise standard deviation is less than 1.0, or if 1899 this bin and the next 2 higher bins each have counts less than the peak bin 1900 count (entire histogram) – 3\*histogram standard deviation, then set all 1901 d\_flag = 0 for all heights above this level and exit bin-by-bin loop 1902 8. For a final check, construct a new histogram, with median and standard 1903 deviation, using the corrected d flag results and only where d flag = 1. If 1904 the histogram median is greater than 0.0 and the standard deviation is 1905 greater than 0.75\* median, set all d\_flag in this geosegment = 0. This 1906 indicates results not well constrained about a detectible surface.

1908	4.3.2 Preprocessing to dynamically determine a DRAGANN parameter
1909	While a default value of P=20 was found to work well when testing with MABEL
1910	flight data, further testing with simulated data showed that P=20 is not sufficient in
1911	cases of very low or very high noise. Additional testing with real ATL03 data have
1912	shown the ground signal to be much stronger, and the canopy signal to be much
1913	weaker, than originally anticipated. Therefore, a preprocessing step for dynamically
1914	calculating P and running the core DRAGANN function is described in this
1915	subsection. This assumes $L$ - $km$ to be 10 km (with additional $L$ - $km$ buffering).
1916	1. Define a DRAGANN processing window of 170 segments (∼3.4 km),
1917	and a buffer of 10 segments ( $\sim$ 200 m).
1918	2. The buffer is applied to both sides of each DRAGANN processing
1919	window to create buffered DRAGANN processing windows
1920	(referenced as "buffered window" for the rest of this section) that will
1921	overlap the DRAGANN processing windows next to them.
1922	3. For each buffered window within the $L$ - $km$ segment, calculate a
1923	histogram of points with 1 m elevation bins.
1924	4. For each buffered window histogram, calculate the median counts.
1925	5. Bins with counts below the buffered window median count value are
1926	estimated to be noise. Calculate the mean count of noise bins.
1927	6. Bins with counts above the buffered window median count value are
1928	estimated to be signal. Calculate the mean count of signal bins.
1929	7. Determine the time elapsed over the buffered window.
1930	8. Calculate estimated noise and signal rates for each buffered window
1931	by multiplying each window's mean counts of noise bins and signal
1932	bins, determined from steps 5 and 6 above, by 1/(elapsed time) to
1933	return the rates in terms of points/meter[elevation]/second[across].
1934	9. Calculate a noise ratio for each window by dividing the noise rate by
1935	the signal rate.
1936	10. If, for all the buffered windows in the $L$ - $km$ segment, the noise rate is
1937	less than 20 and the noise ratio is less than 0.15; OR any noise rate is

1938	0; OR any signal rate is greater than 1000: re-calculate steps 3-9
1939	using the entire $L$ - $km$ segment. Continue with the following steps
1940	using results from the one $L$ - $km$ window (instead of multiple buffered
1941	windows).
1942	11. Now, determine the DRAGANN parameter, P, for each buffered
1943	window based on the following conditionals:
1944	a. If the signal rate is NaN (i.e., an invalid value), set the signal
1945	index array to empty and move on to the next buffered
1946	window.
1947	b. If noise rate < 20    noise ratio < 0.15:
1948	P = signal rate
1949	If signal rate is $< 5$ , $P = 5$ ; if signal rate $> 20$ , $P = 20$
1950	c. Else $P = 20$ .
1951	12. Run DRAGANN on the buffered window points using the calculated P.
1952	13. If DRAGANN fails to find a signal (i.e., only one Gaussian found), run
1953	DRAGANN again with $P = 10$ .
1954	14. If DRAGANN still fails to find a signal, try to determine P a second time
1955	using the following conditionals:
1956	a. If (noise rate >= 20)
1957	&& (signal rate > 100)
1958	&& (signal rate < 250),
1959	P = (signal rate)/2
1960	b. Else if signal rate >= 250,
1961	if noise rate >= 250,
1962	P = (noise rate)*1.1
1963	else,
1964	P = 250
1965	c. Else, P = mean(noise rate, signal rate)
1966	15. Run DRAGANN on the buffered window points using the newly
1967	calculated P.
1968	a. If still no signal points are found, set a dragannError flag.

1969	16. If signal points were found by DRAGANN, for each buffered window
1970	calculate a signal check by dividing the number of signal points found
1971	via DRAGANN by the number of total points in the buffered window.
1972	17. If dragannError has been set, or there are suspect signal statistics, the
1973	following snippet of pseudocode will check those conditionals and try
1974	to iteratively find a better P value to run DRAGANN with:
1975	
1976	try_count = 0
1977	
1978	While dragannError
1979	( (noise rate >= 30)
1980	&& (signal check > noise ratio)
1981	&& (noise ratio >= 0.15) )
1982	(signal check < 0.001):
1983	
1984	if $P < 3$ ,
1985	break
1986	else,
1987	P = P*0.75
1988	end
1989	
1990	if try_count < 2
1991	Clear out signal index results from previous DRAGANN run
1992	Re-run DRAGANN with new P value
1993	Recalculate the signal check
1994	end
1995	
1996	if no signal index results are returned
1997	P = P*0.75
1998	end
1999	
2000	try_count = try_count + 1
2001	
2002	end
2003	
2004	18. If no signal photons are found by DRAGANN because only one
2005	Gaussian was found, set the threshold as b+c (i.e., one standard
2006	deviation away from the Gaussian peak location) for a final DRAGANN
2007	run. Otherwise, set the signal index array to empty and move on to the
2008	next buffered window.

2009 19. Assign the signal values found from DRAGANN for each buffered 2010 window to the original DRAGANN processing window range of points. 2011 20. Combine signal points from each DRAGANN processing window back 2012 into one *L-km* array of signal points for further processing. 2013 2014 4.3.3 Iterative DRAGANN processing 2015 It is possible in processing segments with high noise rates that DRAGANN will 2016 incorrectly identify clusters of noise as signal. One way to reduce these false positive 2017 noise clusters is to run the alternative DRAGANN process (Sec 4.3.1) again with the 2018 input being the signal output photons from the first run through alternative 2019 DRAGANN. Note that this methodology is still being tested, so by default this option 2020 should not be set. 2021 1. If SNR < 1 (TBD) from alternative DRAGANN run, run alternative DRAGANN 2022 process again using the output signal photons from first DRAGANN run as the 2023 input to the second DRAGANN run. 2024 2. Recalculate SNR based on output of second DRAGANN run. 2025 2026 2027 4.4 Compute Filtering Window 2028 1. Next step is to run a surface filter with a variable window size (variable in 2029 that it will change from *L-km* segment to *L-km* segment). The window-size is 2030 denoted as Window. 2.  $Sspan = ceil[5 + 46 * (1 - e^{-a*length})]$ , where length is the number of 2031 2032 photons in the segment. 3.  $a = \frac{\log(1 - \frac{21}{51 - 5})}{-29114} \approx 21x10^{-6}$ , where *a* is the shape parameter for the window 2033 2034 span.

#### 2036 **4.5 De-trend Data**

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2050

2051

2052

2053

- The input data are the signal photons identified by DRAGANN and the ATL03
   classification (signal\_conf\_ph) values of 3-4.
- 2039 2. Generate a rough surface by connecting all unique (time) photons to each other. Let's call this surface interp\_A.
- 3. Run a median filter through interp\_A using the window size set by the software. Output = Asmooth.
- 4. Define a reference DEM limit (ref\_dem\_limit) as 120 m (TBD).
- 5. Remove any Asmooth values further than the ref\_dem\_limit threshold from the reference DEM, and interpolate the Asmooth surface based on the remaining Asmooth values. The interpolation method to use is the shape preserving piecewise cubic Hermite interpolating polynomial hereafter labeled as "pchip" (Fritsch & Carlson, 1980).
  - 6. Compute the approximate relief of the *L-km* segment using the 95<sup>th</sup> 5<sup>th</sup> percentile heights of the signal photons. We are going to filter Asmooth again and the smoothing is a function of the relief.
  - 7. Define the SmoothSize using the conditional statements below. The SmoothSize will be used to detrend the data as well as to create an interpolated ground surface later.
- 2055 SmoothSize = 2 \* Window
- If relief>=900, SmoothSize= round(SmoothSize/4)
- If relief>=400 && <=900, SmoothSize=round(SmoothSize/3)
- If relief>=200 && <=400, SmoothSize=round(SmoothSize/2)
- 8. Greatly smooth Asmooth by first running Asmooth 10 times through a median filter then a smoothing filter with a moving average method on the result. Both the median filter and the smoothing filter use a window size of SmoothSize.
- 9. Create a second smooth line (Asmooth2) that roughly follows the ground and Asmooth2 will be used only for detrending data during initial ground

2065	е	estimation. Asmooth2 is created by running five iterations of a median filter
2066	а	and smoothing using SmoothSize defined in 4.6.7. The threshold for removing
2067	p	photons is 1 m above each iteration.
2068		
2069	4.6 F	ilter outlier noise from signal
2070	1.	If there are any signal data that are 150 meters above Asmooth, remove them
2071		from the signal data set.
2072	2.	If the standard deviation of the detrended signal is greater than 10 meters,
2073		remove any signal value from the signal data set that is 2 times the standard
2074		deviation of the detrended signal below Asmooth or Asmooth2.
2075	3.	Calculate a new Asmooth surface by interpolating (pchip method) a surface
2076		from the remaining signal photons and median filtering using the Window
2077		size, then median filter and smooth (moving average method) 10 times again
2078		using the SmoothSize.
2079	4.	Calculate a new Asmooth2 surface by interpolating a surface from remaining
2080		signal photons and repeat <b>step 4.6.9.</b>
2081	5.	Detrend the signal photons by subtracting the signal height values from the
2082		Asmooth2 surface height values. Use the Asmooth2 detrended heights for the
2083		initial ground estimate surface finding.
2084	6.	Other calculations for canopy and ground finding will utilize detrended from
2085		the original Asmooth.
2086		
2087	4.7 F	inding the initial ground estimate
2088	1.	At this point, the initial signal photons have been noise filtered and de-
2089		trended and should have the following format: X, Y, detrended Z, T (T=time).
2090		From this, the input data into the ground finding will be the ATD (along track
2091		distance) metric (such as time) and the detrended Z height values.
2092	2.	Define a medianSpan as round(Window*2/3).

- 3. Calculate the background neighbor density of the subsurface photons using
  ALL available photons (the non-detrended data). This step is run on all
  photons including noise photons. Histogram the photons in 0.5 m vertical
  bins and a 60 m horizontal bin.
  - 4. To avoid including zero population bins in the histogram signal tracking process, identify the bin with the maximum bin count among bins 3 7 (starting at the lowest height) across each 60 m within the 10-km processing window.
    - 5. Calculate the mean of those maximum bin values to represent the noise count for the 10-km window.
- 2103 6. The following steps are run on the detrended signal photons.
- 7. Calculate the brightness of the surface for each 60 m to be histogrammed via the calculation in Section 2.4.21. If a bright surface is detected, skip steps 8 and 9
  - 8. Determine the lowest 0.5 m histogram height bin for each 60 m along track, in the detrended heights where:
    - a. The neighbor density is 10 x greater than the background density and
    - b. The neighbor density is greater than the histogram population median plus 1/3 of the population standard deviation.
  - 9. The photons with detrended heights above this bin are masked from consideration in the initial ground height estimate. Detrended signal photons implies that the d\_flag photons.
- 2115 10. Identifying the ground surface is an iterative process. Start by assuming that
  2116 all the input signal height photons are the ground. The first goal is the cut
  2117 out the lower height excess photons in order to find a lower bound for
  2118 potential ground photons. This process is done 5 times and an offset of 4
  2119 meters is subtracted from the resulting lower bound. The smoothing filter
  2120 uses a moving average again:
- 2121 for j=1:5

2098

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2114

2122 cutOff = median filter (ground, medianSpan)

```
2123
                         cutOff = smooth filter (cutOff, Window)
2124
                         ground = ground((cutOff - ground) > -1)
2125
                      end
2126
                      lowerbound = median filter (ground, medianSpan*3)
2127
                      middlebound = smooth filter (lowerbound, Window)
2128
                      lowerbound = smooth filter (lowerbound, Window) - 4
2129
                  end:
2130
            11. Create a linearly interpolated surface along the lower bound points and only
2131
               keep input photons above that line as potential ground points:
2132
                  top = input(input > interp(lowerbound))
2133
            12. The next goal is to cut out excess higher elevation photons in order to find an
2134
               upper bound to the ground photons. This process is done 3 times and an
2135
               offset of 1 meter is added to the resulting upper bound. The smoothing filter
2136
               uses a moving average:
2137
                  for j = 1:3
2138
                      cutOff = median filter (top, medianSpan)
2139
                      cutOff = smooth filter (cutOff, Window)
2140
                      top = top((cutOff - top) > -1)
2141
                  end
2142
                  upperbound = median filter (top, medianSpan)
2143
                  upperbound = smooth filter (upperbound, Window) + 1
2144
            13. Create a linearly interpolated surface along the upper bound points and
2145
               extract the points between the upper and lower bounds as potential ground
2146
               points:
2147
                  ground = input( (input > interp(lowerbound)) & ...
2148
                                   (input < interp(upperbound)))
2149
            14. Refine the extracted ground points to cut out more canopy, again using the
2150
               moving average smoothing:
```

2151	For $j = 1:2$
2152	<pre>cutOff = median filter (ground, medianSpan)</pre>
2153	<pre>cutOff = smooth filter (cutOff, Window)</pre>
2154	ground = ground( (cutOff – ground) > -1)
2155	end
2156	15. Run the ground output once more through a median filter using window side
2157	medianSpan and a smoothing filter using window size Window, but this time
2158	with the Savitzky-Golay method.
2159	16. Finally, linearly interpolate a surface from the ground points.
2160	17. The first estimate of canopy points are those indices of points that are
2161	between 2 and 150 meters above the estimated ground surface. Save these
2162	indices for the next section on finding the top of canopy.
2163	18. The output from the final iteration of ground points is temp_interpA – an
2164	interpolated ground estimate.
2165	19. Find ground indices that lie within 10 m below and 0.5 m above of
2166	temp_interpA .
2167	20. Apply the ground indices to the original heights (i.e., not the de-trended data)
2168	to label ground photons.
2169	21. Interpolate a ground surface using the pchip method based on the ground
2170	photons. Output is interp_Aground.
2171	
2172	4.8 Find the top of the canopy
2173	1. The input are the ATD metric (i.e., time), and the de-trended Z values indexed
2174	by the canopy indices extracted from step 4.7(17).
2175	2. Flip this data over so that we can find a canopy "surface" by multiplying the
2176	de-trended canopy heights by -1.0 and adding the mean(heights).
2177	3. Finding the top of canopy is also an iterative process. Follow the same steps
2178	described in $4.7(2) - 4.7(16)$ , but use the canopy indexed and flipped Z
2179	values in place of the ground input.

- 4. Final retained photons are considered top of canopy photons. Use the indices of these photons to define top of canopy photons in the original (not detrended) Z values.
  - 5. Build a kd-tree on canopy indices using elevation data detrended with Asmooth.
  - 6. If there are less than three canopy indices within a 100m radius, reassign these photons to noise photons. Initially, a value of 15 m was used for the search radius. In Release 004 of the algorithm, this value was increased to 100 m to include more top of canopy photons that were not captured in the initial canopy spline estimate.

# 4.9 Compute statistics on de-trended (Asmooth) data

- 1. The input data have been noise filtered and de-trended (Asmooth) and should have the following input format: X, Y, detrended Z, T.
- 2. The input data will contain signal photons as well as a few noise photons near the surface.
- 3. Compute statistics of heights in the along-track direction using a sliding window. Using the window size (window), compute height statistics for all photons that fall within each window. These include max height, median height, mean height, min height, and standard deviation of all photon heights. Additionally, in each window compute the median height and standard deviation of just the initially classified top of canopy photons, and the standard deviation of just the initially classified ground photon heights. Currently only the median top of canopy, and all STD variables are being utilized, but it's possible that other statistics may be incorporated as changes/improvements are made to the code.
  - 4. Slide the window ¼ of the window span and recompute statistics along the entire *L-km* segment. This results in one value for each statistic for each window.

2209	5.	Determine canopy index categories for each window based upon the total
2210		distribution of STD values for all signal photons along the <i>L-km</i> segment
2211		based on STD quartiles.
2212	6.	Open canopy have STD values falling within the $1^{\rm st}$ quartile.
2213	7.	Canopy Level 1 has STD values falling from $1^{\rm st}$ quartile to median STD value.
2214	8.	Canopy Level 2 has STD values falling from median STD value to $3^{\rm rd}$ quartile.
2215	9.	Canopy Level 3 has STD values falling from 3 <sup>rd</sup> quartile to max STD.
2216	10	Linearly interpolate the window STD values (both for all photons and
2217		ground-only photons) back to the native along-track resolution and calculate
2218		the interpolated all-photon STD quartiles to create an interpolated canopy
2219		level index. This will be used later for interpolating a ground surface.
2220		
2221	4 10	Refine Ground Estimates
<i></i>	1.10	nejme di vana Estimates
2222	1.	Detrend the interpolated ground surface using Asmooth. Smooth the
2223		detrended interpolated ground surface 10 times. All further ground surface
2224		smoothing use the moving average method:
2225		For j= 1:10
2226		A 10 11 C1 C1 A 1 C 10 *O
2226		AgroundSmooth = median filter (interp_Aground, SmoothSize*3)
2227		AgroundSmooth = smooth filter (AgroundSmooth, SmoothSize)
2228		End
2229		
2230	2.	This output (AgroundSmooth) from the filtering/smoothing function is an
2231		intermediate ground solution and it will be used to estimate the final
2232		solution.
2233	3.	
2234		>400 m
2235		FINALGROUND = median filter (Asmooth, SmoothSize)
2236		FINAL GROUND = smooth filter (FINAL GROUND, Smooth Size)

2237		Else
2238		FINALGROUND = AgroundSmooth
2239		end
2240	4.	If there are <b>canopy indices</b> identified along the segment:
2241		If there is a canopy photon identified at a location along-track above the
2242		ground surface, then at that location along-track
2243		FINALGROUND = AgroundSmooth
2244		else if there is a location along-track where the interpolated ground STD has
2245		an interpolated canopy level>=3
2246		FINALGROUND = Interp_Aground*1/3 + AgroundSmooth*2/3
2247		else
2248		FINALGROUND = Interp_Aground*1/2 + Asmooth*1/2
2249		end
2250	5.	Smooth the resulting interpolated ground surface (FINALGROUND) once
2251		using a median filter with window size of 9 then a smoothing filter twice with
2252		window size of 9. Select ground photons that lie within the point spread
2253		function (PSF) of FINALGROUND.
2254	6.	PSF is determined by sigma_atlas_land (Eq. 1.2) calculated at the photon
2255		resolution and thresholded between 0.5 to 1 m.
2256		a. Estimate the terrain slope by taking the gradient of FINALGROUND.
2257		Gradient is reported at the center of ((finalground(n+1)-
2258		$final ground (n-1))/(dist_x(n+1)-dist_x(n-1))/2$
2259		b. Linearly interpolate the sigma_h values to the photon resolution.
2260		c. Calculate sigma_topo (Eq. 1.3) at the photon resolution.
2261		d. Calculate sigma_atlas_land at the photon resolution using the sigma_h
2262		and sigma_topo values at the photon resolution.
2263		e. Set PSF equal to sigma_atlas_land.
2264		i. Any PSF $< 0.5$ m is set to 0.5 m as the minimum PSF.

2265	ii. Any PSF > 1 m is set to 1 m as the maximum PSF. Set psf_flag to
2266	true.
2267	
2268	4.11 Canopy Photon Filtering
2269	1. The first canopy filter will remove photons classified as top of canopy that
2270	are significantly above a smoothed median top of canopy surface. To
2271	calculate the smoothed median top of canopy surface:
2272	a. Linearly interpolate the median and standard deviation canopy
2273	window statistics, calculated from 4.9 (3), to the top of canopy photon
2274	resolution. Output variables: interpMedianC, interpStdC.
2275	b. Calculate a canopy window size using Eq. 3.4, where <i>length</i> = number
2276	of top of canopy photons. Output variable: winC.
2277	c. Create the median filtered and smoothed top of canopy surface,
2278	smoothedC, using a locally weighted linear regression smoothing
2279	method, "lowess" (Cleveland, 1979):
2280	<pre>smoothedC = median filter ( interpMedianC, winC )</pre>
2281	
2282	if SNR > 1, canopySmoothSpan = winC*2;
2283	else, canopySmoothSpan = smoothSpan;
2284	
2285	<pre>smoothedC = smooth filter ( smoothedC, canopySmoothSpan )</pre>
2286	d. Add the detrended heights back into the smoothedC surface:
2287	smoothedC = smoothedC + Asmooth
2288	2. Set canopy height thresholds based on the interpolated top of canopy STD:
2289	If SNR > 1, canopySTDthresh = 3; else, canopySTDthresh = 2;
2290	<pre>canopy_height_thresh = canopySTDthresh*interpStdC</pre>
2291	high_cStd = canopy_height_thresh > 10

2292		low_cStd = canopy_height_thresh < 3
2293		canopy_height_thresh(high_cStd) =
2294		canopy_height_thresh(high_cStd)/2
2295		<pre>canopy_height_thresh(low_cStd) = 3</pre>
2296	3.	Relabel as noise any top of canopy photons that are higher than smoothedC +
2297		canopy_height_thresh.
2298	4.	Next, interpolate a top of canopy surface using the remaining top of canopy
2299		photons (here we are trying to create an upper bound on canopy points). The
2300		interpolation method used is pchip. This output is named interp_Acanopy.
2301	5.	Photons falling below interp_Acanopy and above FINALGROUND+PSF are
2302		labeled as canopy points.
2303	6.	For 500 signal photon segments, if number of all canopy photons (i.e., canopy
2304		and top of canopy) is:
2305		< 5% of the total (when SNR > 1), OR
2306		< 10% of the total (when SNR $<= 1$ ),
2307		relabel the canopy photons as noise.
2308	7.	Interpolate, using the pchip method, a new top of canopy surface from the
2309		filtered top of canopy photons. This output is again named interp_Acanopy.
2310	8.	Again, label photons that lie between interp_Acanopy and
2311		FINALGROUND+PSF as canopy photons.
2312	9.	Since the canopy points have been relabeled, we need to do a final
2313		refinement of the ground surface:
2314		If canopy is present at any location along-track
2315		FINALGROUND = AgroundSmooth (at that location)
2316		Else if canopy is not present at a location along-track
2317		FINALGROUND = interp_Aground

2318	Smooth the resulting interpolated ground surface (FINALGROUND) once
2319	using a median filter with window size of SmoothSize (SmoothSize = 9), then
2320	a moving average smoothing filter twice with window size of SmoothSize
2321	(SmoothSize = 9)
2322	10. Relabel ground photons based on this new (and last) FINALGROUND solution
2323	+/- a recalculated PSF (via steps in 4.10 (6)). Points falling below the buffer
2324	are labeled as noise.
2325	11. Using Interp_Acanopy and this last FINALGROUND solution + PSF buffer,
2326	label all photons that lie between the two as canopy photons.
2327	12. Repeat the canopy cover filtering: For 500 signal photon segments, if
2328	number of all canopy photons (i.e., canopy and top of canopy) is:
2329	< 5% of the total (when SNR > 1), OR
2330	< 10% of the total (when SNR $<= 1$ ),
2331	relabel the canopy photons as noise. This is the last canopy labeling step.
2332	
2333	4.12 Compute individual Canopy Heights
2334	1. At this point, each photon will have its final label assigned in
2335	<b>classed_pc_flag</b> : 0 = noise, 1 = ground, 2 = canopy, 3 = top of canopy.
2336	2. For each individual photon labeled as canopy or top of canopy, subtract the Z
2337	height value from the interpolated terrain surface, FINALGROUND, at that
2338	particular position in the along-track direction.
2339	3. The relative height for each individual canopy or top of canopy photon will
2340	be used to calculate canopy products described in Section 4.15. Additional
2341	canopy products will be calculated using the absolute heights, as described in
2342	Section 4.15.1.
2343	

#### 2344 4.13 Final photon classification QA check 2345 1. Find any ground, canopy, or top of canopy photons that have elevations 2346 further than the ref\_dem\_limit from the reference DEM elevation value. 2347 Convert these to the noise classification. 2348 2. Find any relative heights of canopy or top of canopy photons that are greater 2349 than 150 m above the interpolated ground surface, FINALGROUND. Convert 2350 these to the noise classification. 2351 3. Find any FINALGROUND elevations that are further than the ref\_dem\_limit from the reference DEM elevation value. Convert those FINALGROUND 2352 2353 elevations to an invalid value, and convert any classified photons at the same 2354 indices to noise. 2355 4. If more than 50% of photons are removed in a segment, set ph removal flag 2356 to true. 2357 2358 4.14 Compute segment parameters for the Land Products 2359 1. For each 100 m segment, determine the classed photons (photons classified 2360 as ground, canopy, or top of canopy). 2361 a. If there are fewer than 50 classed photons in a 100 m segment, do not 2362 calculate land or canopy products. 2363 b. If there are 50 or more classed photons in a 100 m segment, extract 2364 the ground photons to create the land products. 2. If the number of ground photons > 5% of the total number of classed photons 2365 2366 within the segment (this control value of 5% can be modified once on orbit): 2367 a. Compute statistics on the ground photons: mean, median, min, max, 2368 standard deviation, mode, and skew. These heights will be reported 2369 on the product as h\_te\_mean, h\_te\_median, h\_te\_min, h\_te\_max, 2370 **h te mode**, and **h te skew** respectively described in Table 2.1. 2371 b. Compute the standard deviation of the ground photons about the 2372 interpolated terrain surface, FINALGROUND. This value is reported as 2373 **h\_te\_std** in Table 2.1.

2374 c. Compute the residuals of the ground photon Z heights about the 2375 interpolated terrain surface, FINALGROUND. The product is the root 2376 sum of squares of the ground photon residuals combined with the 2377 **sigma\_atlas\_land** term in Table 2.5 as described in Equation 1.4. This 2378 parameter reported as **h\_te\_uncertainty** in Table 2.1. 2379 d. Compute a linear fit on the ground photons and report the slope. This 2380 parameter is **terrain\_slope** in Table 2.1. 2381 e. Calculate a best fit terrain elevation at the mid-point location of the 2382 100 m segment: 2383 i. Calculate each terrain photon's distance along-track into the 2384 100 m segment using the corresponding ATL03 20 m products 2385 segment\_length and dist\_ph\_along, and determine the mid-2386 segment distance (expected to be  $50 \text{ m} \pm 0.5 \text{ m}$ ). 2387 1. Use the mid-segment distance to linearly interpolate a 2388 mid-segment time (**delta time** in Table 2.4). Use the 2389 mid-segment time to linearly interpolate other mid-2390 segment parameters: interpolated terrain surface, 2391 FINALGROUND, as **h\_te\_interp** (Table 2.1); **latitude** 2392 and longitude (Table 2.4). 2393 ii. Calculate a linear fit, as well as 3<sup>rd</sup> and 4<sup>th</sup> order polynomial fits 2394 to the terrain photons in the segment. 2395 iii. Create a slope-adjusted and weighted mid-segment variable, 2396 weightedZ, from the linear fit: Use terrain\_slope to apply a 2397 slope correction to each terrain photon by subtracting the 2398 terrain photon heights from the linear fit. Determine the mid-2399 segment location of the linear fit, and add that height to the 2400 slope corrected terrain photons. Apply a linear weighting to 2401 each photon based on its distance to the mid-segment location: 2402 1 / sqrt( (photon distance along – mid-segment distance)^2). 2403 Calculate the weighted mid-segment terrain height, weightedZ:

2404			sum( each adjusted terrain height * its weight ) / sum(all
2405			weights).
2406		iv.	Determine which of the three fits is best by calculating the
2407			mean and standard deviation of the fit errors. If one of the fits
2408			has both the smallest mean and standard deviations, use that
2409			fit. Else, use the fit with the smallest standard deviation. If
2410			more than one fit has the same smallest mean and/or standard
2411			deviation, use the fit with the higher polynomial.
2412		v.	Use the best fit to define the mid-segment elevation. This
2413			parameter is <b>h_te_best_fit</b> in Table 2.1.
2414			1. If h_te_best_fit is farther than 3 m from h_te_interp (best
2415			fit diff threshold), check if: there are terrain photons on
2416			both sides of the mid-segment location; or the elevation
2417			difference between weightedZ and h_te_interp is
2418			greater than the best fit diff threshold; or the number of
2419			ground photons in the segment is $\leq 5\%$ of total
2420			number of classified photons per segment. If any of
2421			those cases are present, use h_te_interp as the corrected
2422			h_te_best_fit. Otherwise use weightedZ as the corrected
2423			h_te_best_fit.
2424		f. Comp	ute the difference of the median ground height from the
2425		refere	nce DTM height. This parameter is <b>h_dif_ref</b> in Table 2.4.
2426			
2427	3.	If the number	r of ground photons in the segment <= 5% of total number of
2428		classified pho	otons per segment,
2429		a. Repor	t an invalid value for terrain products: <b>h_te_mean,</b>
2430		h_te_r	nedian, h_te_min, h_te_max, h_te_mode, h_te_skew, h_te_std,
2431		and h	_te_uncertainty respectively as described in Table 2.1.
2432		b. If the	number of ground photons in the segment is <= 5% of total
2433		numb	er of classified photons in the segment, compute <b>terrain_slope</b>

2434	via a linear fit of the interpolated ground surface, FINALGROUND,
2435	instead of the ground photons.
2436	c. Report the mid-segment interpolated terrain surface, FinalGround, as
2437	$h_{te_interp}$ as described in Table 2.1, and report $h_{te_best_fit}$ as the
2438	h_te_interp value.
2439	
2440	4.15 Compute segment parameters for the Canopy Products
2441	1. For each 100 m segment, determine the classed photons (photons classified as
2442	ground, canopy, or top of canopy).
2443	a) If there are fewer than 50 classed photons in a 100 m segment, do not
2444	calculate land or canopy products.
2445	b) If there are 50 or more classed photons in a 100 m segment, extract all
2446	canopy photons (i.e., canopy and top of canopy; henceforth referred to
2447	as "canopy" unless otherwise noted) to create the canopy products.
2448	2. Only compute canopy height products if the number of canopy photons is >
2449	5% of the total number of classed photons within the segment (this control
2450	value of 5% can be modified once on orbit).
2451	a) If the number of ground photons is also > 5% of the total number of
2452	classed photons within the segment, set <b>canopy_rh_conf</b> to 2.
2453	b) If the number of ground photons is $< 5\%$ of the total number of classed
2454	photons within the segment, continue with the relative canopy height
2455	calculations, but set canopy_rh_conf to 1.
2456	c) If the number of canopy photons is < 5% of the total number of classed
2457	photons within the segment, regardless of ground percentage, set
2458	canopy_rh_conf to 0 and report an invalid value for each canopy height
2459	variable.
2460	3. Again, the relative heights (height above the interpolated ground surface,
2461	FINALGROUND) have been computed already. All parameters derived in the
2462	section are based on relative heights.

- 4. Sort the heights and compute a cumulative distribution of the heights. Select the height associated with the 98% maximum height. This value is **h\_canopy** listed in Table 2.2.
  - Compute statistics on the relative canopy heights. Min, Mean, Median, Max and standard deviation. These values are reported on the product as h\_min\_canopy, h\_mean\_canopy, h\_max\_canopy, and canopy\_openness respectively in Table 2.2.
    - 6. Using the cumulative distribution of relative canopy heights, select the heights associated with the **canopy\_h\_metrics** percentile distributions (10, 15, 20, 25, 30, 35, 40, 45, 50, 55, 60, 65, 70, 75, 80, 85, 90, 95), and report as listed in Table 2.2.
    - 7. Compute the difference between h\_canopy and canopy\_h\_metrics(50). This parameter is **h\_dif\_canopy** reported in Table 2.2 and represents an amount of canopy depth.
    - 8. Compute the standard deviation of all photons that were labeled as Top of Canopy (flag 3) in the photon labeling portion. This value is reported on the data product as **toc\_roughness** listed in Table 2.2.
    - 9. The quadratic mean height, **h\_canopy\_quad** is computed by

$$2481 qmh = \sqrt{\sum_{i=1}^{Nca} \frac{h_i^2}{Nca}}$$

where  $N_{\text{ca}}$  is the number of canopy photons in the segment and  $h_i$  are the individual canopy heights.

### 4.15.1 Canopy Products calculated with absolute heights

1. The absolute canopy height products are calculated if the number of canopy photons is > 5% of the total number of classed photons within the segment. No number of ground photons threshold is applied for these. Absolute canopy heights are first determined as the relative heights of individual photons above the estimated terrain surface. Once those cumulative

2491 distribution is made, the absolute heights are the relative heights plus the 2492 best fit terrain height (h\_te\_bestfit). 2. The **centroid height** parameter in Table 2.2 is represented by all the classed 2493 2494 photons for the segment (canopy & ground). To determine the centroid 2495 height, compute a cumulative distribution of all absolute classified heights 2496 and select the median height. 2497 3. Calculate **h\_canopy\_abs**, the 98<sup>th</sup> percentile of the absolute canopy heights. 2498 4. Compute statistics on the absolute canopy heights: Min, Mean, Median, and 2499 Max. These values are reported on the product as **h min canopy abs.** 2500 **h\_mean\_canopy\_abs**, and **h\_max\_canopy\_abs**, respectively, as described in Table 2.2. 2501 2502 5. Again, using the cumulative distribution of relative canopy heights, select the 2503 heights associated with the **canopy h** metrics abs percentile distributions 2504 (10, 15, 20, 25, 30, 35, 40, 45, 50, 55, 60, 65, 70, 75, 80, 85, 90, 95) and then 2505 added to the h\_te\_bestfit, and report as listed in Table 2.2. 2506 4.16 Record final product without buffer 2507 1. Now that all products have be determined via processing of the *L-km* 2508 segment with the buffer included, remove the products that lie within the 2509 buffer zone on each end of the *L-km* segment. 2510 2. Record the final *L-km* products and move on to process the next *L-km* 2511 segment. 2512 2513

#### 5 DATA PRODUCT VALIDATION STRATEGY

Although there are no Level-1 requirements related to the accuracy and precision of the ATL08 data products, we are presenting a methodology for validating terrain height, canopy height, and canopy cover once ATL08 data products are created. Parameters for the terrain and canopy will be provided at a fixed size of 100 m along the ground track referred to as a segment. Validation of the data parameters should occur at the 100 m segment scale and residuals of uncertainties are quantified (i.e. averaged) at the 5-km scale. This 5-km length scale will allow for quantification of errors and uncertainties at a local scale which should reflect uncertainties as a function of surface type and topography.

#### 5.1 Validation Data

Swath mapping airborne lidar is the preferred source of validation data for the ICESat-2 mission due to the fact that it is widely available and the errors associated with most small-footprint, discrete return data sets are well understood and quantified. Profiling airborne lidar systems (such as MABEL) are more challenging to use for validation due to the low probability of exact overlap of flightlines between two profiling systems (e.g. ICESat-2 and MABEL). In order for the ICESat-2 validation exercise to be statistically relevant, the airborne data should meet the requirements listed in Table 5.1. Validation data sets should preferably have a minimum average point density of 5 pts/m². In some instances, however, validation data sets with a lower point density that still meet the requirements in Table 5.1 may be utilized for validation to provide sufficient spatial coverage.

Table 5.1. Airborne lidar data vertical height (Z accuracy) requirements for validation data.

ICESat-2 ATL08 Parameter	Airborne lidar (rms)
Terrain height	<0.3 m over open ground (vertical)
	<0.5 m (horizontal)

Canopy height	<2 m temperate forest, < 3 m tropical forest
Canopy cover	n/a

Terrain and canopy heights will be validated by computing the residuals between the ATL08 terrain and canopy height value, respectively, for a given 100 m segment and the terrain height (or canopy height) of the validation data for that same representative distance. Canopy cover on the ATL08 data product shall be validated by computing the relative canopy cover (cc = canopy returns/total returns) for the same representative distance in the airborne lidar data.

It is recommended that the validation process include the use of ancillary data sets (i.e. Landsat-derived annual forest change maps) to ensure that the validation results are not errantly biased due to non-equivalent content between the data sets.

Using a synergistic approach, we present two options for acquiring the required validation airborne lidar data sets.

# Option 1:

We will identify and utilize freely available, open source airborne lidar data as the validation data. Potential repositories of this data include OpenTopo (a NSF repository or airborne lidar data), NEON (a NSF repository of ecological monitoring in the United States), and NASA GSFC (repository of G-LiHT data). In addition to small-footprint lidar data sets, NASA Mission data (i.e. ICESat and GEDI) can also be used in a validation effort for large scale calculations.

# Option 2:

Option 2 will include Option 1 as well as the acquisition of additional airborne lidar data that will benefit multiple NASA efforts.

GEDI: With the launch of the Global Ecosystems Dynamic Investigation (GEDI) mission in 2018, there are tremendous synergistic activities for data validation between both the ICESat-2 and GEDI missions. Since the GEDI mission, housed on the International Space Station, has a maximum latitude of 51.6 degrees, much of the Boreal zone will not be mapped by GEDI. The density of GEDI data will increase as latitude increases north to 51.6 degrees. Since the data density for GEDI would be at its highest near 51.6 degrees, we would propose to acquire airborne lidar data in a "GEDI overlap zone" that would ample opportunity to have sufficient coverage of benefit to both ICESat-2 and GEDI for calibration and validation.

We recommend the acquisition of new airborne lidar collections that will meet our requirements to best validate ICESat-2 as well as be beneficial for the GEDI mission. In particular, we would like to obtain data over the following two areas:

- 1) Boreal forest (as this forest type will NOT be mapped with GEDI)
- 2) GEDI high density zone (between 50 to 51.6 degrees N). Airborne lidar data in the GEDI/ICESat-2 overlap zone will ensure cross-calibration between these two critical datasets which will allow for the creation of a global, seamless terrain, canopy height, and canopy cover product for the ecosystem community.
- In both cases, we would fly data with the following scenario:
- Small-footprint, full-waveform, dual wavelength (green and NIR), high point density (>20 pts/m²) and, over low and high relief locations. In addition, the newly acquired lidar data must meet the error accuracies listed in Table 5.1.
- Potential candidate acquisition areas include: Southern Canadian Rocky Mountains (near Banff), Pacific Northwest mountains (Olympic National Park, Mt. Baker-Snoqualmie National Forest), and Sweden/Norway. It is recommended that the

airborne lidar acquisitions occur during the summer months to avoid snow cover in either 2016 or 2017 prior to launch of ICESat-2.

### 5.2 Internal QC Monitoring

In addition to the data product validation, internal monitoring of data parameters and variables is required to ensure that the final ATL08 data quality output is trustworthy. Table 5.2 lists a few of the computed parameters that should provide insight into the performance of the surface finding algorithm within the ATL08 processing chain.

Table 5.2. ATL08 parameter monitoring.

Group	Description	Source	Monitor	Validate in Field
h_te_median	Median terrain height for segment	computed		Yes against airborne lidar data. The airborne lidar data should have an absolute accuracy of <30 cm rms.
n_te_photons n_ca_photons n_toc_photons	Number of classed (sum of terrain, canopy, and top of canopy) photons in a 100 m segment	computed	Yes. Build an internal counter for the number of segments in a row where there aren't enough photons (currently a minimum of 50 photons	

			per 100 m	
			segment is	
			used)	
h_te_interp	Interpolated terrain surface height,	computed	Difference	
	FINALGROUND		h_te_interp	
			and	
			h_te_median	
			and	
			determine if	
			the value is	
			> a specified	
			threshold. 2	
			m is	
			suggested	
			as the	
			threshold	
			value. This	
			is an	
			internal	
			check to	
			evaluate	
			whether the	
			median	
			elevation	
			for a	
			segment is	
			roughly the	
			same as the	
			interpolated	
			surface	
			height.	
h_dif_ref	Difference between h_te_median and	computed	This value	
	ref_dem		will be	
			computed	
			and flagged	
			if the	
			difference is	
			> 25 m. The	
			reference	
			DEM is the	
			onboard	
			DEM.	
h_canopy	95% height of individual canopy	computed	Yes, > a	Yes
	heights for segment		specified	against
			threshold	airborne
			(e.g. 60 m)	lidar data.
				The

				canopy heights derived from airborne lidar data should have a relative accuracy <2 m in temperate forest, <3 m in tropical
h_dif_canopy	Difference between h_canopy and canopy_h_metrics(50)	computed	Yes, this is an internal check to make sure the calculations on canopy height are not suspect	forest
psf_flag	Flag is set if computed PSF exceeds 1m	computed	Yes, this is an internal check to make sure the calculations are not suspect	
ph_removal_flag	Flag is set if more than 50% of classified photons in a segment is removed during final QA check	computed		
dem_removal_flag	Flag is set if more than 20% of classified photons in a segment is removed due to a large distance from the reference DEM	computed	Yes, this will check if bad results are due to bad DEM values or because too much noise was labeled as signal	

In addition to the monitoring parameters listed in Table 5.2, a plot such as what is shown in Figure 5.1 would be helpful for internal monitoring and quality assessment of the ATL08 data product. Figure 5.1 illustrates in graphical form what the input point cloud look like in the along-track direction, the classifications of each photon, and the estimated ground surface (FINALGROUND).

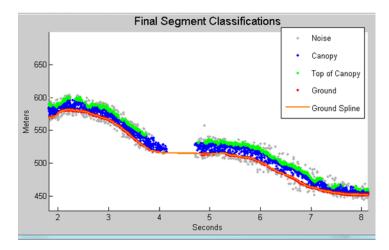


Figure 5.1. Example of *L-km* segment classifications and interpolated ground surface.

The following parameters are to be calculated and placed in the QA/QC group on the HDF5 data file, based on Table 5.2 of the ATL08 ATBD. Statistics shall be computed on a per-granule basis and reported on the data product. If any parameter meets the QA trigger conditional, an alert will be sent to the ATL08 ATBD team for product review.

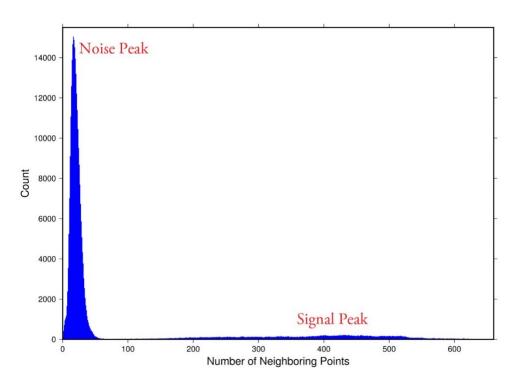
Table 5.3. QA/QC trending and triggers.

QA/QC trending description	QA trigger conditional
Percentage of segments with > 50 classed photons	None
Max, median, and mean of the number of contiguous	None
segments with < 50 classed photons	
Number and percentage of segments with difference in	> 50 segments in a row
h_te_interp - h_te_median is greater than a specified	
threshold (2 m TBD)	
Max, median, and mean of h_diff_ref over all segments	None
Percentage of segments where h_diff_ref > 25 m	Percentage > 75%
Percentage of segments where the h_canopy is > 60m	None
Max, median, and mean of h_diff	None
Percentage of segments where psf_flag is set	Percentage > 75%
Percentage of classified photons removed in a segment	Percentage > 50%
during final photon QA check	(i.e., ph_removal_flag is
	set to true)
Percentage of classified photons removed in a segment	Percentage > 20%
during the reference DEM threshold removal process	(i.e., dem_removal_flag is
	set to true)

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2674 Introduction	ructs ing a
	ructs ing a
This document provides a verbal description of how the DRAGANN (Differenti Regressive, and Gaussian Adaptive Nearest Neighbor) filtering system deconst a histogram into Gaussian components, which can also be called <i>iteratively fitt sum of Gaussian Curves</i> . The purpose is to provide enough detail for ASAS to cr operational ICESat-2 code required for the production of the ATL08, Land and Vegetation product. This document covers the following Matlab functions with DRAGANN:	
<ul> <li>mainGaussian_dragann</li> <li>findpeaks_dragann</li> <li>peakWidth_dragann</li> <li>checkFit_dragann</li> </ul>	
Components of the k-d tree nearest-neighbor search processing and histogram creation were covered in the document, <i>DRAGANN k-d Tree Investigations</i> , and been determined to function consistently with UTexas DRAGANN Matlab software.	have
2690	
2691 Histogram Creation	
Steps to produce a histogram of nearest-neighbor counts from a normalized place cloud segment have been completed and confirmed. Figure A.1 provides an exposed of such a histogram. The development, below, is specific to the two-dimension case and is provided as a review.	ample
The histogram represents the frequency (count) of the number of nearby phot within a specified radius, as ascertained for each point within the photon cloud radius, <i>R</i> , is established by first normalizing the photon cloud in time (x-axis) a height (y-axis), i.e., both sets of coordinates (time & height) run from 0 to 1; th average radius for finding 20 points is determined based on forming the ratio to the total number of the photons in the cloud ( <i>Ntotal</i> ): 20/ <i>Ntotal</i> .	d. The and in en an



**Figure A.1**. Histogram for Mabel data, channel 43 from SE-AK flight on July 30, 2014 at 20:16.

Given that the total area of the normalized photon cloud is, by definition, 1, then this ratio gives the average area, A, in which to find 20 points. A corresponding radius is found by the square root of  $A/\pi$ . A single equation describing the radius, as a function of the total number of photons in the cloud (remembering that this is done in the cloud normalized, two-dimensional space), is given by

$$R = \sqrt{\frac{20/N_{total}}{\pi}} \tag{A.1}$$

For the example in Figure A.1, *R* was found to be 0.00447122. The number of photons falling into this radius, at each point in the photon cloud, is given along the x-axis; a count of their number (or frequency) is given along the y-axis.

#### **Gaussian Peak Removal**

At this point, the function, mainGaussian\_dragann, is called, which passes the histogram and the number of peaks to detect (typically set to 10).

This function essentially estimates (i.e., fits) a sequence of Gaussian curves, from larger to smaller. It determines a Gaussian fit for the highest histogram peak, then removes it before determining the fit for the next highest peak, etc. In concept, the process is an iterative sequential-removal of the ten largest Gaussian components

within the histogram.

2725 In the process of *sequential least-squares*, parameters are re-estimated when input 2726 data is incrementally increased and/or improved. The present problem operates in 2727 a slightly reverse way: the data set is fixed (i.e., the histogram), but components 2728 within the histogram (independent Gaussian curve fits) are removed sequentially 2729 from the histogram. The paper by *Goshtasby & O'Neill* (1994) outlines the concepts. 2730 Recall that a Gaussian curve is typically written as  $y = a \cdot exp(-(x-b)^2/2c^2)$ 2731 (A.2)where a = the height of the peak; b = position of the peak; and c = width of the bell 2732 2733 curve. The function, mainGaussian\_dragann, computes the [a, b, c] values for the ten 2734 2735 highest peaks found in the histogram. At initialization, these [a, b, c] values are set to 2736 zero. The process begins by locating histogram peaks via the function, 2737 findpeaks dragann. 2738 2739 **Peak Finding** 2740 As input arguments, the findpeaks dragann function receives the histogram and a 2741 minimum peak size for consideration (typically set to zero, which means all peaks 2742 will be found). An array of index numbers (i.e., the "number of neighboring points", 2743 values along x-axis of Figure A.1) for all peaks is returned and placed into the 2744 variable peaks. 2745 The methodology for locating each peak goes like this: The function first computes 2746 the derivatives of the histogram. In Matlab there is an intrinsic function, called diff, 2747 which creates an array of the derivatives. Diff essentially computes the differences 2748 along sequential, neighboring values. "Y = diff(X) calculates differences between adjacent elements of X." [from Matlab Reference Guide] Once the derivatives are 2749 2750 computed, then findpeaks dragann enters a loop that looks for changes in the sign 2751 of the derivative (positive to negative). It skips any derivatives that equal zero. 2752 For the kth derivative, the "next" derivative is set to k+1. A test is made whereby if 2753 the k+1 derivative equals zero and k+1 is less than the total number of histogram 2754 values, then increment "next" to k+2 (i.e., find the next negative derivative). The test 2755 is iterated until the start of the "down side" of the peak is found (i.e., these iterations 2756 handle cases when the peak has a flat top to it). 2757 When a sign change (positive to negative) is found, the function then computes an

approximate index location (variable *maximum*) of the peak via

 $maximum = round\left(\frac{next-k}{2}\right) + k$ 

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(A.3)

2760 2761	These values of <i>maximum</i> are retained in the peaks array (which can be <i>grown</i> in Matlab) and returned to the function mainGaussian_dragann.
2762 2763 2764 2765	Next, back within mainGaussian_dragann, there are two tests to determine whether the first or last elements of the histogram are peaks. This is done since the findpeaks_dragann function will not detect peaks at the first or last elements, based solely on derivatives. The tests are:
2766 2767 2768 2769	If (histogram(1) > histogram(2) && max(histogram)/histogram(1) < 20) then insert a value of 1 to the very first element of the peaks array (again, Matlab can easily "grow" arrays). Here, max(histogram) is the highest peak value across the whole histogram.
2770	For the case of the last histogram value (say there are N-bins), we have
2771 2772	If ( histogram(N) > histogram(N-1) && max(histogram)/histogram(N) < 4 ) then insert a value of N to the very last element of the peaks array.
2773 2774 2775	One more test is made to determine whether there any peaks were actually found for the whole histogram. If none were found, then the function, mainGaussian_dragann, merely exits.
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2777	Identifying and Processing upon the Ten Highest Peaks
2778 2779 2780 2781	The function, mainGaussian_dragann, now begins a loop to analyze the ten highest peaks. It begins the $n^{\rm th}$ loop (where $n$ goes from 1 to 10) by searching for the largest peak among all remaining peaks. The index number, as well as the magnitude of the peak, are retained in a variable, called maximum, with dimension 2.
2782 2783 2784 2785	In each pass in the loop, the $[a,b,c]$ values (see eq. 2) are retained as output of the function. The values of $a$ and $b$ are set equal to the index number and peak magnitude saved in maximum(1) and maximum(2), respectively. The $c$ -value is determined by calling the function, peakWidth_dragann.
2786	Determination of Gaussian Curve Width
2787 2788 2789 2790 2791 2792 2793	The function, peakWidth_dragann, receives the whole histogram and the index number (maximum(1)) of the peak for which the value $c$ is needed, as arguments. For a specific peak, the function essentially searches for the point on the histogram that is about $\frac{1}{2}$ the size of the peak and that is furthest away from the peak being investigated (left and right of the peak). If the two sides (left and right) are equidistant from the peak, then the side with the smallest value is chosen (> $\frac{1}{2}$ peak).
2794 2795	Upon entry, it first initializes $c$ to zero. Then it initializes the index values left, xL and right, xR as index-1 and index+1, respectively (these will be used in a loop,

described below). It next checks whether the  $n^{th}$  peak is the first or last value in the histogram and treats it as a special case.

2798 At initialization, first and last histogram values are treated as follows:

If first bin of histogram (peak = 1), set left = 1 and xL = 1.

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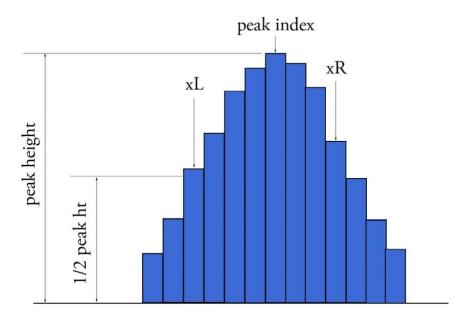
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If last bin of histogram, set right = m and xR = m, where m is the final index of the histogram.

2802 Next, a search is made to the left of the peak for a nearby value that is smaller than 2803 the peak value, but larger than half of the peak value. A while-loop does this, with 2804 the following conditions: (a) left > 0, (b) histogram value at left is  $\geq$  half of histo 2805 value at peak and (c) histo value at left is  $\leq$  histo value at peak. When these 2806 conditions are all true, then xL is set to left and left is decremented by 1, so that the 2807 test can be made again. When the conditions are no longer met (i.e., we've moved to 2808 a bin in the histogram where the value drops below half of the peak value), then the 2809 program breaks out of the while loop.

This is followed by a similar search made upon values to the right of the peak. When these two while-loops are complete, we then have the index numbers from the histogram representing bins that are above half the peak value. This is shown in Figure A.2.



**Figure A.2.** Schematic representation of a histogram showing xL and xR parameters determined by the function peakWidth\_dragann.

A test is made to determine which of these is furthest from the middle of the peak. In Figure A.2, xL is furthest away and the variable x is set to equal xL. The histogram

"height" at x, which we call  $V_x$ , is used (as well as x) in an inversion of Equation A.2

2820 to solve for *c*:

$$c = \sqrt{\frac{-(x-b)^2}{2\ln\left(\frac{V_X}{a}\right)}}$$
 (A.4)

- The function, peakWidth\_dragann, now returns the value of *c* and control returns to
- the function, mainGaussian\_dragann.
- The mainGaussian\_dragann function then picks-up with a test on whether the
- returned value of *c* is zero. If so, then use a value of 4, which is based on an *a priori*
- understanding that *c* usually falls between 4 and 6. If the value of *c* is not zero, then
- 2827 add 0.5 to *c*.
- At this point, we have the [a,b,c] values of the Gaussian for the  $n^{th}$  peak. Based on
- these values, the Gaussian curve is computed (via Equation A.2) and it is removed
- 2830 (subtracted) from the current histogram (and put into a new variable called
- 2831 newWave).
- 2832 After a Gaussian curve is removed from the current histogram, the following peak
- width calculations could potentially have a  $V_x$  value less than 1 from a. This would
- cause the width, c, to be calculated as unrealistically large. Therefore, a check is put
- in place to determine if  $a V_x < 1$ . If so,  $V_x$  is set to a value of a 1.
- 2836 Numeric Optimization Steps
- The first of the optimization steps utilizes a Full Width Half Max (*FWHM*) approach,
- 2838 computed via

$$FWHM = 2c\sqrt{2ln2} \tag{A.5}$$

- A left range,  $L_r$ , is computed by  $L_r$ =round(b-FWHM/2). This tested to make sure it
- doesn't go off the left edge of the histogram. If so, then it is set to 1.
- Similarly, a right range,  $R_r$ , is computed by  $R_r$ =round(b+FWHM/2). This is also tested
- to be sure that it doesn't go off the right edge of the histogram. If so, then it is set to
- the index value for the right-most edge of the histogram.
- Using these new range values, create a temporary segment (between  $L_r$  and  $R_r$ ) of
- the newWave histogram, this is called errorWave. Also, set three delta parameters
- for further optimization:
- 2848 DeltaC = 0.05; DeltaB = 0.02; DeltaA = 1
- The temporary segment, errorWave is passed to the function checkFit\_dragann,
- along with a set of zero values having the same number of elements as errorWave,
- the result, at this point, is saved into a variable called oldError. The function,
- checkFit\_dragann, computes the sum of the squares of the difference between two

2853 histogram segments (in this case, errorWave and zeros with the same number of 2854 elements as errorWave). Hence, the result, oldError, is the sum of the squares of the 2855 values of errorWaye. This function is applied in optimization loops, to refine the 2856 values of b and c, described below. 2857 Optimization of the b-parameter. The do-loop operates at a maximum of 1000 times. 2858 It's purpose is to refine the value of b, in 0.02 increments. It increments the value of 2859 b by DeltaB, to the right, and computes a new Gaussian curve based on  $b+\Delta b$ , which 2860 is then removed from the histogram with the result going into the variable 2861 newWave. As before, checkFit\_dragann is called by passing the range-limited part of 2862 newWave (errorWave) and returning a new estimate of the error (newError) which is then checked against oldError to determine which is smaller. If newError is ≥ 2863 2864 oldError, then the value of b that produced oldError is retained, and the testing loop 2865 is exited. *Optimization of the c-parameter.* Now the value of *c* is optimized, first to the left. 2866 2867 then to the right. It is performed independently of, but similarly, to the *b*-parameter, using do-loops with a maximum of 1000 passes. These loops increment (to right) or 2868 2869 decrement (to left) by a value of 0.05 (DeltaC) and use checkFit dragann to, again, check the quality of the fit. The loops (right and left) kick-out when the fit is found to 2870 2871 be smallest. 2872 The final, optimized Gaussian curve is now removed (subtracted) from the 2873 histogram. After removal, a statement "corrects" any histogram values that may 2874 drop below zero, by setting them to zero. This could happen due to any mis-fit of the 2875 Gaussian. 2876 The  $n^{\text{th}}$  loop is concluded by examining the peaks remaining in the histogram without the peak just processed by sending the  $n^{\text{th}}$ -residual histogram back into the 2877 2878 function findpeaks\_dragann. If the return of peak index numbers from 2879 findpeaks dragann reveals more than 1 peak remaining, then the index numbers for 2880 peaks that meet these three criteria are retained in an array variable called these: 2881 1. The peak must be located above b(n)-2\*c(n), and 2882 2. The peak must be located below b(n)+2\*c(n), and 2883 3. The height of the peak must be < a(n)/5. 2884 2885 The peaks meeting all three of these criteria are to be eliminated from further 2886 consideration. What this accomplishes is eliminate the nearby peaks that have a size 2887 lower than the peak just previously analyzed; thus, after their elimination, only 2888 leaving peaks that are further away from the peak just processed and are

presumably "real" peaks. The  $n^{th}$  iteration ends here, and processing begins with the

revised histogram (after having removed the peak just analyzed).

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### 2892 **Gaussian Rejection**

- The function mainGaussian\_dragann returns the [a,b,c] parameters for the ten
- 2894 highest peaks from the original histogram. The remaining code in dragann examines
- each of the ten Gaussian peaks and eliminates the ones that fail to meet a variety of
- 2896 conditions. This section details how this is accomplished.
- First, an approximate area, area 1=a\*c, is computed for each found peak and b, for all
- ten peaks, being the index of the peaks, are converted to an actual value via
- 2899 b+min(numptsinrad)-1 (call this allb).
- Next, a rejection is made for all peaks that have any component of [a,b,c] that are
- imaginary (Matlab isreal function is used to confirm that all three components are
- real, in which case it passes).
- 2903 To check for a narrow noise peak at the beginning of the histogram in cases of low
- 2904 noise rates, such as during nighttime passes, a check is made to first determine if the
- highest Gaussian amplitude, a, within the first 5% of the histogram is  $\geq 1/10$  \* the
- 2906 maximum amplitude of all Gaussians. If so, that peak's Gaussian width, c, is checked
- 2907 to determine if it is <= 4 bins. If neither of those conditions are met in the first 5%,
- 2908 the conditions are rechecked for the first 10% of the histogram. This process is
- repeated up to 30% of the histogram, in 5% intervals. Once a narrow noise peak is
- 2910 found, the process breaks out of the incremental 5% histogram checks, and the
- 2911 noise peak values are returned as [a0, b0, c0].
- 2912 If a narrow noise peak was found, the remaining peak area values, area1 (a\*c), then
- 2913 pass through a descending sort; if no narrow noise peak was found, all peak areas go
- through the descending sort. So now, the [a,allb,c]-values are sorted from largest
- "area" to smallest, these are placed in arrays [a1, b1, c1]. If a narrow noise peak was
- 2916 found, it is then appended to the beginning of the [a1, b1, c1] arrays, such that a1 =
- 2917 [a0 a1], b1 = [b0 b1], c1 = [c0 c1].
- In the case that a narrow noise peak was not found, a test is made to check that at
- least one of the peaks is within the first 10% of the whole histogram. It is done
- inside a loop that works from peak 1 to the number of peaks left at this point. This
- loop first tests whether the first (sorted) peak is within the first 10% of the
- 2922 histogram; if so, then it simply kicks out of the loop. If not, then it places the loop's
- 2923 current peak into a holder (ihold) variable, increments the loop to the next peak and
- runs the same test on the second peak, etc. Here's a Matlab code snippet:

```
2925
             inds = 1:length(a1);
2926
2927
2928
2929
             for i = 1: length(b1)
                 if b1(i) <= min(numptsinrad) + 1/10*max(numptsinrad)</pre>
                      if i==1
                          break;
2930
                      end
2931
                      ihold = inds(i);
2932
                      for i = i:-1:2
2933
                          inds(j) = inds(j-1);
2934
                      inds(1) = ihold;
```

```
2936 break
2937 end
2938 end
```

2944

2945

2961

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The j-loop expression gives the init\_val:step\_val:final\_val. The semi-colon at the end of statements causes Matlab to execute the expression without printout to the user's screen. When this loop is complete, then the indexes (inds) are re-ordered and placed back into the [a1,b1,c1] and area1 arrays.

Next, are tests to reject any Gaussian peak that is entirely encompassed by another peak. A Matlab code snippet helps to describe the processing.

```
2946
             reject any gaussian if it is fully contained within another
2947
           isR = true(1,length(a1));
2948
           for i = 1: length(a1)
2949
               ai = a1(i);
2950
               bi = b1(i);
2951
               ci = c1(i);
2952
               aset = (1-(c1/ci).^2);
2953
               bset = ((c1/ci).^2*2*bi - 2*b1);
2954
               cset = -(2*c1.^2.*log(a1/ai)-b1.^2+(c1/ci).^2*bi^2);
2955
               realset = (bset.^2 - 4*aset.*cset >= 0) | (a1 > ai);
2956
               isR = isR & realset;
2957
2958
           a2 = a1(isR);
2959
           b2 = b1(isR);
2960
           c2 = c1(isR);
```

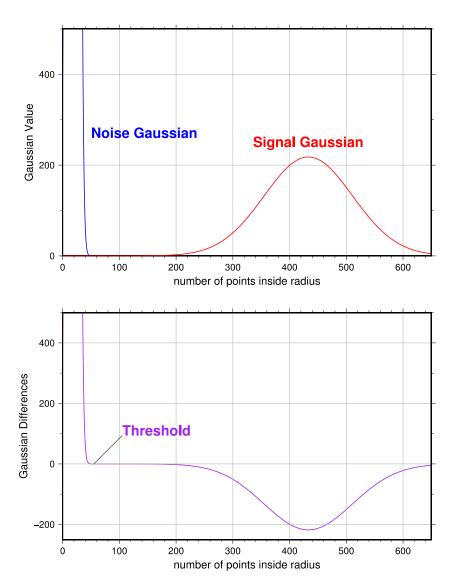
The logical array isR is initialized to all be true. The i-do-loop will run through all peaks. The computations are done in array form with the variables aset, bset, cset all being arrays of length(a1). At the bottom of the loop, isR remains "true" when either of the conditions in the expression for realset is met (the single "|" is a logical "or"). Also, the nomenclature, ".\*" and ".^", denote element-by-element array operations (not matrix operations). Upon exiting the i-loop, the array variables [a2,b2,c2] are set to the [a1,b1,c1] that remain as "true." [At this point, in our test case from channel 43 of East-AK Mable flight on 20140730 @ 20:16, six peaks are still retained: 18, 433, 252, 33, 44.4 and 54.]

Next, reject Gaussian peaks whose centers lay within  $3\sigma$  of another peak, unless only two peaks remain. The code snippet looks like this:

```
2973
2974
2975
2976
             isR = true(1, length(a2));
             for i = 1: length(a2)
                  ai = a2(i);
                  bi = b2(i);
<u> 2</u>977
                  ci = c2(i);
2978
                  realset = (b2 > bi+3*ci | b2 < bi-3*ci | b2 == bi);
2979
                  realset = realset | a2 > ai;
2980
                  isR = isR & realset;
<del>2</del>981
             end
2982
             if length(a2) == 2
2983
                 isR = true(1, 2);
2984
             end
```

2985 2986 2987	a3 = a2(isR); b3 = b2(isR); c3 = c2(isR);
2988	
2989 2990 2991 2992 2993	Once again, the isR array is initially set to "true." Now, the array, realset, is tested twice. In the first line, one of three conditions must be true. In the second line, if realset is true or a2 > ai, then it remains true. At this point, we've pared down, from ten Gaussian peaks, to two Gaussian peaks; one represents the noise part of the histogram; the other represents the signal part.
2994 2995 2996 2997 2998	If there are less than two peaks left, a thresholding/histogram error message is printed out. If the lastTryFlag is not set, DRAGANN ends its processing and an empty IDX value is returned. The lastTryFlag is set in the preprocessing function which calls DRAGANN, as multiple DRAGANN runs may be tried until sufficient signal is found.
2999 3000 3001	If there <u>are</u> two peaks left, then set the array [a,b,c] to those two peaks. [At this point, in our test case from channel 43 of East-AK Mable flight on 20140730 @ 20:16, the two peaks are: 18 and 433.]
3002	
3003	Gaussian Thresholding
3004 3005	With the two Gaussian peaks identified as noise and signal, all that is left is to compute the threshold value between the Gaussians.
3006 3007 3008 3009 3010 3011 3012 3013 3014	An array of xvals is established running from min(numptsinrad) to max(numptsinrad). In our example, xvals has indices between 0 and 653. For each of these xvals, Gaussian curves (allGauss) are computed for the two Gaussian peaks $[a,b,c]$ determined at the end of the previous section. This computation is performed via a function called gaussmaker which receives, as input, the xvals array and the $[a,b,c]$ parameters for the two Gaussian curves. An array of heights of the Gaussian curves is returned by the function, computed with Equation A.2. In Matlab, the allGauss array has dimension $2x654$ . An array, noiseGauss is set to be equal to the $1^{st}$ column of allGauss.
3015 3016 3017	An if-statement checks whether the b array has more than 1 element (i.e., consisting of two peaks), if so, then nextGauss is set to the $2^{nd}$ column of allGauss, and a difference, noiseGauss-nextGauss, is computed.
3018 3019 3020 3021	The following steps are restricted to be between the two main peaks. First, the first index of the absolute value of the difference that is near-zero (defined as 1e-8) is found, if it exists, and put into the variable diffNearZero. This is expected to be found if the two Gaussians are far away from each other in the histogram.
3022 3023	Second, the point (i.e., index) is found of the minimum of the absolute value of the difference; this index is put into variable, signchanges. This point is where the sign

changes from positive to negative as one moves left-to-right, up the Gaussian curve differences (noise minus next will be positive under the peak of the noise curve, and negative under the next (signal) curve). Figure A.3 (top) shows the two Gaussian curves. The bottom plot shows their differences.



**Figure A.3**. Top: two remaining Gaussian curves representing the noise (blue) and signal (red) portions of the histogram in Figure A.1. Bottom: difference noise – signal of the two Gaussian curves. The threshold is defined as the point where the sign of the differences change.

3033 3034 3035	If there is any value stored in diffNearZero, that value is now saved into the variable threshNN. Else, the value of the threshold in signchanges is saved into threshNN, concluding the if-statement for b having more than 1 element.
3036 3037	An else clause (b !> 1), merely sets threshNN to b+c, i.e., 1-standard deviation away from mean of the (presumably) noise peak.
3038 3039 3040 3041	The final step is mask the signal part of the histogram where all indices above the threshNN index are set to logical 1 (true). This is applied to the numptsinrad array, which represents the photon cloud. After application, dragann returns the cloud with points in the cloud identified as "signal" points.
3042 3043	The Matlab code has a few debug statements that follow, along with about 40 lines for plotting.
3044	
3045	References
3046 3047	Goshtasby, A & W. D. O'Neill, Curve Fitting by a Sum of Gaussians, <i>CVGIP: Graphical Models and Image Processing</i> , V. 56, No. 4, 281-288, 1994.